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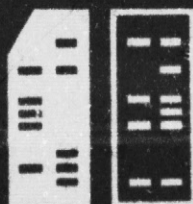
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MONTE CARLO SIMULATION OF NONLINEAR RADIATION  
INDUCED PLASMAS

by

Benjamin Shaw-hu Wang

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THESIS

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# MONTE CARLO SIMULATION OF NONLINEAR RADIATION INDUCED PLASMAS

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A Monte Carlo simulation model for radiation induced plasmas with nonlinear properties due to recombination has been developed, employing a piecewise linearized predict-correct iterative technique. Several important variance reduction techniques have been developed and incorporated into the model, including an antithetic variates technique (negatively correlated variates) which has proven to be most useful. This approach is especially efficient for plasma systems with inhomogeneous media, multidimensions, and irregular boundaries, where other analytic or numerical solutions are either unrealistic or impractical. This model is quite general in scope and should be applicable to a variety of similar particle transport and diffusion-type plasma problems.

The Monte Carlo code developed in this work has been applied to the determination of the electron energy distribution function and related parameters for a noble gas plasma created by  $\alpha$ -particle irradiation. The radiation induced plasma involved was characterized by the following features:

- 1) a continuous internal volume-source of high energy electrons

- 2) an applied electric field (optional)
- 3) secondary electron production via ionization processes
- 4) losses via recombination and leakage across boundaries

The calculations were specifically carried out for a helium gas medium with an electron source rate from  $10^{14}$  to  $10^{22}$  ( $\#/cm^3$ -sec), initial electron energies from 70 to 1000 eV, pressures from 10 to 760 torr and E/P (electric field/pressure) ratios from 1 to 10 V/cm-torr.

For the lower source rates, it is observed that, in the zero electric-field case, the low-energy portion of the distribution function is quite close to a standard Maxwellian distribution. However, the high-energy portion (above the ionization potential) of the distributions is a rapidly decaying parabolic-shaped tail which can be crudely represented by a  $1/(\text{Energy})$  distribution. The addition of the electric field causes the distribution function in the low-energy region to shift towards Smit's (or Druyvesteyn's) distribution while the high-energy portion essentially retains its zero field shape. As the source rate increases beyond certain value (roughly  $10^{18} \text{ cm}^{-3}\text{-sec}^{-1}$  in the present case), the low-energy region begins to depart from the standard Maxwellian or Smit distributions and this departure can become quite large.



A major emphasis in this development was placed on computation efficiency. The present computation scheme with its variance and accelerated convergence techniques is thought to be quite efficient compared to alternate approaches.

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## CHAPTER I

### INTRODUCTION

#### A. Statement of the Problem and Objectives

Two major goals were set for this study. The first was to develop an efficient Monte Carlo simulation technique for solving radiation-induced plasma problems. This required adapting several variance reduction techniques in order to increase the Monte Carlo efficiency. The second goal was to apply the model developed herein to study the characteristics of a noble-gas plasma created by alpha-particle irradiation.

The major point of interest is the determination of the electron energy distribution in the nuclear radiation induced plasma. This distribution is of direct importance to current experimental studies at the University of Illinois where the possibility of directly pumping\* gas lasers with nuclear radiation is under study (6). Recent results from these studies also show that nuclear radiation can be used to enhance the output from some type of electrically pumped lasers (5).

#### B. The Importance of Nuclear Radiation Produced Plasmas and the Physical Model

Nuclear radiation can be used in two ways: to directly pump the laser, or alternately, where an electric discharge

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\*"Pumping" of a laser is defined as the introduction of energy into the lasing medium required for population inversion. The electrical current which creates the discharge in the conventional gas laser is the pump in that case. Flash lamps are used for many solid state lasers.

serves as the main pump, to enhance the operation of lasers. The possibility of efficient coupling of lasers and nuclear reactors has generated an interest in the use of high-energy ion beams to pump lasers. To avoid radiation damage problems, gas lasers are generally thought of in this connection.

There are several important reasons for searching for more direct methods of coupling energy from reactor sources into lasers (8). One is the possibility that the resulting laser may offer unique characteristics, e.g., new frequencies and new methods of modulation and control. Simplified coupling could also offer weight reductions and improved reliability, which may be of advantage in space applications. Also, it may eventually become feasible to simultaneously extract several energy forms, e.g., laser and electrical power from a single nuclear station.

In most gas lasers the excited states, and consequently the inversion of population necessary for laser operation, are formed directly or indirectly by electronic collisions in a plasma generated by an electrical discharge. However, an alternate way to accomplish this goal is via nuclear radiation which can produce ionization and excitation through both direct interactions and interactions due to secondary-electron production. Given the high neutron fluxes available from nuclear reactors, the particles from neutron-induced nuclear reactions (e.g.,  $\alpha$ -particles from a Boron coating on the wall of a laser tube) can furnish the nuclear radiation

source. The description of the transport of both the primary radiation and the secondary electrons, from their point of birth throughout their path in the gas, is a key problem in the quantitative evaluation of the plasma. Therefore the development of nuclear-radiation pumping requires a thorough understanding of the radiation-induced plasma.

The following physical model is applied to the situation of a radiation-induced plasma. Pertinent quantities of interest include: electron energy distribution spectrum; average electron energy; and rate of energy loss.

The heavy-charged particles ( $\alpha$ -particles) are born in coating and enter an adjacent gas. In the present calculations, the cylindrical geometry of the actual laser tube is approximated by a slab geometry (cf. Figure 1) with two parallel plane boron coatings separated by a gaseous medium. The  $\alpha$ -particle radiation continuously creates high energy electrons throughout the gas volume (with energy  $E_0$ ). The resulting plasma is a weakly ionized, three-component plasma (electrons, positive ions, and neutral gas atoms). It is assumed that the alpha-particle trajectory can be treated analytically using methods such as developed in (7). Further, this reference shows how to obtain the primary secondary-electron ( $\delta$ -ray) energy spectrum. The difficult problem, however, is to follow the thermalization and further secondary production of these primary electrons.

Thus, the Monte Carlo simulation starts off by following these electrons on their life time histories, i.e., by simulating the course of their thermalization. Electrons go through various processes of collisions and scattering, resulting in a complex process of diffusion, energy degradation and secondary production, but finally, they recombine with positive ions or escape from the system, ending their histories.

With the radiation rates of interest here, the plasma is only weakly ionized and electrons mostly thermalize by collisions with the neutral background-gas atoms. In other words, collisions with ions and other free electrons can be neglected. In this sense, the problem is simplified in that it remains linear. However, some non-linearity enters in the following way. The recombination process depends on the ion density and spatial distribution. However, the ions are originally formed by ionizations which occur as the electrons slow down.

Another important characteristic is that the plasma is highly non-equilibrium. The thermalized electrons and ions form an approximately Maxwellian background, but superimposed on this is the high energy tail due to the electrons in the process of slowing. Also, the alpha particles themselves represent a non-equilibrium component.

In summary, the special features of the plasma system are:



- 1) Continuous internal volume source.
- 2) An applied electric field may or may not be present.
- 3) Finite cylindrical geometry--approximated here by slab geometry.
- 4) Electron-ion pair production and recombination processes.
- 5) The energy range of the electrons range from kilovolts to thermal (eV.).
- 6) The system represents a non-equilibrium plasma, but in a steady-state condition (or stationary).
- 7) Due to relatively small densities of electrons and ions compared with that of gas atoms, direct interactions between the electrons themselves or electron-ion collisions are neglected.

#### C. General Approaches and Methods of Solution

Possible approaches toward solutions of these types of problems (particle transport and diffusion) can be categorized as follows:

- 1) Analytical solution of the governing Boltzmann transport equation.
- 2) Numerical solution of the Boltzmann transport equation.
- 3) Direct analog simulation (Monte Carlo).
- 4) Mixtures of the above approaches.

The governing equation for this type of problem is the Boltzmann equation, and most classical methods have involved approximate analytic solutions of it. Solutions are possible only after various approximations and simplifications are made, and such solutions are often very restrictive and/or difficult to extend to actual situations of interest. More

or less the same drawbacks apply to numerical solutions. However, in this case approximations are necessary or the computation time becomes too long, especially as the dimensionality and the complexity of the problem increases. In fact, the computation time increases roughly exponentially with the dimensionality of the problem. Usually the governing partial differential and integral equations (coupled Boltzmann equations) are reduced to finite difference equations, and the requirements for stability and convergence severely limit the maximum step sizes of the independent variables. Further difficulties with the practical application of such approaches arise because it is often difficult to include realistic boundary conditions.

Direct simulation by the application of Monte Carlo techniques seems to be the better approach; sometimes the only approach for such problems. It by-passes the direct solution of the partial differential-integral equation, and generally no approximations and/or simplifications are necessary. Complicated boundary conditions can be handled easily, and the computation time increases at the most linearly with the dimensionality of the problem. However, computer time still may be a problem so that judicious selection of numerical schemes and variance reduction techniques are required.

Many methods have been devised in the past which involve a mixture of analytic, numerical, and Monte Carlo approaches.

In such situations, Monte Carlo is usually applied to the most difficult part of the problem. For instance, in the evaluation of the non-linear, five-dimensional collision integral in the Boltzmann equation, straight forward numerical quadrature would require months of computation time even on the present day fast computers. The Monte Carlo (statistical sampling techniques which closely resemble the actual collision phenomena) techniques can produce results within reasonable computation time (24). The Monte Carlo method may be defined as a numerical device (numerical experimentation) for studying an artificial stochastic model of a physical or mathematical process. This study may be approached by one of two methods. The first is the process of simulation (direct analog simulation) in which the particles - electrons and ions in the case of plasma systems, or photons in the case of radiative transfer - are represented by a game of chance in the computer. In the second method, which is non-simulation, the Monte Carlo technique is strictly used as a numerical stochastic device for the solution of a given integral or partial differential equations (24,25).

## CHAPTER II

### THE MATHEMATICAL MODEL: NUMERICAL SCHEME AND PHYSICAL PROBLEM

#### A. General Description of the Physics of the Problem

The plasma under study is a weakly-ionized three-component plasma, and is non-equilibrium as described in Chapter I. The  $\alpha$ -particle radiation continuously creates high-energy electrons throughout the gas volume. The Monte Carlo simulation is to trace these electrons' life histories and their complex processes of diffusion and energy degradation. Since the diffusion process of the electrons through the medium is affected by the density of the positive ions and the other electrons present (background electrons), the resulting problem is non-linear.

##### 1. Physical Processes Involved

The mechanisms by which particles interact with the atoms of the medium are elastic and inelastic collisions, by which the incident particle is slowed and dissipates some or all of its kinetic energy to the medium by the combined action of the above mentioned processes. The particle can also experience so called superelastic collision by which the particle collides with the excited atoms gaining some energy. Finally, the particle can collide and recombine with positive ions or can leak out of the system thus ending its life



history.

As described in Chapter I, the type of interaction which is involved in any particular collision is governed by the laws of chance. For large numbers of interactions, the frequency of occurrence of any given interaction is determined by the relative magnitudes of the interaction cross sections. Secondary knock-on electrons due to the ionization process may be capable of further interacting with the atoms and producing additional secondary electrons.

As the system reaches steady-state, a balance is reached such that the birth rate for new particles, i.e., the rate at which primary or source electrons enter the system plus the production rate of secondary electrons due to ionization should approximately equal the loss rate due to leakage and absorption or recombination.

Since the transport of radiation through a medium affects its absorption properties, the transport of individual particles is indirectly affected by the density of the other particles present, and this introduces some nonlinearities into the system. We shall adopt a piecewise predict-correct technique to approximate the nonlinear system with a linear model.

We are concerned with the simulation, by random sampling of the scattering of the charged particles. The direct simulation of the physical scattering processes in such a way that the diffusion process is imitated by letting

the particles carry out a random walk, each step of which takes into account the combined effect of many collisions. The mechanisms by which electrons interact with the atoms of the medium are elastic, inelastic, and superelastic collisions as well as recombination with positive ions. The incident source electron is slowed and dissipates some or all of its energy to the medium by the combined action of the various scattering processes, until it escapes, or recombines with an ion (radiative energy losses are neglected). The probability that an interaction occurs is determined by the total collision cross-section. The type of interaction is then determined by the relative magnitudes of the individual cross-sections.

## 2. Relation to the Transport Equation

Although no direct use will be made of the governing Boltzmann transport equation in the sequel, we shall write it down briefly, in order to indicate the mathematical problems to be solved implicitly by the Monte Carlo method. It is a nonlinear integro-differential equation of the form

$$\frac{\partial f}{\partial t} + \bar{u} \cdot \nabla_r f - \frac{e\hbar}{m} (\bar{E} + \bar{u} \times \bar{B}) \cdot \nabla_u f = \left( \frac{\partial f}{\partial t} \right)_{\text{collisions}} \quad (1)$$

where  $f$ : number density distribution function

$\bar{B}$ : magnetic field

$\bar{u}$ : velocity of the particles

$\bar{E}$ : electric field strength

t: time

r: spatial variable

In our situation  $\vec{B}=0$  since the magnetic field does not exist.

Further, if we assume an infinite medium and uniform source,

$\nabla_r f=0$ , the distribution function  $f$  depends only on energy.

Then the above equation reduces to

$$\frac{\partial f}{\partial t} - \frac{e}{m} \vec{E} \cdot \nabla_u f = \left( \frac{\partial f}{\partial t} \right)_{\text{collisions}} \quad (2)$$

The right hand side term is the collision term, which is the most difficult part of the equation. The effects of all relevant particle interactions must be included in this term. For the case of electron collisions with stationary gas atoms of uniform density, we can write the collision term as:

$$\left( \frac{\partial f}{\partial t} \right)_{\text{collision}} = \left( \frac{dN_{\text{in}} - dN_{\text{out}}}{dV \cdot dt} \right) \quad (3)$$

where  $dN_{\text{in}}$  gives the number of phase space points (electrons) that are scattered into the phase-space volume  $dV$  by collision in time  $dt$ . Similarly,  $dN_{\text{out}}$  represents the number of points scattered out of  $dV$  by collisions in time  $dt$ . This Eq. 2 states that the change in the distribution function  $f$  are caused by collisions of various kinds and the external force due to the electric field.

The assumptions that enter into the derivation of the transport equation which carry over to the Monte Carlo

simulation include:

a) the scattering centers (background atoms and ions) are distributed at random, i.e., possible correlations between the positions of various atoms and ions are not taken into account.

b) In the absence of an electric field, the trajectory of the particle is idealized as a zig-zag path, consisting of flights interrupted by sudden collisions in which the energy and direction of the particle changed. The quantum-mechanical interference (electron diffraction) resulting from the coherent scattering by several centers are neglected, and a particle, in the course of traversing the medium, interacts with one scattering center at a time.

The diffusion process in terms of the transport equation is analogous to the use of Eulerian coordinates in hydrodynamics, in that one asks about the flux at given points in space. By contrast, the Monte Carlo method uses Lagrangian coordinates; one attaches a label to a particular bit of fluid, i.e., a diffusing particle, and follows its history. Then, by sampling many histories, one is able, at least in principle, to solve any diffusion problem.

#### B. The Stochastic Model - Numerical Scheme

The essential approach to the nonlinear problem adopted here is the use of piecewise linearized steps. In each such step (an amount of time  $\Delta t$  along the time axis), the



dominate parameters, which characterize the nonlinear properties, are predicted and considered to be fixed with respect to the independent variables. During the time  $\Delta t$ , the system can be considered as more or less linear, so the conventional Monte Carlo simulation technique is used. The quantities pertaining to the given system are computed for each of a reasonably large number of independent particle histories. The desired parameter values are obtained by averaging the results over all of the histories. At the end of the step, the dominate parameters are corrected before beginning the next step. In each step (time interval  $\Delta t$ , or cycle), instead of computing the trajectory and angular deflection of an electron after each interaction, we perform a random walk calculation. Since each step of the random walk takes into account the combined effects of many individual interactions, we form a so-called condensed history by this approach. Let us describe the model in some detail, step by step.

#### 1. The Geometry and the Coordinate System

The initial spatial geometry in this problem is a one-dimensional, finite slab (from  $z=0$  to  $z=D$ ) which is subdivided into uniform intervals called zones as shown in Figure 1. The energy range is from 0 to  $E_0$  ( $\approx 1\text{keV}$ ) which is also subdivided into three regions:  $0 \sim 1$  (thermal energy region), 1 to  $ED2$ , and  $ED2$  to  $E_0$ . Each of these

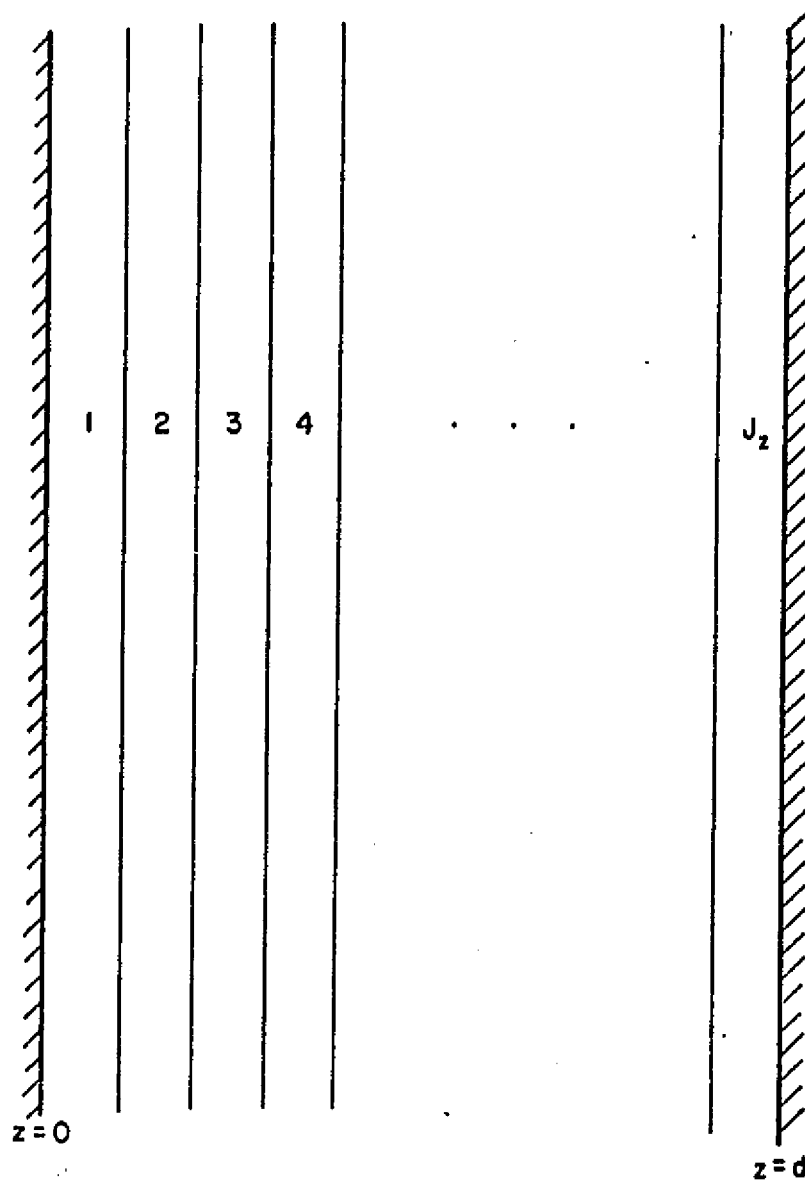


Figure 1. Slab geometry of the medium with subdivided zones.

regions is further divided into smaller uniform or logarithmic spacing (intervals). The particle number density distribution  $f(E,z)$  is a function of energy  $E$  and spatial position  $z$  and is generated such that particles with energy and spatial position falling into the same intervals are grouped together and normalized by the total particles processed. Although the system is stationary, we deliberately introduce the time variable for the convenience of recording output information and carrying out the calculation steps. The time serves as a clock, or we can imagine that the system begins in a transient state. Then, as time goes on it will eventually reach steady state. The storage of the timing information for individual particles is not necessary. Instead of following a single particle from the beginning to the end of its history, we follow a large number of particles, say  $N$  (which represent the sample particles of the whole system population) through a small time interval  $\Delta t$  (called the cycle time). Within this time interval, the system is considered to be linear, and the particles are processed more or less independently as in a conventional linear Monte Carlo model. At the end of  $\Delta t$ , called census time, we record data information for each particle (record particle state information). During this time interval (or present time cycle) some particles may experience collisions of some kind, some may not, and new particles (source particles, secondary particles produced

by ionization) enter the system. Some particles may be absorbed or escape from the system; their histories are therefore ended. This represents a random walk calculation, and each step of the random walk takes into account the combined effect of many individual interactions. The trajectory of a particle which is computed by this approach is called a condensed history. The residue particles (ionization produced secondary particles) are also set into motion, and their histories are followed separately. They are like all the other source particles in all respects. The complete description of particle histories is approximated by the condensed histories, or we take sequences of snapshots of all the particles in the system at various times to provide a moving picture of the history that can be used to estimate the quantities of interest. The path lengths travelled by the particles are also used as a clock to measure time, or equivalently, the variable  $t$  may be replaced by  $S(t) = \int_{t_0}^t v(x)dx$ . The trajectories of the particles (the condensed histories) are described by the arrays as indicated in Figure 2. The superscript denotes the cycle time number; the subscript denotes particle number where the index 0 refers to the initial state of the particle. Other notation includes  $E$ : the particle energy,  $z$ : spatial position, and  $U$ : the direction, and  $wt$ : the weight. When the particle has traveled a path-length  $DCEN$  (in the same duration  $\Delta t$ ), a condensed history

$i \backslash t$	$t_0$	$t_1$		$t_j$		$t_m$
1	$E_1^0$	$E_1^1$	. . .	$E_1^j$	. . .	$E_1^m$
	$Z_1^0$	$Z_1^1$		$Z_1^j$		$Z_1^m$
	$U_1^0$	$U_1^1$		$U_1^j$		$U_1^m$
	$wt_1^0$	$wt_1^1$		$wt_1^j$		$wt_1^m$
-----						
2	$E_2^0$	$E_2^1$	. . .	$E_2^j$	. . .	$E_2^m$
	$Z_2^0$	$Z_2^1$		$Z_2^j$		$Z_2^m$
	$U_2^0$	$U_2^1$		$U_2^j$		$U_2^m$
	$wt_2^0$	$wt_2^1$		$wt_2^j$		$wt_2^m$
-----						
.	.	.		.		.
.	.	.		.		.
.	.	.		.		.
-----						
N	$E_n^0$	$E_n^1$		$E_n^j$		$E_n^m$
	$Z_n^0$	$Z_n^1$		$Z_n^j$		$Z_n^m$
	$U_n^0$	$U_n^1$		$U_n^j$		$U_n^m$
	$wt_n^0$	$wt_n^1$		$wt_n^j$		$wt_n^m$

Figure 2. Snapshots of the particles' condensed histories.

is sampled. Letting the particle carry out a random walk in each step (from state  $M$  to state  $M+1$ ) takes into account the combined effect of many collisions. The sizes of the steps of the random walk, i.e., the pathlength intervals DCEN (equivalent to  $\Delta t$ ) are chosen such that the total number of steps should be kept as small as possible. Too large a step size ( $\Delta t$ ) may cause inaccuracies in the approximation. However, the length of the Monte Carlo calculation is directly proportional to the total number of steps. Thus, a compromise must be reached such that we choose as big a step size as possible without introducing intolerable inaccuracies. A reasonable initial value is chosen such that the particle with the average energy of the time cycle travels one mean free path, i.e.,  $\Delta t = D_{\text{avg}}/V_{\text{avg}}$ . Different time cycles have different  $E_{\text{avg}}$  ( $V_{\text{avg}}$ ) and  $D_{\text{avg}}$  so  $\Delta t$  may vary, but as steady-state is approached,  $\Delta t$  will not change much. By this technique we not only randomize and average the different particle samples, ensemblewise, but we also do the same (on the particle ensemble sets) along the time axis (averaging independent but equivalent time cycles). Trial and error results indicate that the optimum values of  $\Delta t$  (or DCEN) lie between  $0.5 (D_{\text{avg}}/V_{\text{avg}})$  and  $2.0 (D_{\text{avg}}/V_{\text{avg}})$ .

The maximum number of particles ( $N$ ) that can be processed in each cycle is limited by the amount of storage available in the computer system. For the IBM 360/75 system

at the Illinois Computer Center, the storage of the data for 5,000 particles is easily accomplished. By playing Russian roulette, splitting, or similar techniques, the total number of artificial particles ( $N$ ) can be increased by an order of magnitude. The end of one time cycle  $t_i$  is the start of the next time cycle  $t_{i+1}$ . While keeping track of the particles in a current time cycle at one time, we simultaneously record the data information after each particle reaches the census and extract the output parameter values after all the particles have been processed. In this way the storage requirements are reduced to a minimum level possible.

The limitation on  $N$  by the storage requirement may be further relaxed by the following technique. Instead of recording and tallying the output information and clearing or reinitializing all parameter values at the end of every cycle, we record the accumulated information for  $K$  continuous, and similar but statistically independent time cycles, then renormalize and reinitialize. In this way the total number of particles processed is  $KN$ , and if processing of a large number of particles is desired, we can use a larger value of  $K$  ( $K > 1$ ) at the expense of more processing time. The reason for this is based on the fact that the system being simulated is time independent. We have introduced the time variable artificially so that ergodicity holds, namely ensemble averaging is the same as time averaging.

## 2. Tracing Procedure

The simulation is initialized by introducing  $N$  particles with energy  $E_{in}$  [sampled from an initial guess for the number density distribution vs. energy], with spatial position  $z_i^0$  (uniformly distributed throughout the spatial region), with direction  $u_i^0$  ( $u = \cos\theta$ ;  $\theta$ : angle of flight with  $z$  axis) which is uniformly distributed between  $-1$  and  $+1$  and with weight  $wt_i^0$ . Volume particles are emitted uniformly throughout the medium and isotropically in direction; namely, at  $t=t_0$ :

$$E_i^0 = E_{in} \quad (4a)$$

$$z_i^0 = r_i \cdot D \quad (4b)$$

$$u_i^0 = 2r_i - 1 \quad (4c)$$

where  $r_i$ : random number, uniformly distributed between 0, 1, or  $r_i \in (0, 1)$

$d$ : slab width or cylinder diameter

$wt_i, E_{in}$ : weight and particle energy determined from initial distribution.

Once the particle data are determined, the geometry routine is entered. This includes the determination of the distance to boundary  $d_B$ , the distance to collision  $d_{col}$ , and the distance to census time  $d_{cen}$  (clock distance equivalent to  $\Delta t$ ). These distances are defined by:



$$\begin{aligned}
 d_B &= [(j+1)\Delta Z - Z]/u \text{ if } u > 0 \\
 &= (Z - j\Delta Z)/|u| \text{ if } u < 0 \\
 &= DLB \text{ (a large number } DLB \gg \Delta Z) \text{ if } u \approx 0
 \end{aligned} \tag{5}$$

$$d_{col} = \lambda(E) |\ln r| = |\ln r| / \sum_t(E) \tag{6}$$

$$d_{cen} = D_{avg} \cdot DI = V_{avg} \cdot \Delta t \cdot DI \tag{7}$$

where r: random number uniformly distributed

j: zone number

$\Delta z$ : spatial interval size

$\sum_t$ : the total cross-section evaluated at the particles' energy

$D_{avg}$ : the average value of  $D_{col}$  of N particles of the previous cycle. It is approximately one mean free path corresponding to  $E_{avg}$ : the average energy of the cycle.

DI: input parameter for controlling the cycle time interval.

The next step depends on which of these distances is the smallest.

a. The System Without Electric Field ( $E_f=0$ )

i) If the smallest distance is  $d_{cen}$ : the particle is advanced in position and time to census time.

$$z' = z + u \cdot d_{cen} \tag{8a}$$

$$u' = u \tag{8b}$$

$$E' = E \tag{8c}$$

The particle data is stored, the energy is tallied, and

the density counter is incremented. Then we go on to process the next particle.

ii) If the smallest distance is  $d_B$ , the particle is advanced to the boundary such that:

$$z' = z + (d_B + 0.001 \cdot \Delta z) \cdot u \quad (9a)$$

$$j = \text{INT} (z' / \Delta z) + 1 \quad (9b)$$

$$d'_{\text{col}} = d_{\text{col}} - d_B \quad (9c)$$

$$d'_{\text{cen}} = d_{\text{cen}} - d_B \quad (9d)$$

where  $j$  is the zone number. If  $j > JZ$  or  $j < 1$ , then the particle is considered lost or has escaped from the system, and its history is ended. Otherwise the calculation and selection is repeated until the particle reaches the census time.

iii) If  $d_{\text{col}}$  is the smallest, the particle is first advanced to collision point and

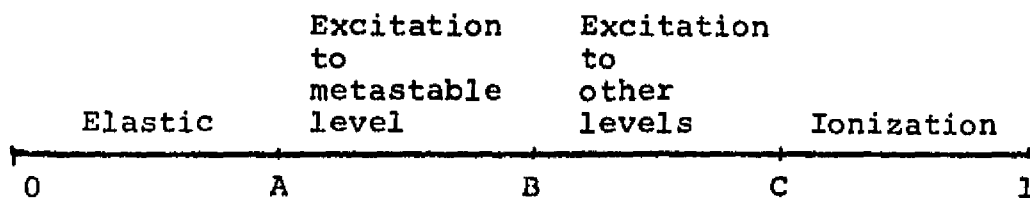
$$z' = z + u \cdot d_{\text{col}}$$

$$d'_{\text{cen}} = d_{\text{cen}} - d_{\text{col}} \quad (10)$$

Then the type of collision is specified. For specifying the collisional processes, we divide the whole energy range into two regions. Region I: energy range 0 to  $Ed1$  eV. Region II; energy range  $Ed1$  eV to  $E_0$  eV.  $Ed1$  and  $E_0$  are specified by input parameters and are referred to as the energy dividing line and source particle energy, respectively.

If the particle's energy is in the range of Region I, three kinds of collisions are possible: elastic collisions, electron-ion recombination, and superelastic collisions (where electrons hitting an excited atom pick-up the excitation energy). If the particle's energy falls in Region II, it can experience four kinds of collisions; namely, elastic collisions and three types of inelastic collisions--excitation to metastable level, excitation to other levels, and ionization. For ionization, the secondary electron so produced shares energy with the primary electron, and the energy of primary electron before collision less the ionization potential is shared between the two electrons. The partition of the shared energy could be determined according to Goodrich's experimental distribution curve (31). But based upon Thomas' (1) calculation experience, the way in which the energy is partitioned does not have a strong influence on any of the calculated mean properties. Thus, for computational simplicity, as Thomas did in all his calculations, we choose the partition ratio 9:1. The secondary electrons produced via ionization are considered as new particles introduced into the system. Thus they are followed from their point of birth just like other particles. This is equivalent to the generation method for processing branched trees of samples.

The type of collision is determined according to Figure 3 using a random number (uniformly distributed



$$A = P_{el}$$

$$B = P_{el} + P_{met}$$

$$C = P_{el} + P_{met} + P_{exc}$$

$$P_{el} + P_{met} + P_{exc} + P_{ion} = 1$$

Figure 3. The probability model for determining the types of collision.

between 0 and 1) to play the law of chances.

b. With Electric Field Present

The coordinate axis  $z$  is defined as antiparallel to the electric field  $E_f$ . In the presence of electric field  $E_f$ , the state of motion of an electron is specified by its speed  $V$  and the direction cosine  $\cos \theta$ , where  $\theta$  is the angle made by the line of flight with the  $z$ -axis at the time  $t$ . If an electron is in a state  $[E(V), u(\cos \theta), z, t]$ , it will assume (after a flight along the orbit by  $\Delta t$ ) the new state  $[E'(V'), u'(\cos \theta'), z + \Delta z, t + \Delta t]$ . These parameters are connected by the following relations:

$$V' \cdot \sin \theta' = V \cdot \sin \theta \quad (11)$$

$$mV' \cos \theta' = mV \cos \theta + eE_f \cdot \Delta t \quad (12)$$

$$1/2 mV'^2 = 1/2 mV^2 + eE_f \cdot \Delta z \quad (13)$$

$$E' = E + eE_f \cdot \Delta z \quad (14)$$

where  $m$  and  $-e$  are the mass and the electric charge of the electron, respectively. Here  $\Delta z$  is the projection of the path length  $L$  on the  $z$  axis. It can be obtained as follows;

$$\begin{aligned} \Delta z &= \int_0^{\Delta z} V_z dt = \int_0^{\Delta t} (V \cdot u + \frac{E_f}{m} \Delta t) dt \\ &= V \cdot u \cdot \Delta t + 1/2 \cdot \frac{E_f}{m} \cdot \Delta t^2 \end{aligned} \quad (15)$$

In carrying out the geometry routine, it is suitable to use time directly as the clock. The determination of the distance to collision  $d_{col}$  (DCOL) and the distance to census

$d_{cen}$  (DCEN) is accomplished as follows:

$$DCOL = \lambda |\ln r| \quad (16)$$

$$= 2 / (1/\lambda(E) + 1/\lambda(E')) \quad (17)$$

$$DCEN = \sqrt{V_y^2 + V_z^2} \cdot \Delta t \quad (18)$$

$$V_y = V \cdot \sin \theta \quad (19)$$

$$V_z = V \cdot \cos \theta + \frac{E_f}{2m} \Delta t \quad (20)$$

where  $E$ : the initial energy of the particle

$E'$ : the energy of the particle after flight time  $\Delta t$ .

The next step is to determine which of the distances is smaller. If  $DCEN < DCOL$ , then the particle is advanced to the new state

$$\begin{aligned} E' &= E + E_f \cdot \Delta z \\ z' &= z + \Delta z \\ u' &= \sqrt{E/E'} \cdot u + e E_f \cdot \Delta t \end{aligned} \quad (21)$$

If  $E' \leq 0$  or  $z' > D$ ,  $z' < 0$  then the particle is considered to have escaped from the system. If  $DCEN \geq DCOL$ , the particle is advanced to collision point.

$$\begin{aligned} \Delta t' &= \Delta t - DCOL \cdot \Delta t / DCEN \\ DCEN &= DCEN - DCOL \\ z' &= z + \Delta z \cdot DCOL / DCEN \\ E' &= E + \Delta z \cdot E_f \cdot DCOL / DCEN \end{aligned} \quad (22)$$

then the specific collision type must be determined.

### c. Types of Collisions and Cross Sections

For the elastic collisions, the particle scatters off with angle of deflection  $\theta$  and energy loss  $\Delta E$ , the following formula is used (36).

$$\Delta E = \frac{2m}{M} (E - E_n) (1 - \cos\theta_s) \quad (23)$$

where  $\theta_s$ : scattering angle (angle of deflection)

$E$ : particle energy

$E_n$ : target gas atom energy

$m$ : electron mass

$M$ : gas atom mass

$E_n$  is selected according to a Maxwellian distribution, and the  $\Delta E$  so formed may have either positive or negative values, indicating the loss or gain of energy respectively. For calculational simplicity, instead of selecting the deflection from the elaborate distributions, we use  $\pi/4$  or  $\pi/2$  deflection approximations, namely  $\theta_s = \pi/4$  or  $\pi/2$ , then  $\cos\theta_s = 0$  or  $0.70711$ . Another reasonable alternative is to assume that the direction after collision is equally probably in all directions (isotropic scattering). We select this direction at random with uniform distribution, and based on Eq. 23 calculate  $\cos\theta_s$  and  $\Delta E$ , respectively, where

$$\cos\theta = 2r_i - 1 \quad (24)$$

$$\theta_s = \theta - \theta_i$$

where  $\theta_i$ : incident angle

$\theta$ : outgoing angle after collision

For the three kinds of inelastic collision cross sections we use approximate formula based on experimental data as correlated by Itoh and Musha (2). The formulas are expressed in the form:

$$\Sigma(E) = \frac{B(E-E_i)}{A+(E-E_i)^2} + C \quad (25)$$

where the constants A, B, C, and  $E_i$  depend on the kind of collision and scattering angle. For the elastic cross section in helium, we use a single approximate formula for whole effective energy range based on Heylen and Lewis's paper (10), namely:

$$\Sigma_{el}(E) = 26 \exp(-0.04E) \text{ (cm}^{-1}\text{)} \quad (26)$$

For the recombination cross section for helium (35), we assume the simple form:

$$\begin{aligned} \Sigma_{RCB}(E) &= \mu(E) \cdot N^+ \\ \mu(E) &= 10^{-9}/V(E) \end{aligned} \quad (27)$$

where  $N^+$ : ion density

V: electron velocity

the superelastic cross section is

$$\Sigma_{sel}(E) = \frac{0.109(19.8-E)}{0.25 + (E-19.8)^2} \cdot \frac{N^+}{3 \cdot D_n} \quad (28)$$



where  $N^+$ : positive ion density

$D_n$ : gas atom density

$E$ : electron's energy

### 3. The Prediction and Correction of Dominant System Parameters

There exists a single parameter, the ion density distribution  $N^+(E, z)$ , which completely characterizes the nonlinearity of the system. This simplifies the problem and suggests the following predict-and-correct scheme. We assume that the average mean free path (or  $\Delta t$ ) and the ion density distribution  $N^+(E, z)$  in the present time cycle also applies to the next time cycle. In other words, in each time cycle we use the predicted values of  $\Delta t$  and  $N^+(E, z)$  from previous time cycle. At the end of present time cycle, we use the up to date information to renormalize and correct the value to be used for the next time cycle. As steady state is reached, the predicted and corrected values of the parameters converge. From the continuity equation

$$\nabla \cdot J = S - L \quad (29)$$

where  $J$ : the particle current flow in or out of the system

$S$ : the source

$L$ : the sum of losses

To start with, we assume that  $\nabla \cdot J = 0$ , so that

$$S = L = \alpha(E) N^+(E, z) N^-(E, z) \quad (30)$$

Here  $N^-(E, z)$  is the electron number density, the key quantity we are calculating. To initialize the run we have to predict  $N^+(E, z)$ . It is reasonable to initially assume this distribution is uniform in  $z$ , and a maxwellian distribution in  $E$ , i.e.,  $N^+(E, z) = N_0 M(E)$ . We also assume that  $N^-(E, z)$  is the same as  $N^+(E, z)$ . Then we obtain

$$\begin{aligned} S &= L = \alpha(E) N^+(E, z) N^-(E, z) \\ &= \alpha(E) [N^+(E, z)]^2 \end{aligned} \quad (31)$$

$$\text{or } S_0 \delta(E-E_0) = \alpha(E) N_0^2 M^2(E)$$

Integration of both sides gives

$$S_0 = N_0^2 \int_0^{E_0} \alpha(E) M^2(E) dE \quad (32)$$

$$. . N_0 = [S_0 / \int_0^{E_0} \alpha(E) M^2(E) dE]^{1/2}$$

This gives a starting value for the calculations. In later cycles the renormalization process involves integration of Eq. 29, with  $\nabla \cdot J \neq 0$ , i.e.,

$$\begin{aligned} \iint \frac{dJ(E, z)}{dz} dz dE &= \iint S(E, z) dz dE \\ &- \iint \alpha(E) N^+(E, z) N^-(E, z) dz dE - J_{\text{esc}} + J_s \end{aligned}$$

which reduces to

$$\begin{aligned} J(Z=d) - J(z=0) &= S_0 d - \iint \alpha(E) N^+(E, z) N^-(E, z) dz dE \\ &- J_{\text{esc}} + J_s \end{aligned}$$

$$N_i^2 = \frac{S_0 d + J_s - [J(z=d) - J(z=0)] - J_{esc}}{\iint \alpha(E) M(E) n^-(E, z) dz dE} \quad (33)$$

where  $i$  refers to  $i$ th time cycle

$J_{esc}$ : current due to escaped electrons

$J_s$ : current due to secondary electrons produced by ionizations

$M(E)$ : Maxwellian distribution (normalized)

$n^-(E)$ : electron number density distribution (normalized)

$S_0$ : source rate

$N_i$ : ion density distribution (normalized)

The value of  $N_i$  replaces  $N_0$  in equation (33) at later time cycles. The value of  $N_i$  ( $N^+$ ) enters the formula for the recombination cross section and superelastic cross section, thus it directly affects the rates of recombination and superelastic collision. Indirectly, it affects the balance of the system, or equivalently the rate of convergence of the Monte Carlo calculation. An adaptive correction procedure is used here. It is based on the value of the "balance factor" BF, defined as the ratio of the particles introduced into the system by source and ionization processes to the particles removed from the system due to recombination and leakage. We adopt a over-relaxation type of correction technique, the over-correction is used to speed up the over-all convergence rate. As defined before, if  $BF > 1$ , the system has more particles introduced

into than eliminated, but if  $BF < 1$ , the opposite is true. The over-correction factor is proportional to  $(BF)^k$ , i.e., to the  $k$ th power of the balance factor. The value of  $k$  initially chosen was  $k=2$ , but an optimum value can be obtained by trial and error.

# CHAPTER III

## VARIANCE REDUCTION TECHNIQUES

### A. General Background

The Monte Carlo method (or method of statistical trials) consists of solving various problems of computational mathematics by the construction of some random process for each such problem, with the parameters of the process set equal to the required quantities of the problem. These quantities are then approximated by observations of the random process and the computation of its statistical characteristics, which are approximately equal to required parameters [from Schreider (15)]. Every Monte Carlo computation that leads to quantitative results may be regarded as estimating the value of a multiple integral. Suppose we have  $M$  random numbers in the computation system. The results will be a vector valued function  $R(\xi_1, \xi_2, \dots, \xi_M)$  involving the sequence of random numbers  $\xi_1, \xi_2, \dots, \xi_M$ . This is an unbiased estimator of  $\int_0^1 \dots \int_0^1 R(x_1, x_2, \dots, x_M) dx_1 \dots dx_M$ . For the sake of simplicity, we take the one-dimensional integral as a standard example. Let

$$\theta = \int_0^1 f(x) dx \quad (34)$$

where  $f \in L^2(0,1)$  or  $\int_0^1 [f(x)]^2 dx$  exists.

Define the relative efficiency of two Monte Carlo methods as follows. Let the methods call for  $n_1$  and  $n_2$  units of computing time, respectively, and let the resulting

estimates of  $\theta$  have variances  $\sigma_1^2$  and  $\sigma_2^2$ . Then the efficiency of method 2 relative to method 1 is defined as

$$\frac{n_1 \cdot \sigma_1^2}{n_2 \cdot \sigma_2^2} \quad (35)$$

Equation 35 is the product of two terms, the variance ratio  $\sigma_1^2 / \sigma_2^2$  and the labor ratio  $n_1/n_2$ . If  $\gamma_1, \gamma_2, \dots, \gamma_N$  are independent random numbers with uniform distribution between 0, 1 or  $\gamma_i \in (0,1)$ ; then the quantities  $f_i = f(\gamma_i)$  are the independent random variates with expectation  $\theta$ . Therefore

$$\bar{F} = \frac{1}{N} \sum_{i=1}^N f_i \quad (36)$$

is an unbiased estimator of  $\theta$ , and its variance is

$$\frac{1}{n} \int_0^1 (f(x) - \theta)^2 dx = \frac{\sigma^2}{n} \quad (37)$$

Accordingly the quantity  $\bar{F}$ , which has been determined by observation of the random process, is approximately equal to the required quantity  $\theta$  (or in other words, it is an unbiased estimator of  $\theta$ ) with a probability which can be made as close as required to unity if a large number of trials is practical. We refer to the estimator  $\bar{F}$  as the crude Monte Carlo estimator of  $\theta$  (commonly referred to as Monte Carlo estimation).

Let us introduce another even less efficient method, namely hit or miss Monte Carlo. Suppose that  $0 \leq f(x) \leq 1$

when  $0 \leq x \leq 1$ . We may draw the curve  $y = f(x)$  in the unit square  $0 \leq x, y \leq 1$ , and  $\theta$  is the proportion of the area of the square beneath the curve. Stating this formally, we write

$$f(x) = \int_0^1 g(x,y) dy \quad (38)$$

where  $g(x,y) = 0$  if  $f(x) < y$   
 $= 1$  if  $f(x) \geq y$

Then

$$\theta = \int_0^1 \int_0^1 g(x,y) dx dy \quad (39)$$

The estimator

$$\bar{g} = \frac{1}{n} \sum_{i=1}^n g(\xi_{2i-1}, \xi_{2i}) = \frac{n^*}{n} \quad (40)$$

$n^*$  is the number of occasions where  $f(\xi_{2i-1}) \geq \xi_{2i}$

In other words, we take  $n$  trials at random in the unit square, and count the proportion of them which lie below the curve  $y = f(x)$ .

Historically, a hit or miss method was first propounded in the explanation of Monte Carlo techniques. It is the easiest to understand but one of the least efficient techniques. The rate of convergence and the variance (statistical error) is in general proportional to  $1/\sqrt{N}$  where  $N$  is the number of samples. It is obviously impractical to gain a significant improvement of accuracy by merely increasing  $N$ . Therefore more efficient techniques for reducing the variance and/or increasing the efficiency

are very important. Monte Carlo experimentalists need wide experience and background in both the mathematics and the physics of the problem, and they have to exercise considerable ingenuity in distorting and modifying problems in the pursuit of variance reduction techniques. Statistical and inferential procedures are also important in order to extract the most reliable conclusions from the observational data.

The basic requirement of the Monte Carlo calculations is to establish how random numbers may be used to sample a function that describes, in a probabilistic fashion, a physical event. The procedure by which this is done follows from the fundamental principle of Monte Carlo (16), which may be stated as follows. If  $p(x)dx$  is the probability of  $x$  lying between  $x$  and  $x+dx$ , with  $a \leq x \leq b$

$$\int_a^b p(x)dx = 1 \quad (41)$$

$$\text{Then } r = P(x) = \int_a^x p(y)dy \quad (42)$$

determines  $x$  uniquely as a function of the random number  $r$  which is uniformly distributed on range  $(0, 1)$ . The quantities  $p(x)$ ,  $P(x)$  are the probability density and the probability distribution functions, respectively.

With the development of large computers, the use of the Monte Carlo method in a wide range of problems has



increased rapidly. In particular, the method is applied to various physical problems such as plasmas (15), radiative transfer problems (21), neutron transport and nuclear reactor physics (29), etc.

The random walk type simulation can be briefly described in the following manner. One first introduces the basic physics of the problem into the computer in a probabilistic fashion. A system of coordinates and boundaries are defined and then, as a computer experiment, particles are released from the source. These particles are traced as they diffuse through a prescribed medium following the probabilistic interaction laws. The particles are followed until they escape from the medium or are absorbed or are converted to the thermal field.

The parameters pertinent to the evaluation of the desired quantities are recorded. One continues processing additional particles until adequate statistical estimates of the quantities of interest have been obtained.

When doing a Monte Carlo problem one focusses attention on three main topics. They are:

1. The development of the analogy for the probability processes (simulation).
2. The generation of sample values of the random variables on a given computing machine.
3. The design and use of variance reduction techniques.

The variance reduction methods are often strongly dependent on the probabilistic model, thus the greatest gains are often made by exploiting specific details of the problem, rather than by routine application of general principles.

#### B. Importance Sampling

The general idea of importance sampling is to draw samples from a distribution other than the one suggested by the problem. Then an appropriate weighting factor is introduced to correct the biasing caused by changing the original distribution. If done correctly the final results are essentially unbiased. The object is to concentrate the distribution of sample points in the parts of the region that are most "important" instead of spreading them out evenly. Thus, the probability of sampling from an interesting region is increased; the probability of sampling from unimportant or less interesting region is correspondingly decreased. Stating this formally, we have

$$\theta = \int_0^1 f(x) dx = \int_0^1 \frac{f(x)}{g(x)} g(x) dx = \int_0^1 \frac{f(x)}{g(x)} dG(x) \quad (43)$$

for any functions  $g$  and  $G$  satisfying

$$G(x) = \int_0^x g(y) dy \quad (44)$$

where  $g$  must be a positive valued function such that

$$G(1) = \int_0^1 g(y) dy \quad (45)$$

if  $\gamma$  is a random number sampled from the distribution  $G$ , then  $f(\gamma)/g(\gamma)$  has the expectation  $\theta$  and variance

$$\sigma_{f/g}^2 = \int_0^1 \left( \frac{f(x)}{g(x)} - \theta \right)^2 dG(x) \quad (46)$$

We notice that if  $g(x) = cf(x)$ , or if  $g$  is proportional to  $f$  and if  $c = \frac{1}{\theta}$  then  $\sigma_{f/g}^2 = 0$ . This perfect situation does not exist, since we do not know  $\theta$ . We always obtain an unbiased estimate for positive function  $g$ . Our object is to select  $g$  to reduce the standard error of our estimate.

Because the estimate is the average of observed values of  $f/g$ , we choose  $g$  such that  $f/g$  is as constant as possible in order to achieve a small sampling variance. While we intend  $g$  to mimic  $f$ , we generally restrict our choice of  $g$  to functions simple enough to allow analytic integration.

For the present plasma problem, the importance sampling technique is developed as follows: during the particle's (electron's) random walk, it suffers a large number of collisions with gas atoms, and these collisions may be elastic or inelastic or of other kinds. The probabilities of their occurrence depend on their cross section  $\Sigma_i$ . We have

$$\Sigma_t = \Sigma_1 + \Sigma_2 + \dots + \Sigma_M \quad (47)$$

where  $\Sigma_i$ : the cross section for  $i$ th process

$\Sigma_t$ : total cross section or the summations of the cross sections of all the possible processes that the particle may experience.

Dividing both sides of equation 47 by  $\Sigma_t$ , we obtain:

$$1 = p_1 + p_2 + \dots + p_M \quad (48)$$

where  $p_i = \Sigma_i / \Sigma_t$  the probability of  $i$ th process occurring as the particle experiences a collision. If we introduce weights  $w_i$  into equation 47 we obtain

$$\Sigma_t' = w_1 \cdot \Sigma_1 + w_2 \cdot \Sigma_2 + \dots + w_M \cdot \Sigma_M \quad (49)$$

Similarly we have

$$1 = p_1' + p_2' + \dots + p_M' \quad (50)$$

After the application of the weights, we put more weights on the interesting or more important events. Then their artificial probabilities are greater, therefore we can get better estimates of the important parameters. Coincidentally, such a linear weighting scheme can only be applied to the elastic collision process in our system. The other collision processes involve appreciable changes in the particle's energy (catastrophic collision). Also cross sections are energy dependent, these nonlinear effects prevent the application of above mentioned weighting scheme to inelastic collision processes. By looking at the cross section values, in the low energy region, the elastic cross section is by

far the largest by order of magnitude. Others have shown that the elastic process does not strongly influence the estimated parameters that we are interested in (1, 2). As an electron makes an elastic collision, it loses only a small amount of energy while making a small angle deflection. In fact, most other workers simply neglect elastic losses if an electric field is present (1, 2, 10). In our case, with  $E_f=0$ , elastic effects must be included. In the low energy region ( $\approx 0$  eV to 20 eV), the energy of a majority of the particles lies below 1 eV and over 90% collisions are elastic collisions, and only a few collisions are inelastic.

The collision process subroutines are a time-consuming part of the simulation calculations. If we can suppress the elastic processes by a factor of  $k_{el}$  the chances of all other processes are increased by approximately the same factor. Therefore we convert more computing time to processes which contribute more strongly to the quantities of interest. The following formulation provides the theoretical support for this. In lieu of equations 47-50, the  $p_i$  are the original probabilities (undistorted probabilities) while the  $p'_i$  are the weighted probabilities. Then the correction factors are introduced as follows; at energy  $E$ , we have

$$C_1 = P_1/p'_1, C_2 = P_2/p'_2 \dots C_M = P_M/p'_M \quad (51)$$

$$\text{or } C_i = p_i/p_i' = \frac{\sum_i \cdot \sum_t'}{\sum_t \cdot \sum_i w_i} = \frac{\sum_t'}{\sum_t} \frac{1}{w_i} \quad (51a)$$

If we let  $w_1 = \frac{1}{k_{el}}$ ,  $w_2 = w_3 = \dots w_M = 1$

then from (51) we have

$$C_1 = \frac{\sum_1 \cdot \sum_t'}{\sum_t \sum_1 w_1} = \frac{\sum_t'}{\sum_t} \frac{1}{w_1}, \quad C_i = \frac{\sum_t'}{\sum_t} = 1, \quad i \neq 1 \quad (52)$$

let  $\sum_1 = \sum_{el}$  namely we introduce weight on elastic cross section only. Since  $\sum_{el}(E) \gg \sum_i(E)$ ,  $i > 1$ ,  $\forall_i$  then

$$\frac{\sum_t'(E)}{\sum_t(E)} \approx \frac{\sum_{el}'(E)}{\sum_{el}(E)} \quad w_1 = \frac{1}{k_{el}} \quad (53)$$

$$\text{and } C_1 \approx w_1 \frac{1}{w_1} = 1, \quad C_i \approx \frac{1}{k_{el}} \quad (i \neq 1) \quad (53a)$$

Equation 53 holds for low energy region, which is approximately independent of energy.

### C. Russian Roulette and Splitting

In general, the sampling is done such that we examine the sample and classify it as being in some sense "interesting" or "uninteresting". We are willing to spend more than average amount of work on the interesting samples and on the contrary, we want to spend less effort on the uninteresting ones. This can be done by splitting the interesting samples into independent branches, thus resulting in more of them, and by killing off some of the uninteresting ones. The first process is splitting and the second Russian Roulette.

The "killing off" is done by a supplementary game of chance. If this game is lost, the sample is killed; if it is won, the sample is counted with an extra weight to make up for the fact that some other samples have been "killed". The game has a certain similarity to the Russian game of chance played with revolvers - whence the name.

For the present plasma system we divide the whole energy spectrum into two regions, the low energy region I (0 eV to ED1 eV) and the high energy region II (ED1 to  $E_0$ ). Here  $E_0$  is energy of source particles, which is the highest energy the particle can have. ED1 is the energy region dividing line, which can be chosen as any value between 5 eV and 100 eV depending on the initial guessed energy distribution. In region I, the number density is high, and the elastic process is the dominant process. As we mentioned earlier, the elastic process does not contribute much, so Region I is classified as uninteresting. Each particle in this region (particles with energy less than ED1) is assigned a weight  $k_1$  ( $k_1 \geq 1$ ). In other words, each particle plays the role of  $k_1$  particles.

In region II, classified as interesting region, each particle carries a weight  $w_2$  ( $w_2 = 1/k_2$ ,  $w_2 \leq 1$ ). Namely, each particle in this region is split into  $k_2$  particles with weight  $w_2$ .

As particles cross over to other regions, care must be taken to account for killing off and splitting effects.

For instance, a particle in region I may acquire enough energy, either from electric field or by superelastic collision with the excited atoms, to cross over to region II. Then we should generate  $k_3$  particles ( $k_3 = k_1 \cdot k_2$ ) in region II with the same energy.

Particles from the high-energy region may lose part of their energy by inelastic collision such that they cross over to the low-energy region. Then killing off should be effective. For every  $k_3$  such particles that cross over from region II to region I, only one particle survives. The values of  $k_1$  and  $k_2$  can be chosen according to the total number of particles in the system to be processed, the source rate, and related accuracy requirements. They are input parameters.

#### D. Initial Guess

Monte Carlo methods comprise that branch of experimental mathematics which is concerned with experiments on random numbers. The user, like the experimental physicist needs theory and knowledge to give structure and purpose to his experiments, and as experimental work provides growing insight into the nature of a problem and suggest appropriate theory. Good Monte Carlo practice should keep this relationship as a general maxim.

The basic procedure of the Monte Carlo method is the manipulation of random numbers, but these numbers should not



be employed prodigally. Each random number is a potential source of added uncertainty in the final result. Thus it will usually pay to scrutinize each part of a Monte Carlo experiment to see whether that part can not be replaced by an exact theoretical analysis contributing no uncertainty. In other words, exact analysis should replace Monte Carlo sampling wherever possible. Sometimes reliable intuition would aid in increasing the efficiency of Monte Carlo calculations. For the plasma system we are simulating, the straight-forward approach would be to follow  $N$  source particles on their life time history, until they lose energy sufficient by collisions to reach a thermal equilibrium. Over 50 inelastic collisions generally occur, and the probability (inelastic cross sections) for inelastic processes is low. Consequently, the particles travel a relatively large distance which consumes a good amount of time. Thus a large number of time cycles would be required before the system reaches steady state. To avoid this, we use our intuition, to start off with a guessed initial distribution. The electron energy distribution  $f(E)$  (or electron number density distribution) is a function of the electron's energy  $E$ . Initial particle data is generated using this distribution, and this can save considerable irrelevant computations. This idea is simple, yet very useful and effective. Without an initial distribution, the problem would require hundreds of time cycles to

converge. However, with a reasonable guess, good results have been obtained in as little as tens of time cycles. Of course, the better the initial guess, the faster the convergence.

### E. Antithetic Variates

When we estimate an unknown parameter  $\theta$  by means of an estimator  $t$ , we may seek another estimator  $t'$  having the same (unknown) expectation  $\theta$  as  $t$  but a strong negative correlation with  $t$ . Then  $1/2(t + t')$  will be an unbiased estimator of  $\theta$ , and its variance is

$$\text{Var}[\frac{1}{2}(t+t')] = \frac{1}{4}\text{var}(t) + \frac{1}{4}\text{var}(t') + \frac{1}{2}\text{cov}(t,t') \quad (54)$$

where  $\text{cov}(t,t')$  is negative, and

$$\text{var}(t) = \mathcal{E}[(t-u)^2]$$

$$\text{cov}(t,t') = \mathcal{E}[(t-u)(t'-u')]$$

$u, u'$  are the mean of  $t, t'$  respectively, and  $\mathcal{E}(x)$  is the expectation of the random variable  $x$ .

It is possible to make  $\text{var}(t+t')$  smaller than  $\text{var}(t)$  by suitably selection  $t'$ . For example,  $1-r$  is uniformly distributed whenever  $r$  is, and if  $f(r)$  is the unbiased estimator of  $\theta$ , so is  $f(1-r)$ . When  $f$  is a monotone function,  $f(r)$  and  $f(1-r)$  will be negatively correlated. Thus we could take

$$\frac{1}{2}(t+t') = \frac{1}{2} t(r) + \frac{1}{2} f(1-r) \quad (55)$$

as an estimator of  $\theta$ .

The name and idea of antithetic variates was first introduced by Hammersley and Morton (13, 14) during attempts to improve Buffon's needle experiment for estimating the value of  $\pi$  (3.1415926).

The main idea of antithetic variates is based on the underlining theorem [from Hammersley and Morton (12)] which can be stated informally as follows: "Whenever we have an estimator consisting of a sum of random variables, it is possible to arrange for a strict functional dependence between them, such that the estimator remains unbiased, while its variance comes arbitrarily close to the smallest that can be attained with these variables". As the name implies, antithetic variates are the set of estimators which mutually compensate each other's variations. Essentially, we rearrange the random variables by permuting finite subintervals in order to make the sum of the rearranged functions as nearly constant as possible. From the practical viewpoint, the mathematical conditions imposed by antithetic variates to the Monte Carlo calculation are quite loose and flexible. It is relatively easy to find a negatively correlated unbiased estimator to produce an efficient variance reduction scheme.

A system of antithetic variates is obtained by simple transformations based on a stratification of the interval (12). The transformations are constructed in such a way

that the efficiency of the method will increase as a higher power of  $M$  (the number of antithetic variates taken). The transformations considered are linear combinations of the values of the function  $f$  at a number of points. Thus a correlated stratified sampling technique is developed, and its efficiency generally exceeds that of crude sampling by a factor  $M^3$ .

The following useful transformations are obtained in this manner:

$$\begin{aligned} t_1 &= \frac{1}{n} \sum_{i=1}^n \{ \alpha f(\alpha x_i) + (1-\alpha) f[\alpha + (1-\alpha)x_i] \} \\ &= \frac{1}{n} \sum_{i=1}^n G_\alpha f(x_i) \end{aligned} \quad (56)$$

$$\begin{aligned} t_2 &= \frac{1}{n} \sum_{i=1}^n \{ \alpha f(\alpha x_i) + (1-\alpha) f[1 - (1-\alpha)x_i] \} \\ &= \frac{1}{n} \sum_{i=1}^n \mathcal{L}_\alpha f(x_i) \end{aligned} \quad (57)$$

The value of  $\alpha$  is between 0 and 1, the variance  $\text{var}(t_j)$  has a minimum at some value of  $\alpha$ . An adequate rule of thumb is to choose  $\alpha$  by finding the root of

$$f(\alpha) = (1-\alpha) f(1) + \alpha f(0) \quad (58)$$

Another useful transformation is

$$t_3 = \frac{1}{m} \sum_{j=0}^m f\left(\frac{r+j}{m}\right) = U_m f(r) \quad (59)$$

In general, the variance of crude Monte Carlo is  $\text{var}(t) = O(m^{-1})$  while the variance of  $t_3$  is  $\text{var}(t_3) = O(m^{-2M})$ , ( $M > 1$ ). The simplest form of antithetic transformations have proved quite successful in neutron transport calculations.

The general notion and a simple form of antithetic variates are applied in our present calculations. High efficiency gain can be obtained easily if it is used sparingly and at judicious points in the computation. We have used the random number  $r_i$  ( $r_i \in (0,1)$ ) most of the time for playing the law of chances, and  $1-r_i$  is another random number of the same kind which is the image of  $r_i$ . The two are used together in good many places where repeated use of  $r_i$  is required. This serves the purpose of two antithetic variates and cuts the labor in half (same random number  $r_i$  being used twice). Equations 55 and 59 have been used in the initial data preparation program and source routine where initial particles states  $(E_i, u_i, z_i, wt_i)$ ,  $E_i$  are generated according to input initial distribution. For instance:

$$\begin{aligned} u_i &= 2r_i - 1, \quad u_{i+1} = 2(1-r_i) - 1 \\ z_i &= r_i \cdot d, \quad z_{i+1} = (1-r_i) \cdot d \\ E_i &= E_k + r_i \cdot \Delta E_k, \quad E_{i+1} = E_k + (1-r_i) \cdot \Delta E_k \\ E_i &= E_0 \text{ for source particles.} \end{aligned} \tag{60}$$

Here  $k$  is an energy index, or  $k$ th energy interval.

Wherever  $r_i$  appears,  $r_i$  and  $1-r_i$  are used interchangeably.

Thus we are generating source particles from diametrically opposite directions, or from symmetric locations.

#### F. Double History (Double Processing) Technique

In the particle tracing routine, using the general idea of antithetic variate technique, we have adopted a "so-called" double history processing technique which is described as follows;

a) In the course of tracing the particle's history, in each time cycle, we use the transformed random number  $\lambda \cdot |\ln r|$  to decide the distance to collision  $d_{col 1}$  and go on until the particle reaches census time. However, we process the same particle twice under the identical initial condition except that we use  $\lambda \cdot |\ln(1-r)|$  as the distance to collision  $d_{col 2}$ . Both sets of history are tallied but only one set of particle data is stored. This way each particle is processed twice independently but correlated in such a way that they compensate each other's variation (variance).

b) Another alternative for double history processing is, to play a dual direction random walk instead of dual collision distances. We process the same particle twice in such a way it looks like there are two identical particles emerging into the system in two diametrically opposite directions. Again both histories are tallied, but only the original particle's data is stored.

The two versions of the double history processing technique increase the labor (computation time) by a factor of 1.2 to 1.5. The efficiency gain is difficult to specify. At least it is certain that the gain is more than double the total number of the particles in the system.

The purpose of dual processing is to compensate each particle's variation and minimize the variance (statistical errors). One set of particle data in dual processing is thrown away, while the other set is stored just like normal processing. Which one of the two resulting sets is saved is completely random.

## CHAPTER IV

### MONTE CARLO SIMULATION CODE

#### A. Introduction

This chapter presents the complete Monte Carlo simulation program, including an input, output, main flow chart, and the background theory and techniques incorporated. This program implements the condensed history approach that was described in Chapter II together with various variance reduction techniques described in Chapter III.

The basic Monte Carlo calculation can be described most easily by reference to the simplified flow chart shown in Figure 4. By means of this flow chart we shall follow the particle tracing process, making relatively brief comments on the various steps implied by each box, on each routine, and on the main program variables. The program is written in FORTRAN IV and is designed to be run on IBM 360/75 under the HASP-MVT system.

#### B. Calculational Procedure and Program Description

To obtain an overall view of the calculations, we note that there are three basic steps involved in the simulation process. The first is the initialization of all parameter values, i.e., prepare the particle data and predict the unknown dominate parameter values. The second is the main loop of particle tracing which generates census information. The third uses the census information and takes tallies to estimate the histograms (electron energy or number density



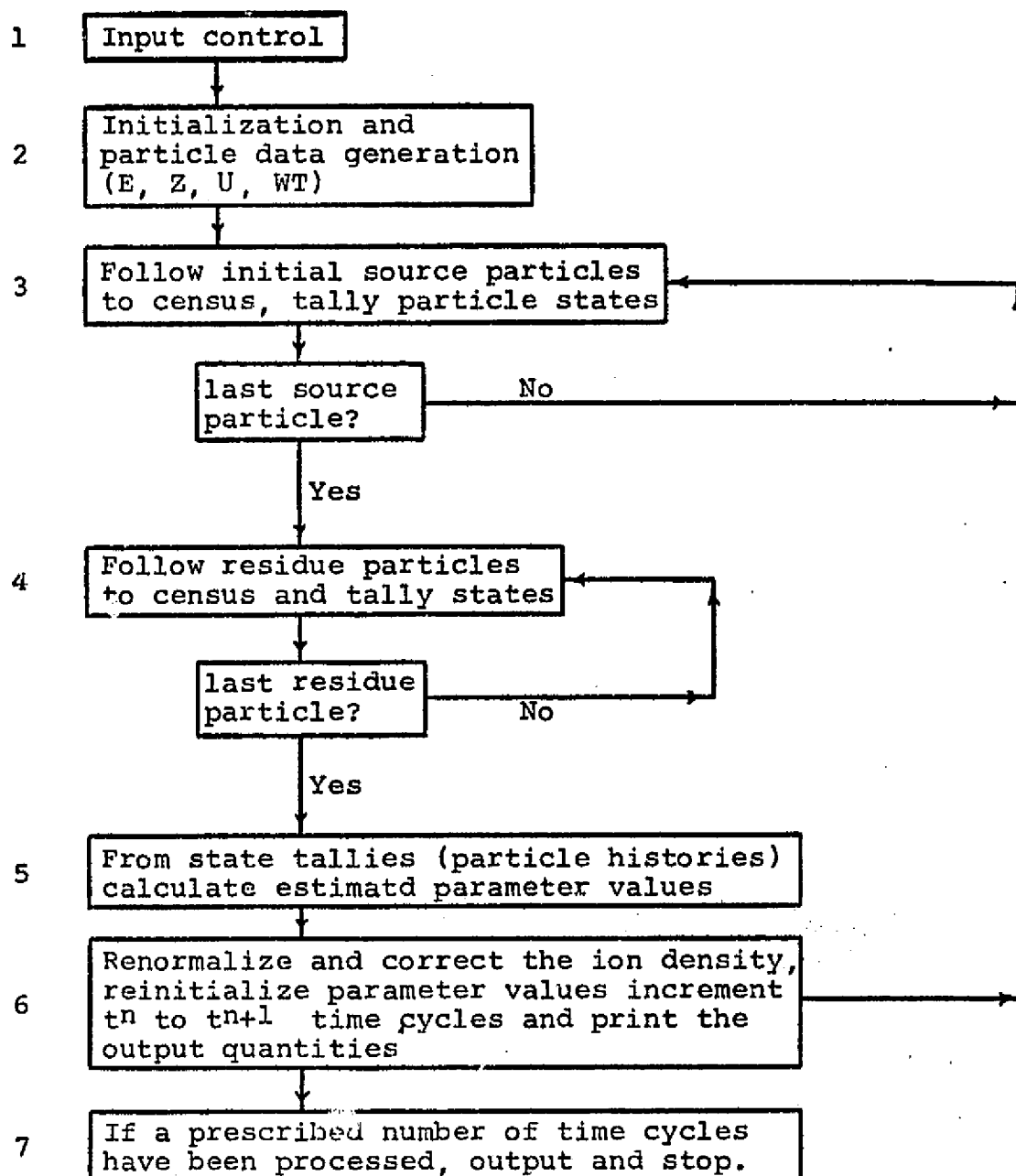


Figure 4. Flow chart of the Monte Carlo simulation code.

frequency distribution vs. energy), and parameter values. Finally the dominate parameter values are corrected and the program returns to the first step.

Let us follow the logic flow diagram (Figure 4) in the order of the number marked on the left hand side of the boxes, and describe their individual functions.

1. Input Control: The main program is usually compiled and the loaded version is then stored on disk storage, ready for execution. In order to offer a wide flexibility, the following input parameters are used to control the program runs.

a) NDISK: program start control. The program can start from the beginning (cycle 0), or it can start from the stopping point of the previous run. This feature enables long runs to be broken up into several short runs. It also offers debugging flexibility and economy of computer time. While a long run may run up to 150 time cycles, we can always break it up into multiple numbers of fives and tens cycles. Thus we can adjust other parameter values on a cut and try basis without wasting much computing time. The input parameter for this control purpose is NDISK: input 0 for starting from very beginning; input 1 for starting from the latest tenth cycle of last

run; input 2 for starting from the latest fifth cycle.

- b) K1, K2, and ED1: Regional weights and dividing energy. The Russian Roulette and splitting variance reduction techniques use three input parameters to control the amount of killing off and/or splitting of particles. This also determines where to divide energy regions for the Russian Roulette and splitting, assuming that the option to use Roulette is selected. The input parameters are K1: the weight of particles in region I ( $K1 \geq 1$ ), K2: the inverse of the weight  $w2$  of particles in region II where  $w2 = 1/K2$  ( $K2 \geq 1$ ), ED1: the energy dividing line, namely (0 to ED1) is region I, (ED1 to  $E_0$ ) region II where the energies are in electron volts.
- c) NTC: The total number of timecycles to be run.
- DTC: The input parameter to control the length of census time  $\Delta t$  of the time cycles. In terms of distance travelled by the particles,  $\Delta t$  is equal to a fraction or multiple of one mean free path, corresponding to average energy of particles in the system.

- d) NTP: The total number of particles that actually exist in the simulation system.
- NTW: The artificial\* total number of particles in the system during Russian Roulette and splitting games.
- NS: The number of source particles injected into the system during each time cycle.
- e) KELC: The input parameter that controls the importance sampling option of the program, i.e., the multiplying factor used to suppress the elastic cross section (reduces the probability of elastic process by a factor KELC).
- f) EDP: The input value of the ratio of the electric field strength to the pressure.
- PDPO: The input value for the reduced pressure of the system ( $P/p_0$  where  $p_0 = 1$  torr).
- g) NIP: The input parameter used to control double processing: NIP = 2 for double processing, NIP = 1 for single processing.
- IUSN: IUSN = 1 for dual-collision type double processing  
IUSN = -1 for dual-direction type double processing

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\*The total number modified by scaling and weighting factors.

- h) D: The input value of the width of the slab (diameter of the cylinder) which defines the external boundaries of the medium.
- JZ: The number of zones the slab is divided into, or the number of internal zones.
- i) NCUM: Input control to give cumulative results (time average) over the time cycles or a cumulative average over NCUM time cycles.
- NAVP: Input constant to control the power of  $\Delta t$ , (or  $\Delta D$ ) namely  $\Delta t_i^{**NAVP}$  is the weight of the cycle to be used for cumulative results. These two parameters are used to implement antithetic variates techniques along the time axis.
- j) EO: The energy of source particles.
- SR: The rate of the source particles entering the system.
- k) FID: An array that contains the initial guess distribution (number density vs. energy).
- l) NUSC: An input parameter used to control the up-scattering elastic collision process; NUSC = 1: with upscattering, NUSC = 0: without upscattering.

2. Data preparation: The initialization is entered only once during the calculation, where the time cycle is 0 (or at time  $t_0$ ). After the input phase, all the pertinent

parameters are set to their proper permanent or temporary values. Before entering the next phase to trace the particle histories, several dominate parameter values must be predicted. The data preparation routine is called in to generate particle data ( $E_i$ : energy,  $u_i$ : direction of travel,  $Z_i$ : position,  $Wt_i$ : weight) based on the guess for the initial distribution (generate machine particles). Then the average energy of the system is calculated and is used to predict the census time ( $\Delta t$  or equivalently DCEN) as well as the number of ions( $N^+$ ).

3. Particle tracing: With the point of origin, direction of propagation, and energy (velocity) known for the initial source particles, the next phase follows the particles through a series of stochastic interactions until the census time is reached, and the particle data and parameters values are tallied and updated. During the course of tracing the particle histories, all the relevant events are recorded and later on can be interrogated to provide the output quantities of interest. This section of the program is the central part of the main program. It contains the main loop for tracing all the particles in the system in a self consistent way. Most of the variance reduction techniques are incorporated here, including double processing, Russian Roulette, and splitting, as well as the various versions of

importance sampling techniques.

4. Secondary particle tracing: The residue particles, i.e., particles generated by ionization collision process (the secondary electrons so produced) are treated in all respects like the regular source particles, except that their starting positions and time coordinates as well as directions are those associated with the particles that generated them. These residue particles may go further and produce other residue particles (secondary electrons) if their energy is high enough, thus forming a short production chain. The probability of such chain production is small however, since the average energy of the secondaries is quite low. The end of the main loop comes after all such residue particles are processed.
5. Data calculations: The details of the phase of the computations depend upon the specific goals and character of study. Our primary interest is the particle energy distribution (or number density vs. energy). This frequency distribution and other pertinent parameters are estimated (or calculated) using the stored data from various parameter counters in the main loop (tallied information) as well as tallied histories. The quantities of interest include local values (for the current time cycle only) and the cumulated values from several time cycles. After these calculations, the next phase is to prepare for the processing of new time

cycle.

6. Normalization and printout: Up to this stage, all the output quantities have been calculated. Next the output information is printed out, and at the same time, if the cycle number is the multiple of five, all the particle data and related parameter values are stored on the disk. Then the next run can be read off of the disk and the process can proceed from where we left off the last time. All the parameter values are again initialized, and new source particles are generated for the next time cycle. Particles that have escaped from the system or have been absorbed create holes in the particle data storage which are filled up by the new source particles and the residue particles produced during the time cycle. Before going on to the next cycle, the ion density is renormalized and corrected in a semiadaptive way according to a balance factor BF. The ion density distribution parameter  $N^+$  is a dominate parameter since it affects several processes such as the rate of recombination and the rate of super-elastic collision (refer to Equations 27 and 28). It is corrected in such a way as to accelerate convergence, till the system reaches a balanced situation. Then increment time from  $t^n$  to  $t^{n+1}$ . The output quantities are the following:

- a) The kind of gas in the medium, source particle energy, pressure, slab thickness, and  $\Delta t$  the time interval of this cycle.



- b) The number of time cycles, the time in seconds, the number of source particles generated during the cycle time, and the average energy of the particles in the system.
  - c) The number density distribution vs. energy, the single cycle distribution, the cumulated distribution, the number-density deviation, the average-energy deviation, the average energy, and the energy interval vs. energy.
  - d) The number density and the average energy vs. z (position zone).
  - e) The ion distribution ( $N_+$ ), the number of ionizations ( $N_{IZN}$ ), the number of escaped particles ( $NES$ ), the number of absorbed or recombined particles ( $NAS$ ), the number of super-elastic collisions ( $NSE$ ), the number of excitations to metastable levels ( $NMET$ ), and the number of excitations to other levels ( $NEXC$ ).
  - f) All the input control parameters as listed in section IV-1.
  - g) The electric field strength ( $EF$ ), the source rate, the W-value, the average energy of escaped particles, and the average energy of absorbed particles.
  - h) The number check on the density distributions.
- All the above outputs include both current cycle values and the cumulated values over the past cycles.

7. Final checks: The last stage is to check whether the prescribed number of time cycles have been processed, adequate statistics have been collected, and all the useful data have been stored on the disk. If so, the program stops.

Further programming details are provided in Appendix

## CHAPTER V

## RESULTS

A. Convergence and Reliability

The present Monte Carlo Model is a predict-correct iterative scheme, with semiadaptive control ability. For such an iterative procedure, our main criteria for the evaluation and determination of convergence are based on the following inferences:

1. The balance factor (BF) as defined in Chapter II (Section II-B-3) signifies whether or not the system has reached steady state, i.e., converged to a final solution. At steady state, the value of BF should be approximately equal to unity. It may approach to unity in both directions, and due to semiadaptive control procedure used, it may oscillate around unity. Larger amplitude oscillations are observed at first, but they rapidly damp out.
2. The estimated values of the major parameters, e.g. the average energy and the W-value show a similar fluctuation, and then they converge to nearly constant values as steady state is approached.
3. The validity of the convergence has been checked using test problems where analytic results are available. A good comparison has been obtained, showing that the program and the model are working correctly, free of obscure errors.

## B. Error Estimation and Analysis

### 1. Error Analysis

1. As mentioned in early chapters, Monte Carlo simulation is equivalent to performing mathematical experiments, the results being based on observations of such experiments. From this analogy we can derive the following procedures for the analysis of errors or accuracies of our calculations.

It is convenient to subdivide experimental errors into three broad types, namely: random errors, systematic errors, and blunders. In general, the experimental error is some additive function of all three, while blunders can hopefully be eliminated. We shall describe them in detail in following sections.

- a. Random errors - such errors are of great concern in Monte Carlo type calculations. They generally involve statistical fluctuations or deviations, representing the difference between the singly measured value and the best value of a set of measurements, i.e., the difference between the arithmetic mean and the true mean. Sometimes such errors are referred to as elementary (or inherent) errors in the measurements. The commonly used measure of random errors is the variance or standard deviation.
- b. Systematic error - a systematic error tends to have the same algebraic sign, i.e. it is either an additive or subtractive quantity introduced in the measurement

process. It is an unpleasant and insidious contribution which is not generally amenable to statistical treatment. Thus such errors seriously impair the reliability of the estimation. Typical examples include:

- i) Incorrect assumptions or approximation in the representation of certain processes.
- ii) Constructional faults and mistakes in the algorithms or subroutines.
- iii) Inadequate regard of constancy of experimental conditions or inadequate sampling techniques (biased).

c. Blunders - these are outright mistakes which should be corrected by all means. Possible examples include:

- i) Incorrect logic, misunderstanding of the problems.
- ii) Errors in transcription of data.
- iii) Mistakes in constants used.
- iv) Confusion of units.

In any type of calculation, the systematic error as well as blunders should be removed. One way for detecting such errors is to compare the program against some known reliable solutions. Only the random errors are subject to reduction by the various treatments amenable to attack by variance reduction techniques.

The precision in an estimated (mean) value is proportional to the reciprocal of the statistical error and is

high if the statistical error is small, i.e. the accuracy is high if the net systematic error is small. Usually, but not necessarily, high accuracy implies small statistical errors as well.

As discussed in later sections, our model and program has been checked against several test problems, and good agreement has been obtained. Therefore, it can be assumed that any serious systematic errors have been removed and that the program is free of blunders. Thus, we shall concentrate on the analysis of random errors.

No thorough analysis of variance has been made. Due to the intricate way that the histories calculated during a given time cycle depend on previous time cycles, a meaningful variance estimate beyond the first time cycle seems to be out of the question. However as the system reaches steady state, the deviations or variations (statistical fluctuations) of the estimated parameters can be estimated by means of standard deviation (or variance) over subsequent time cycles. Namely, we use the following formulation:

$$s_i = \left[ \frac{\sum_{j=1}^n (f_{ij} - \bar{f}_i)^2}{n} \right]^{1/2} \quad \forall i \quad (61)$$

where  $s_i$ : the standard deviation at the  $i$ th energy interval over  $n$  consecutive time cycles.

$f_{ij}$ : the parameter value (density distribution function of  $i$ th energy interval, at  $j$ th time cycle.

$\bar{f}_i$ : the average parameter value of  $i$ th interval over  $n$  time cycles.

A heuristic approach was adopted to measure local deviations or statistical fluctuations at each time cycle. This provides a relative magnitude of deviation or dispersion which can be used to compare the various variance reduction techniques. The approach is described as follows:

The estimated major frequency distribution (the electron energy density distribution function in our case) is obtained in the form of histogram. The abscissa of the histogram is divided into so-called class intervals (energy intervals in our calculations). In each such interval an erect rectangle (block) of height  $f_i$  and width  $E_i$  is formed. This block-area type of distribution is the fitted frequency distribution curve. We devise a numerical description of local deviations by defining:

- i) the location index of the center of the erected rectangles in the histogram as  $i$  and with abscissa value  $E_i$ .
- ii) the spread or dispersion around the center.

The total number of particles in each class interval (energy interval) is proportional to the area of the rectangular flock ( $f_i \Delta E_i$ ), the particle's energy in this class interval falls in the range  $(E_i - \Delta E_i/2, E_i + \Delta E_i/2)$  with overall average energy  $\bar{E}_i = E_i$ . Then the following relationship holds (refer to Figure 5):

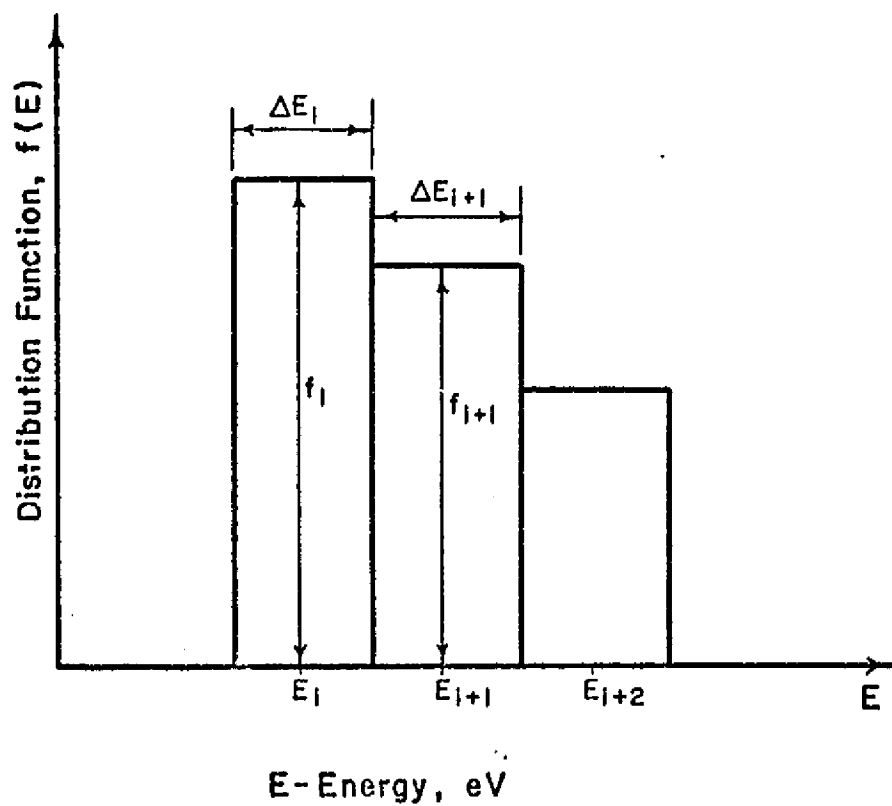


Figure 5. Rectangular block representation of the electron energy distribution histogram.



$$\frac{\Delta \bar{E}_i}{\bar{E}_i} = \frac{\Delta f_i}{f_i}$$

$$\text{or } \Delta f_i = f_i \frac{\Delta \bar{E}_i}{\bar{E}_i} \quad (62)$$

where  $\Delta \bar{E}_i$  is the deviation of the average energy  $\bar{E}_i$  from  $E_i$  or  $\Delta \bar{E}_i = \bar{E}_i - E_i$ . Also  $\Delta f_i$  is the corresponding deviation of the frequency distribution at location index  $i$ .

Equation (62) provides a relative measure of the amplitude of the deviation of density distribution curve at energy  $E_i$  which can be used for comparing the efficiency of different variance reduction techniques as well as provide an heuristic clue to the modification of the fitted density distribution curve.

## 2. Data Filtering

The random errors (statistical fluctuations) result in bumps and ridges in the estimated frequency distribution function. It may be possible to alleviate many of these distortions by means of data smoothing techniques (data filtering).

Two types of data smoothing techniques commonly used in the processing of time series might be used here:

- a. Moving averages - this technique operates by replacing each point of the frequency distribution function (height of the histogram at  $E_i V_i$ ) at time cycle  $t_j$  with an average value of several subsequent points in the time cycle series. Thus if  $f_{ij}$  is the value

of frequency function at  $i$ th energy interval and  $j$ th time cycle, then  $f_{ij}$  is replaced by

$$\sum_{j=1}^M f_{ij} \cdot W_j / M$$

where  $W_j$ , the weighting factor for the  $j$ th time cycle, usually has the value of 1 or we can define it as a function of  $\Delta t_j$ . The value of  $M$  is the number of successive points of the time cycles to be included in each average.

- b. Parabolic filtering - this is similar to the moving averages technique, but  $M$  is selected as an odd number, the point to be replaced is located in the center of these  $M$  points, and a parabola is fitted to these points by means of least squares. Thus it is replaced by the corresponding point  $f_i$  on the derived parabola. The process is repeated by shifting the center point to the next one in the time cycle series.

The technique of moving averages (a) is employed in the program, because of its simplicity and the fact that the parabolic smoothing technique is too sensitive to the changes in the distribution function. Note that this smoothing (filtering) technique only becomes effective as the simulated system reaches steady state.

## C. Solutions of the Problem

### 1. Preliminary Checks

In order to verify that the model is valid and correctly working, and the computer program is free of bugs, blunders, and systematic errors, we solved a specific problem first reported by J. A. Smit (11) and later extended by Heylen and Lewis (10). Their solutions for the electron energy distribution were obtained analytically based on the Boltzmann transport equation for the special case characterized as follows:

- i) electric field present
- ii) no external sources
- iii) infinite size medium of background noble gases.

The comparisons are made for helium gas at  $E_{f/p} = 4, 5, 10 \text{ V/cm}^{-1}$ . ( $E_{f/p}$ : electric field strength to pressure ratio). The results are shown in Figures 6, 7 and 8. Both Smit and Heylen's results have taken into account all the collision processes including those due to elastic loss as well as inelastic losses. Heylen's results is more up-to-date, for low  $E_{f/p}$  values only  $E_{f/p} = 5$  is given, which was chosen for the present comparison. In so far as possible, the same cross sections as Smit and Heylen and Lewis used were incorporated in the present calculations. The slight discrepancies apparant at the low-energy end is probably due to differences in treating the scattering process, and

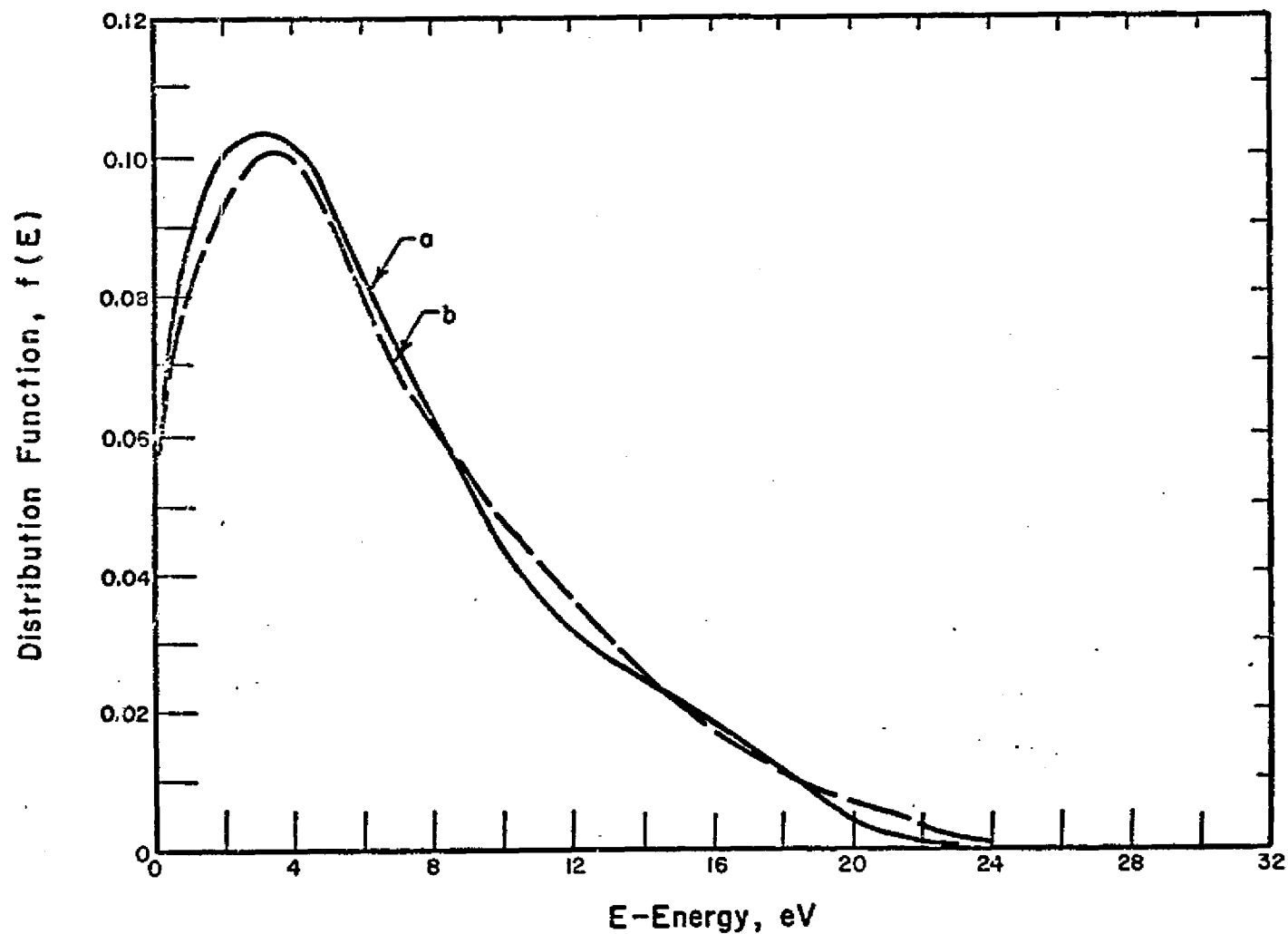


Figure 6. Comparison of the electron energy distribution for He at  $E_f/p = 4$  for  
 a) Smit's calculation (1936) ( $\bar{E} = 6.4$ )  
 b) present calculation ( $\bar{E} = 6.51$ )

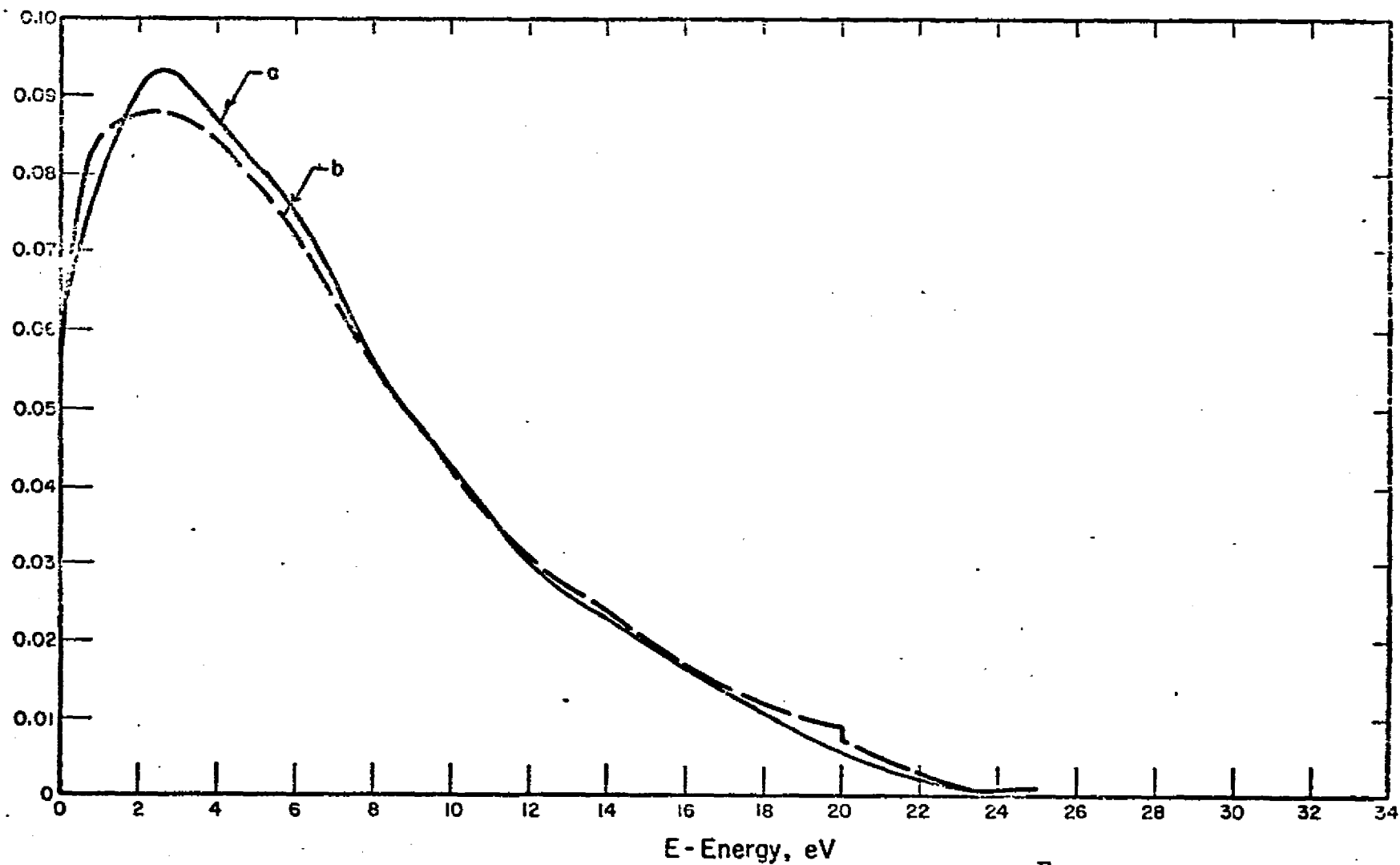


Figure 7. Comparison of electron energy distribution for He at  $E_f/p = 5$   
 a) present calculation ( $\bar{E} = 6.64$ )  
 b) Heylen and Lewis' result ( $\bar{E} = 6.51$ )

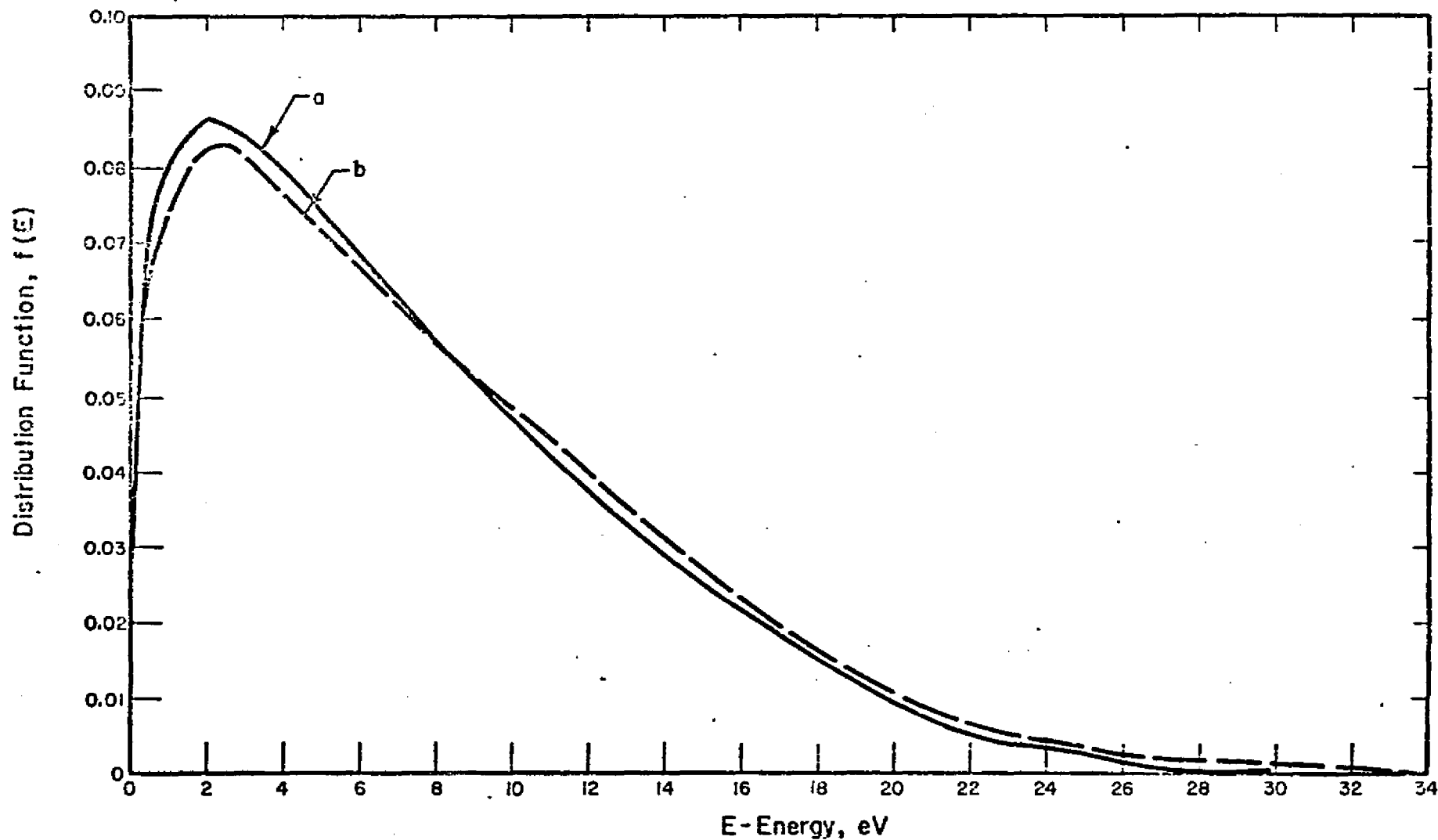


Figure 8. Comparison of electron energy distribution for He at  $E_f/p = 10$   
 a) Smit's result ( $\bar{E} = 7.7$ )  
 b) present calculation ( $\bar{E} = 7.83$ )

the many approximations required in their analytic solutions. Also, some of the cross sections used in this calculations are experimental values, which may differ by some amount from that used for analytic solutions.

For the high energy region, (near the source energy) checks have been made against the analytic solutions reported by Lo and Miley (32). As shown in Figure 9, the general shape of the electron energy flux distributions are in reasonable agreement. In this case it was not practical to use the same cross sections as incorporated in the analytic solutions, and this may account for some of the differences observed. Also the analytic solutions are obtained for an infinite medium vs. a 1 cm slab in the present case. Thus the analytic solutions do not allow leakage, although this becomes important in the present calculation for low pressures. Lo and Miley's results are plotted in Figure 9 for the source energy  $E_0 = 1500$  eV and  $E_0 = 500$  eV, whereas the present results are for  $E_0 = 1000$  eV. The plot shown is for the flux density (normalized) instead of energy number density. They are related as follows:

$$\varphi(E) = f(E) \cdot V(E) \quad (63)$$

where  $\varphi(E)$ : electron flux density  
 $f(E)$ : electron number density  
 $V(E)$ : Electron kinetic velocity

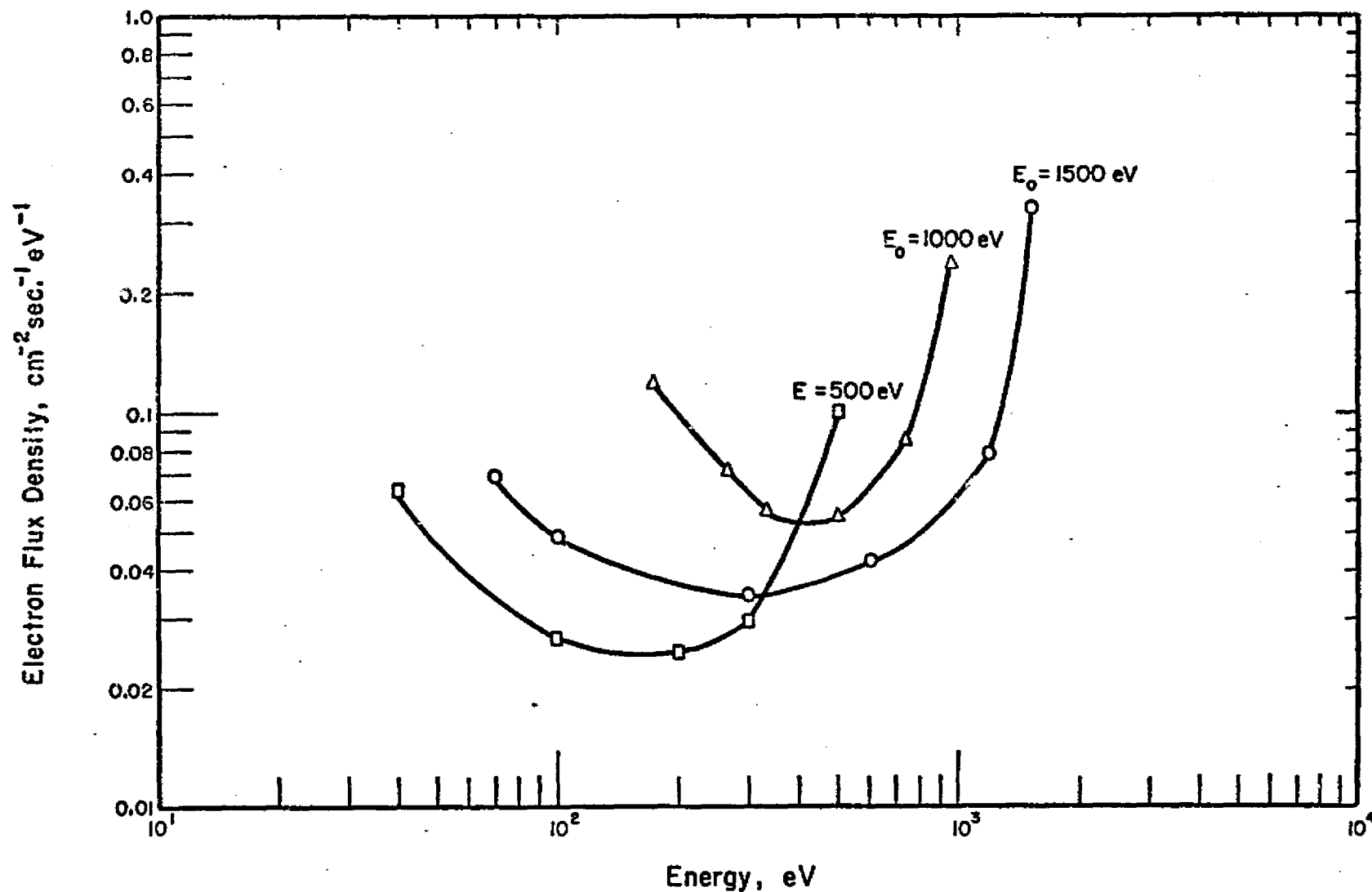


Figure 9. Comparison of the high-energy portion of flux density curve  
 a)  $E_0 = 1500 \text{ eV}$  and  $E_0 = 500 \text{ eV}$ : Lo and Miley's result  
 b)  $E_0 = 1000 \text{ eV}$ : present calculations



Since we only intend to compare general shapes of distribution, the normalizations used for the various curve, shown here are arbitrarily chosen for optimum display of information.

## 2. The Electron Energy Distribution Function without Electric Field

The calculations for electron energy distribution functions for helium with a high energy electron source, but no applied electric field (Figures 10 to 14) were carried out at two different pressures and various source rates. The slab thickness  $D$  is chosen to be fixed at  $D = 1$  cm, which is equal to the actual diameter of cylindrical tube used in the laboratory (7). The two pressures are chosen to be  $P = 10$  torr and  $P = 760$  torr, corresponding to a typical low-pressure case and a normal atmospheric pressure.

We have chosen two different source energies ( $E_0$ ), one at  $E_0 = 1,000$  eV and the other at  $E_0 = 70$  eV. These values roughly correspond to the highest energy and the average energy in the primary electron energy spectrum for MeV alpha-particle irradiation of helium as done in the experimental studies conducted at the University of Illinois (6).

The range of interest for the source rates for a boron coated tube irradiated in the Illinois TRIGA reactor lies between  $10^{14}$  to  $10^{18}$  ( $\#/\text{cm}^3\text{-sec}$ ). However, results shown for source rates above about  $10^{16}$  are questionable since the resulting ionization density becomes large enough that

electron-electron collisions become important, and this process was neglected here.

The major quantity of interest here is the electron energy distribution function which were defined as  $n^-(E)$  in Chapter II. For simplicity, we represent the normalized distribution corresponding to  $n^-(E)$  by  $f(E)$ . This normalization is defined by:

$$\int_0^{E_0} f(E) dE = 1 \quad (64)$$

For the plasma system without electric field ( $E_f = 0$ ), the distribution functions with source energy  $E_0 = 1000$  eV are shown in Figures 10 and 11 for the pressure  $p = 10$  torr, and  $p = 760$  torr respectively. Similar distribution functions for  $E_0 = 70$  eV are shown in Figures 12 and 13.

An idealized Maxwellian distribution plus  $1/E$  tail are shown in Figures 10 to 13 for reference. (The Maxwellian plus  $1/E$  distribution was used as the initial distribution for these calculations. Note that since this curve is for reference only, it is arbitrarily normalized to a point on the Maxwellian curve and does not satisfy Equation 64.) For the range of source rates mentioned, the calculated energy distribution for helium (He) are reasonably Maxwellian at low energy region. The differences due to variations in the source rates show up at the high energy end, especially near the source energy, where the magnitudes seem to depend linearly on the values of source rate. This is somewhat

deceptive, however, since the absolute magnitude at any point is due to the normalization. (This point is stressed again later in connection with Figure 24).

The fact that the distributions drop off rapidly for energies above about 20 eV is attributed to the fact that the threshold energy for inelastic processes is 20 eV, above which energy the ionization and excitation processes bring more electrons into the low-energy region.

In order to investigate the change in distribution function at very high source rates, a calculation for  $S_0 = 10^{22}$  particles/cm<sup>3</sup>-sec was carried out. As mentioned earlier, the neglect of electron-electron collisions makes the accuracy of this calculation questionable. Still the trends are of interest, and it is seen from Figure 14 that the distribution curve changes drastically. In fact, it is no longer Maxwellian at the low-energy end. Due to the high source rate, more high-energy particles and more particles with energy in the range from 0.3 eV to 20 eV are present in the system.

As discussed earlier, the distribution curve drops sharply for energies above the excitation threshold energy (at  $E = 20$  eV) due to the dominance of ionization and excitation collision processes. At the higher pressure, the distribution curves decrease more rapidly compared to low pressure, since increased scattering slows the electrons down more rapidly. This effect is shown in Figures 15 and 16,

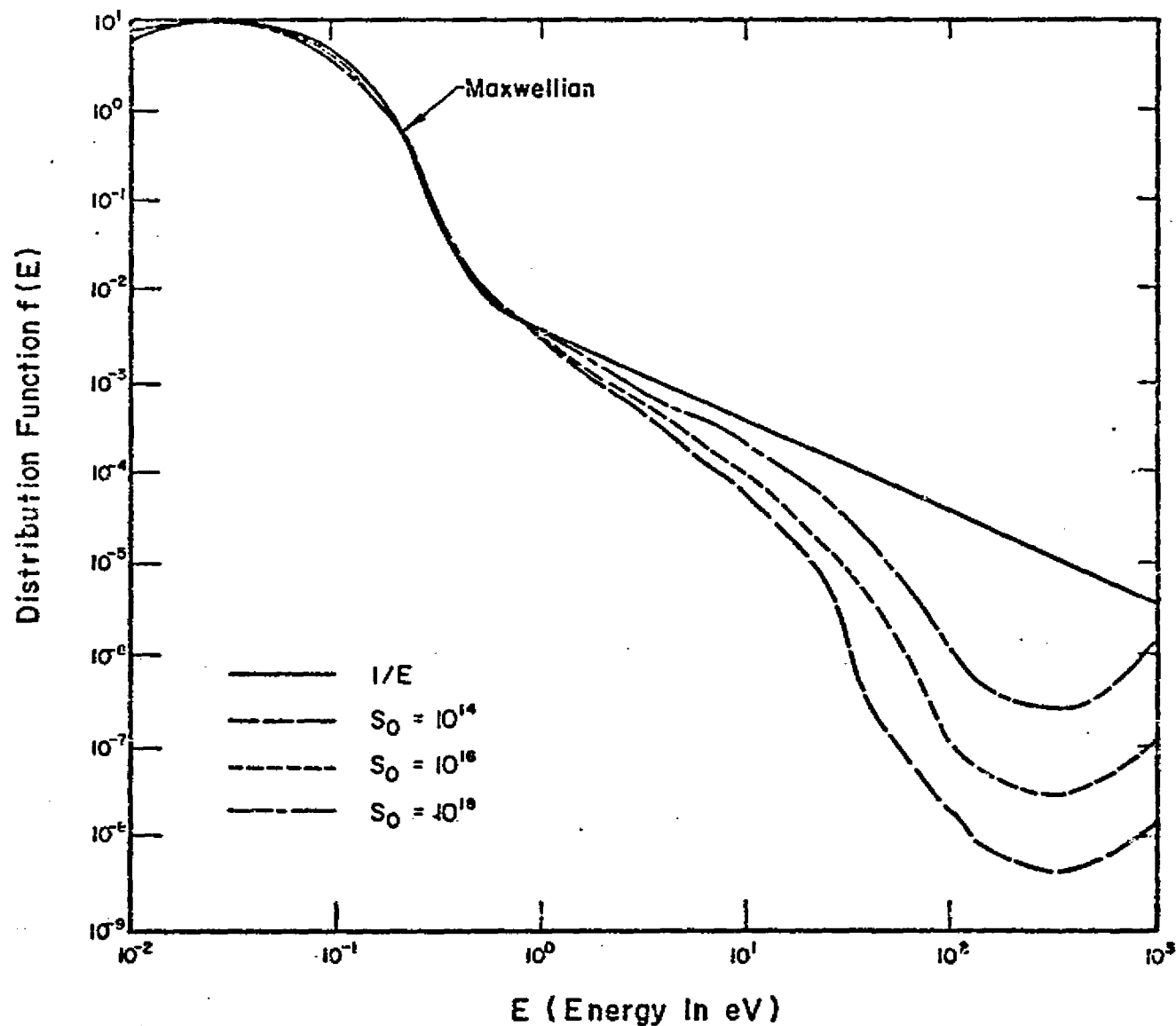


Figure 10. The electron energy distribution for He without an electric field but different source rates at  $p = 10$  torr,  $E_0 = 1000$  eV and Maxwellian plus  $1/E$  distribution.

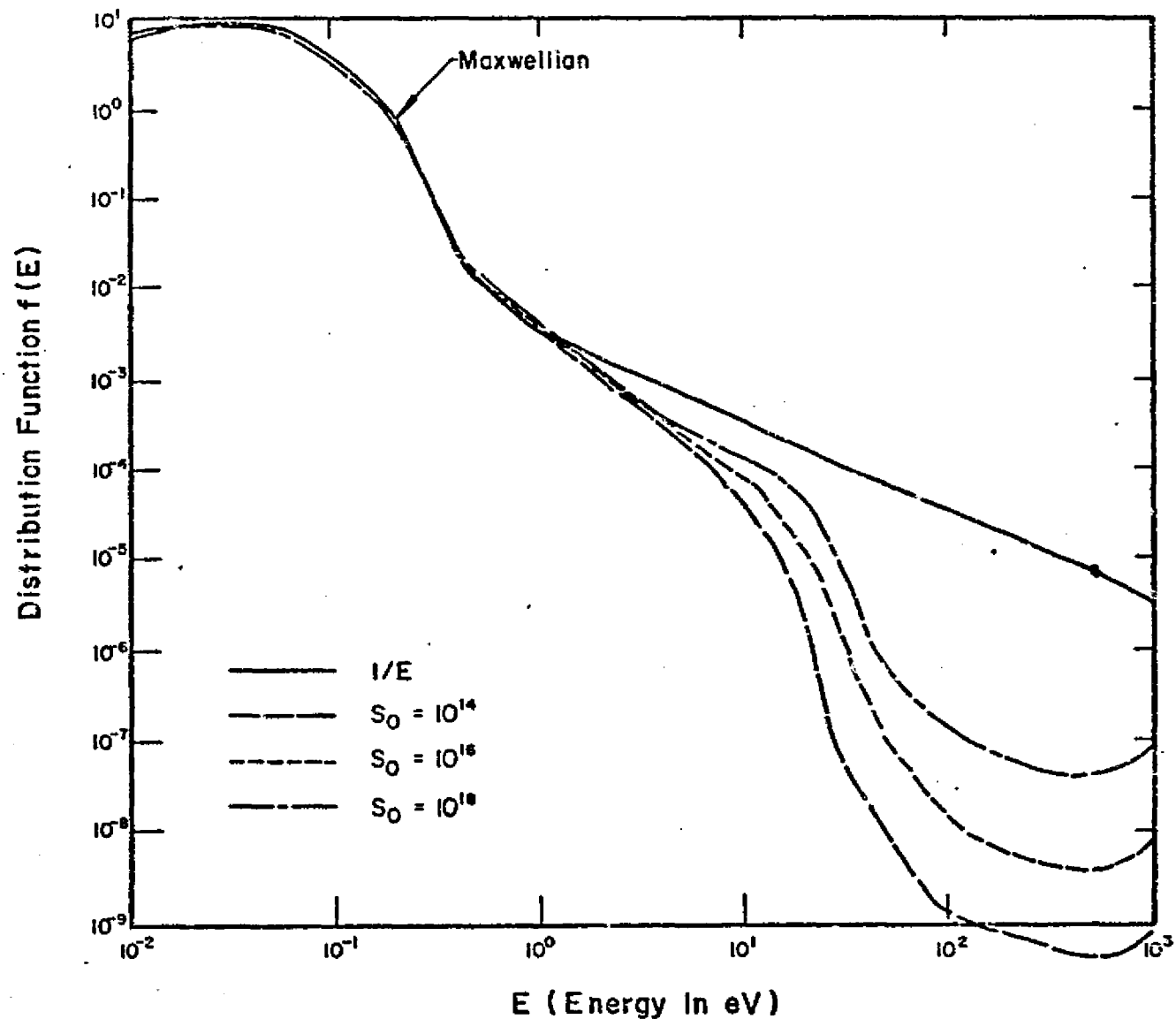


Figure 11. The electron energy distribution for He without an electric field but different source rates at  $p = 760$  torr,  $E_0 = 1000$  eV and Maxwellian plus  $1/E$  distribution.

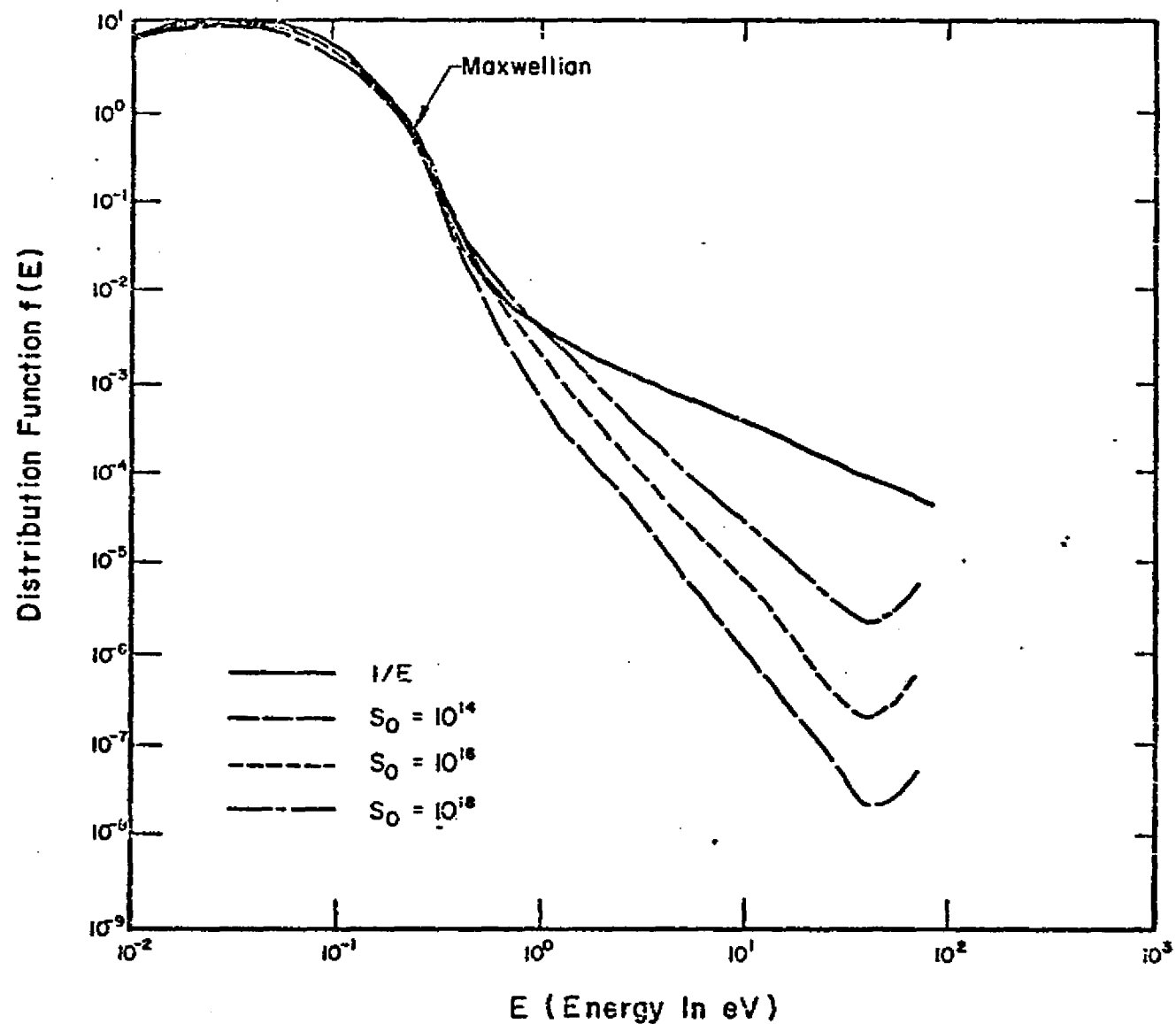


Figure 12. The electron energy distribution for He without an electric field but different source rates at  $p = 10$  torr,  $E_0 = 70$  eV and Maxwellian plus  $1/E$  distribution.

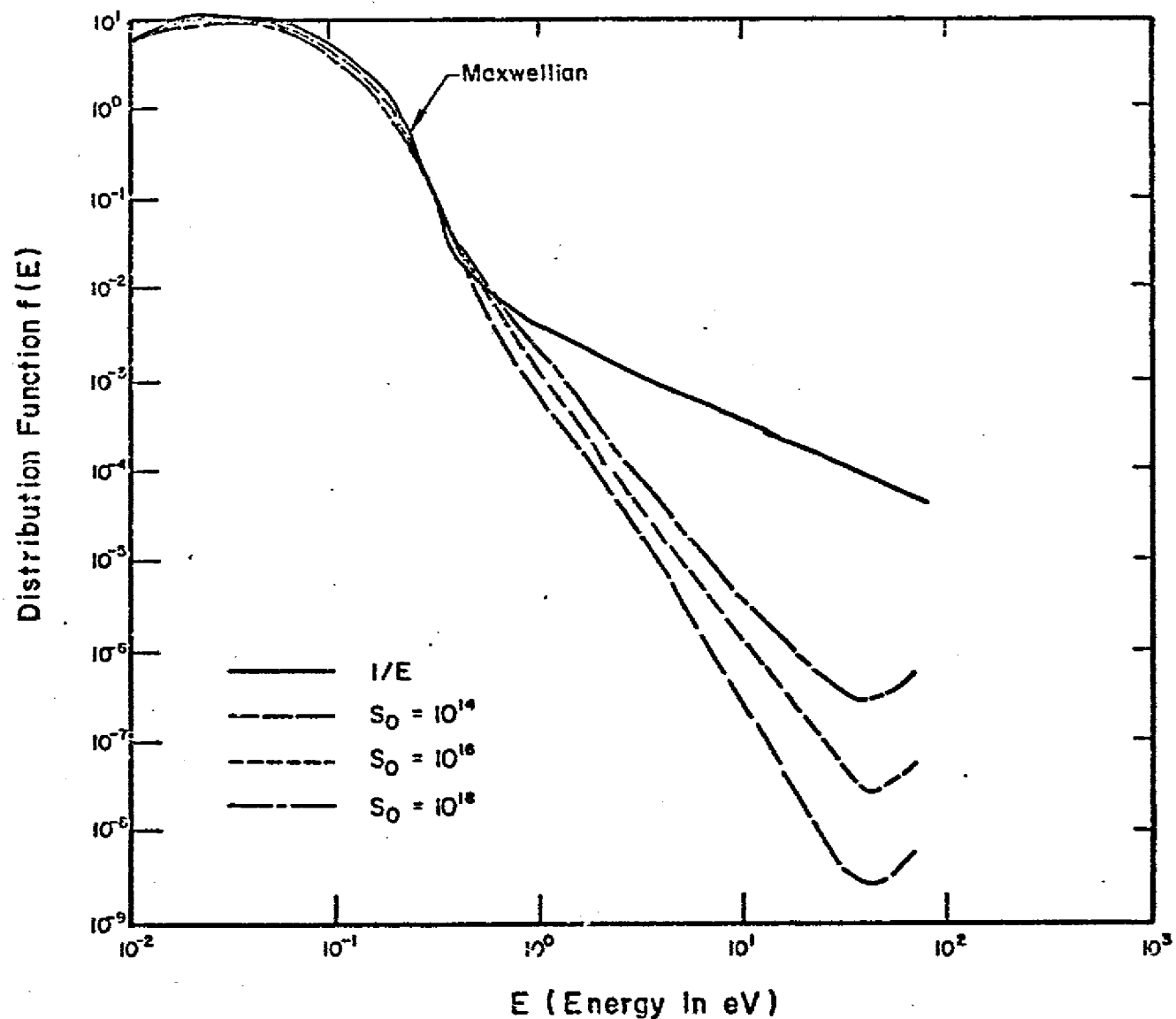


Figure 13. The electron energy distribution for He without an electric field but different source rates at  $p = 760$  torr,  $E_0 = 70$  eV and Maxwellian plus  $1/E$  distribution.

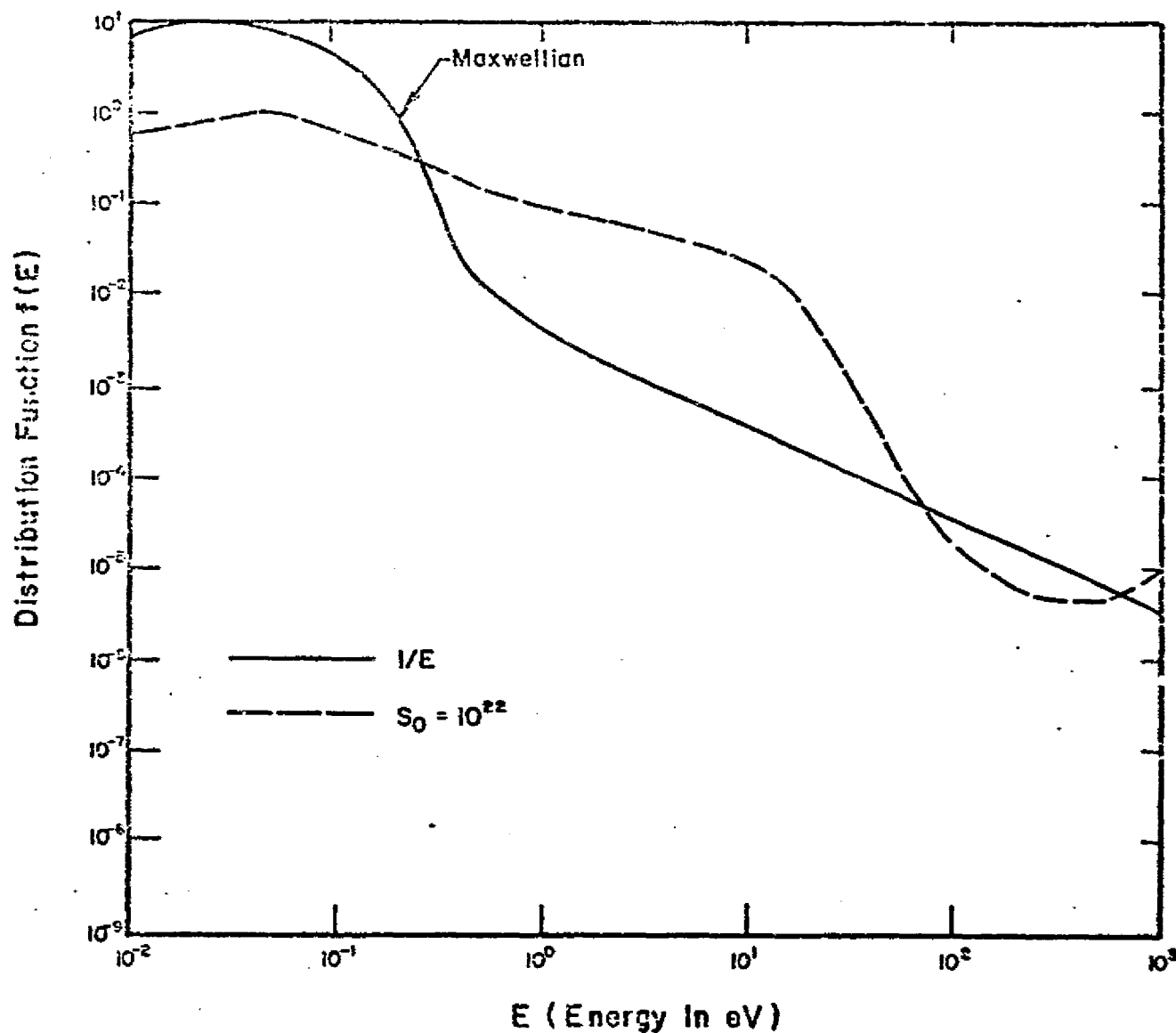


Figure 14. Comparison of  $f(E)$  for He without an electric field (source rate  $S_0 = 10^{22}$  #/cm<sup>3</sup>-sec) with the standard Maxwellian distribution plus  $1/E$  tail.



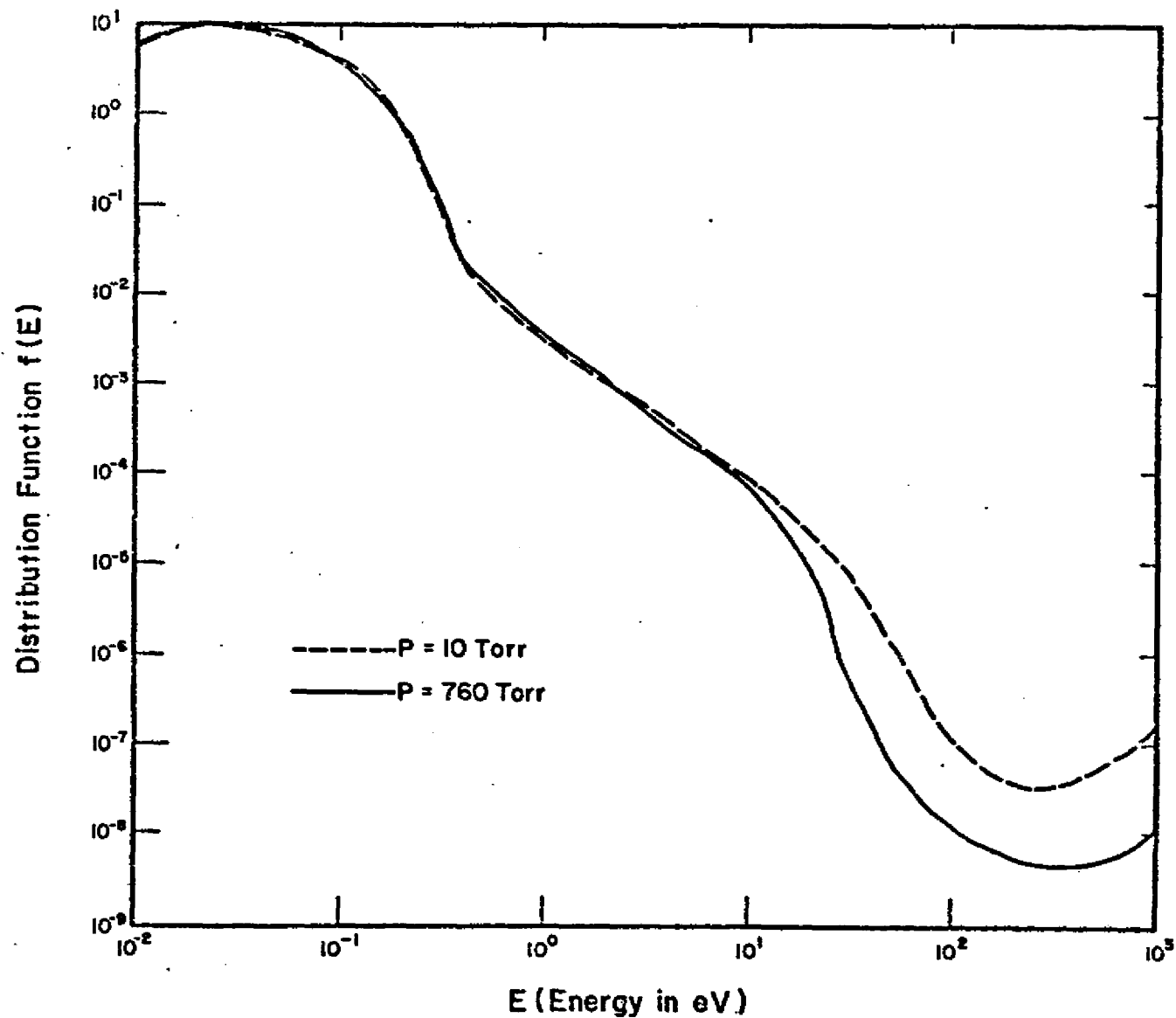


Figure 15. The electron energy distribution for He, without an electric field, for different pressure at  $S_0 = 10^{16}$ ,  $E_0 = 1000$  eV.

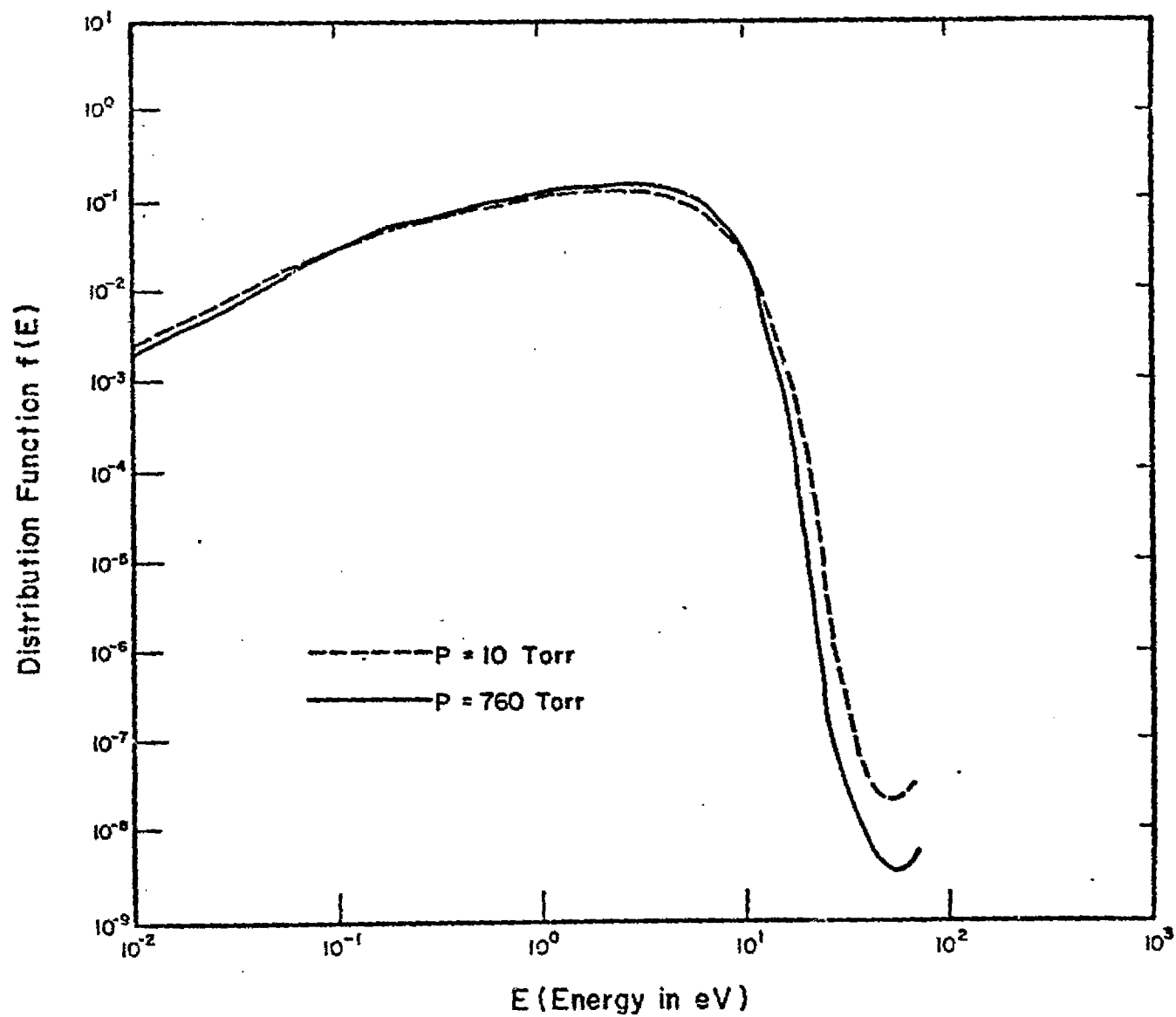


Figure 16. The electron energy distribution for He, without an electric field, for different pressure at  $S_0 = 10^{16}$ ,  $E_0 = 70$  eV.

for the same source rate  $S_0 = 10^{16}$  (#/cm<sup>3</sup>-sec) and source energy, but two different pressures. The reduction in magnitude of the high energy tail as the pressure increases is clearly illustrated.

In these calculations, the rate of both leakage and absorption (recombination) are strongly dependent on the pressure. At higher pressures (e.g.,  $p = 760$  torr), very few particles leak out of the system, and the loss are dominated by the absorption process. For lower pressure ( $p = 10$  torr), fewer particles are absorbed, while more escape or leak out of the system. Also, more particles escape in the high-energy region, but the absorbed particles are mainly in the low-energy region.

The estimated major parameters are summarized in Table 1, which includes the average energy of the system ( $\bar{E}$ ), the number of positive ions ( $N^+$ ) and the W-values for various combinations of source rate  $S_0$ , source energy  $E_0$  and pressure  $p$ .

The W-value is defined as the amount of energy required (input) to produce an ionization pair (positive ion and electron), or

$$W\text{-value} = \frac{(N_s \cdot E_0 - E_{Tes} - E_{Tab})}{N_z} \quad (65)$$

Where  $N_s$ : the number of source particles introduced into the system during the period  $\Delta t$ .

$E_{Tes}$ : the total energy lost with escaping particles during the time  $\Delta t$ .

Table 1. Calculated Parameter Values for the Plasma  
Without a Electric Field

* P (torr)	$E_o$ (eV)	$S_o$ (#/cm <sup>3</sup> -sec)	$\bar{E}$ (eV)	$N^+$ (#)	W-value	Range
10	1000	$10^{18}$	0.171	$0.125 \times 10^{14}$	41 ~	63
10	1000	$10^{16}$	0.157	$0.125 \times 10^{13}$	35 ~	53
10	1000	$10^{14}$	0.111	$0.125 \times 10^{12}$	49 ~	58
760	1000	$10^{18}$	0.167	$0.583 \times 10^{15}$	32 ~	38
760	1000	$10^{16}$	0.123	$0.446 \times 10^{14}$	40 ~	65
760	1000	$10^{14}$	0.104	$0.67 \times 10^{13}$	35 ~	54
10	70	$10^{18}$	0.104	$0.125 \times 10^{14}$	50 ~	62
10	70	$10^{16}$	0.082	$0.125 \times 10^{13}$	47 ~	63
10	70	$10^{14}$	0.075	$0.124 \times 10^{12}$	40 ~	57
760	70	$10^{18}$	0.082	$0.16 \times 10^{15}$	47 ~	65
760	70	$10^{16}$	0.075	$0.141 \times 10^{14}$	42 ~	58
760	70	$10^{14}$	0.075	$0.157 \times 10^{13}$	39 ~	55
Maxwellian distribution						
		0	0.05	---	---	

\*P: pressure,  $E_o$ : source energy,  $S_o$ : source rate,  
 $\bar{E}$ : average energy of the particles,  $N^+$ : number of positive  
ions.

$E_{\text{Tab}}$ : the total energy lost with recombining particles during the time  $\Delta t$

$N_z$ : the total number of ionizations produced during the time period  $\Delta t$ .

As seen from Table 1, the average energy  $\bar{E}$  tends to be higher for higher source rates and higher source energy, and it is lower for higher pressures. However, the differences are in general very small. (All the average energies are consistently greater than that of a Maxwellian distribution). The values for ion density  $N^+$  are roughly proportional to the square root of source rate, and are higher for higher pressures, virtually independent of source energy.

The W-values are quite sensitive to statistical fluctuations, and the ranges of variation obtained over the last 10 time cycles in a given calculation are shown in Table 1. The W-values obtained here do not show any clear dependence on the system parameters, thus variation indicated may be totally due to the statistical fluctuations. This is not too surprising since the W-value is known to be roughly independent of particle energy, etc. The W-value for helium obtained experimentally by Jesse and Sadouski (34) is 43. A reasonable value to choose for the W-value from Table 1 would be 50, which is little bit higher than the experimental result. The reason for this discrepancy is not clear, but it may be due to small inaccuracies in the ionization-excitation cross section employed here. The W-value is, by definition, quite sensitive to the details

of the ratio of the ionization to the excitation cross section as a function of energy.

### 3. The Electron Energy Distribution Function with an Applied Electric Field

The calculated electron energy distribution functions in He with an applied electric field ( $E_f/p = 10$ ) are plotted in Figures 17 to 21, in a manner similar to that in the previous section. All these calculations are carried out at the electric field to pressure ratio  $E_f/p = 10$ . Smit's result (without source or  $S_0 = 0$ ) are plotted together with present calculations at given pressure and source energy ( $p = 10, 760$  and  $E_0 = 1000$  eV, 70 eV), with three different source rates. It is observed that the distribution functions are highly non-Maxwellian. For the source rate in the range  $10^{14}$  #/cm<sup>3</sup>-sec to  $10^{18}$  #/cm<sup>3</sup>-sec, the distribution curves at low energy end are quite close to the Smit's results (11). At the high energy end, they have a shape similar to the curves of no electric fields. Again differences due to the source rate show up at high-energy region. Below 10 eV, all the distribution functions fall on the top of each other, and they begin to spread out as the energy increases above 10 eV. For high pressures or low source energy, more difference (between  $S_0 = 0$  and  $S_0 \neq 0$ ) show up at high-energy end than that of low pressure and/or high source energy. Again at very high source rates  $S_0 = 10^{22}$  #/cm<sup>3</sup>-sec) the distribution function changes appreciably.

In Figure 21, the comparison is made with a very high source rate ( $S_o = 10^{22}/\text{cm}^3\text{-sec}$ ) and with one in the region of interest ( $S_o = 10^{16}$ ,  $p = 760$  torr). For very high source rate, the distribution function changes quite a lot; it still preserves the general shape, but at low-energy end, it is much smaller in amplitude. (However, as stressed earlier, the high source rate calculation is questionable due to the neglect of electron-electron interactions.)

The changes in these distribution functions due to different pressures are also shown in Figures 22 and 23. The observed trends are similar to those in the previous section for the electric field  $E_f = 0$  case.

So far, the normalized distribution functions  $f(E)$  have been displayed. However, to stress the fact that the absolute magnitudes of the curves depends strongly on the source (shapes have received our interest to this point), Figure 24 shows the unnormalized electron energy distribution functions for He (with electric field or  $E_f \neq 0$ ) are plotted, with  $E_f/p = 10$ ,  $p = 10$  torr,  $E_o = 1000$  eV, and three different source rates ( $S_o = 10^{14}$ ,  $S_o = 10^{16}$ ,  $S_o = 10^{18}$ ). The magnitude (unnormalized) distribution is defined as

$$N^-(E) = N^+ \cdot f(E) \quad (66)$$

where  $N^+$ : number of positive ions

$f(E)$ : normalized distribution

The values of  $N^+$  are  $0.24 \times 10^{15}$ ,  $0.24 \times 10^{14}$ , and  $0.24 \times 10^{13}$  corresponding to the three source rates  $S_o = 10^{18}$ ,  $S_o = 10^{16}$

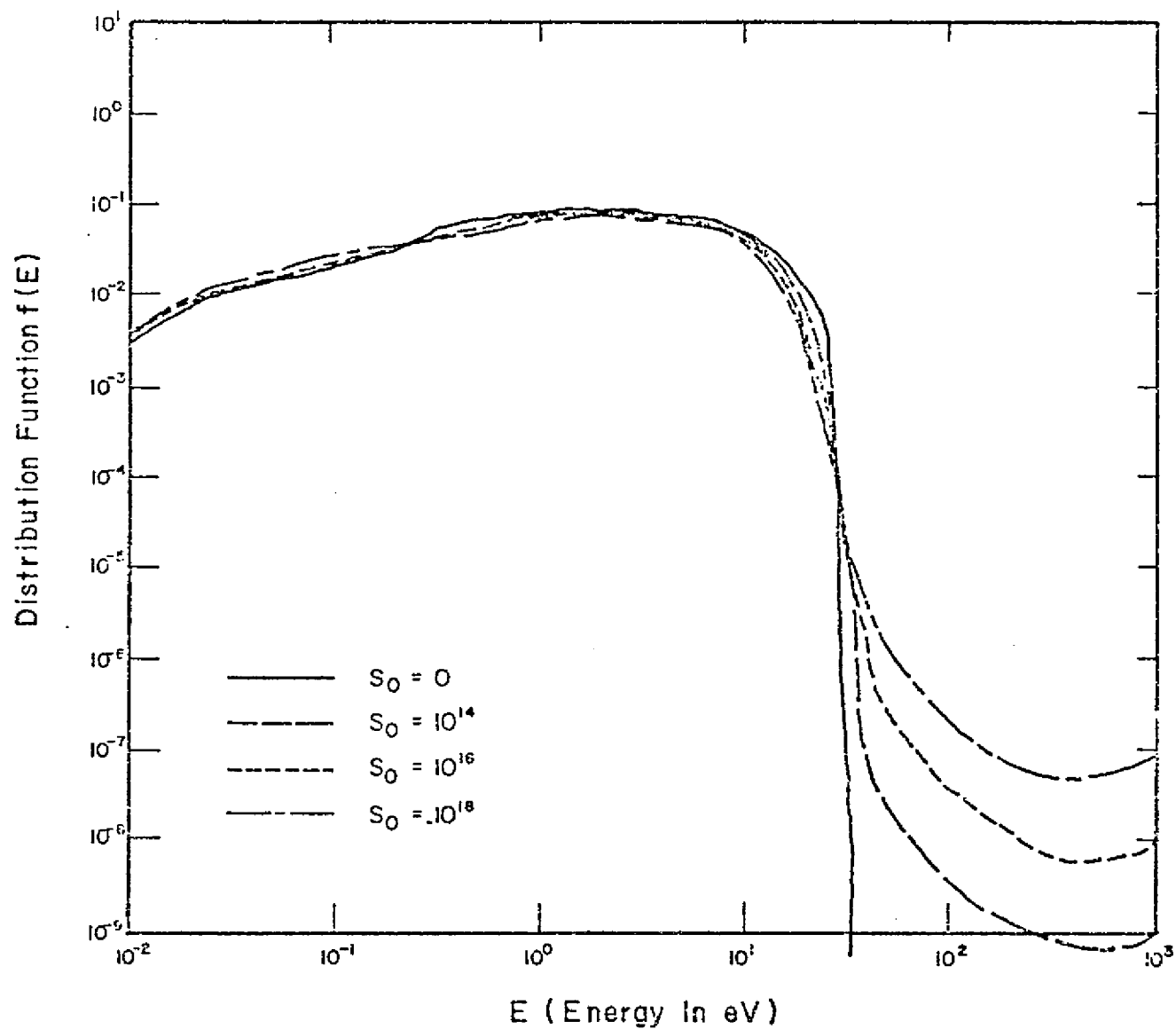


Figure 17. The electron energy distribution for He with an electric field at  $E_f/p = 10$ ,  $p = 10$  torr,  $E_0 = 1000$  eV.



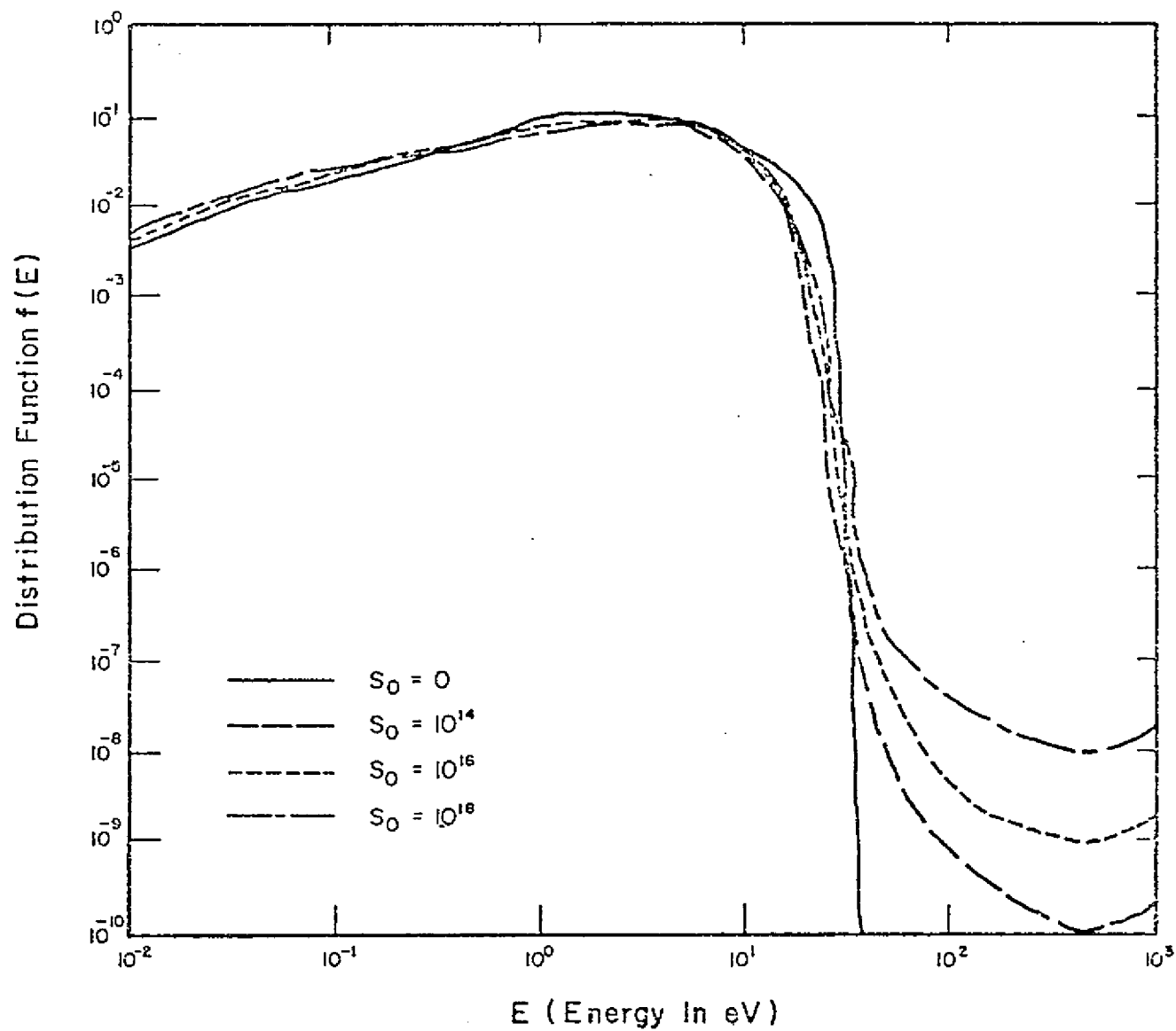


Figure 18. The electron energy distribution for He with an electric field at  $E_f/p = 10$ ,  $p = 760$  torr,  $E_0 = 1000$  eV.

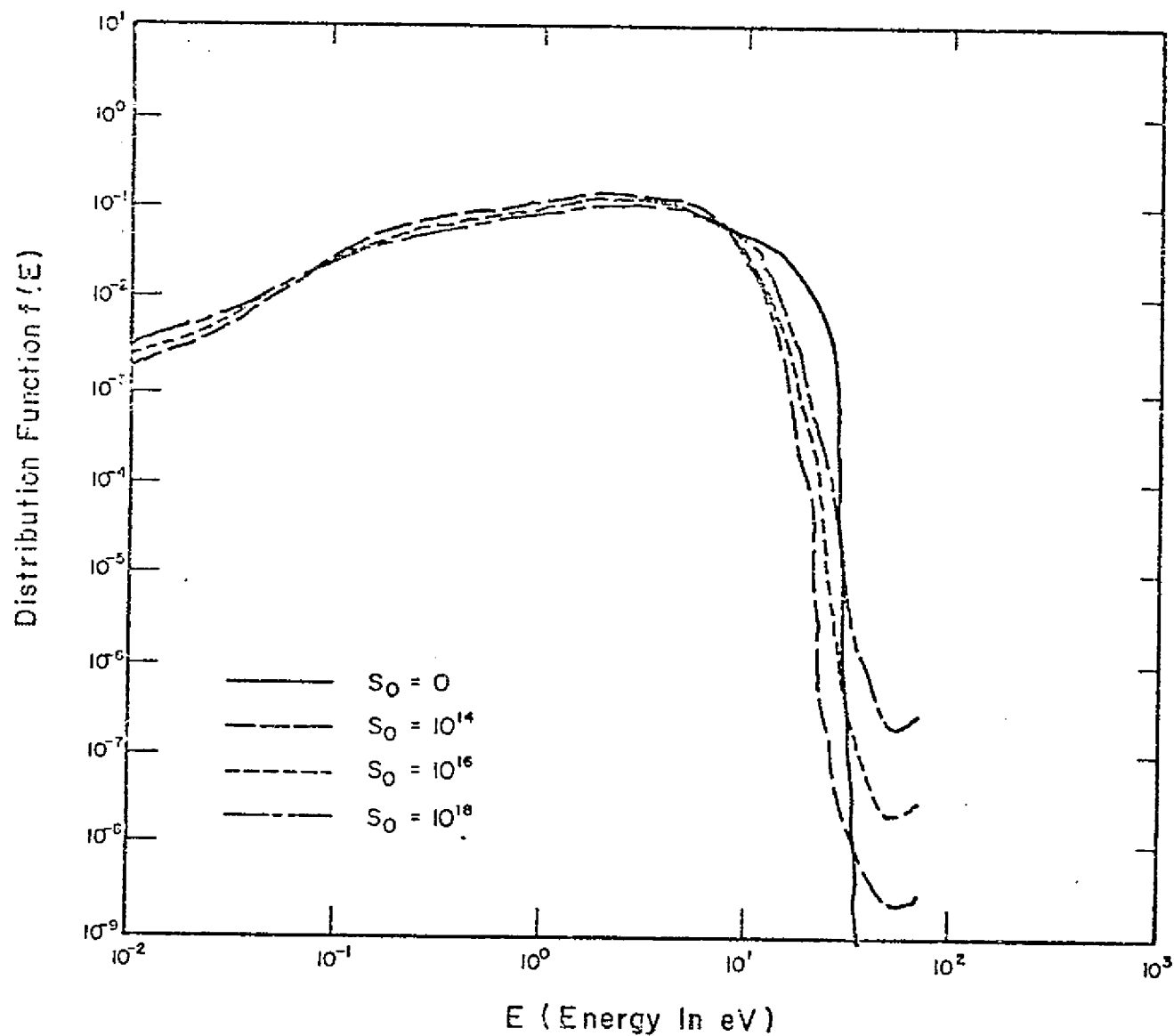


Figure 19. The electron energy distribution for He with an electric field at  $E_{f/p} = 10$ ,  $p = 10$ ,  $E_0 = 70$  eV.

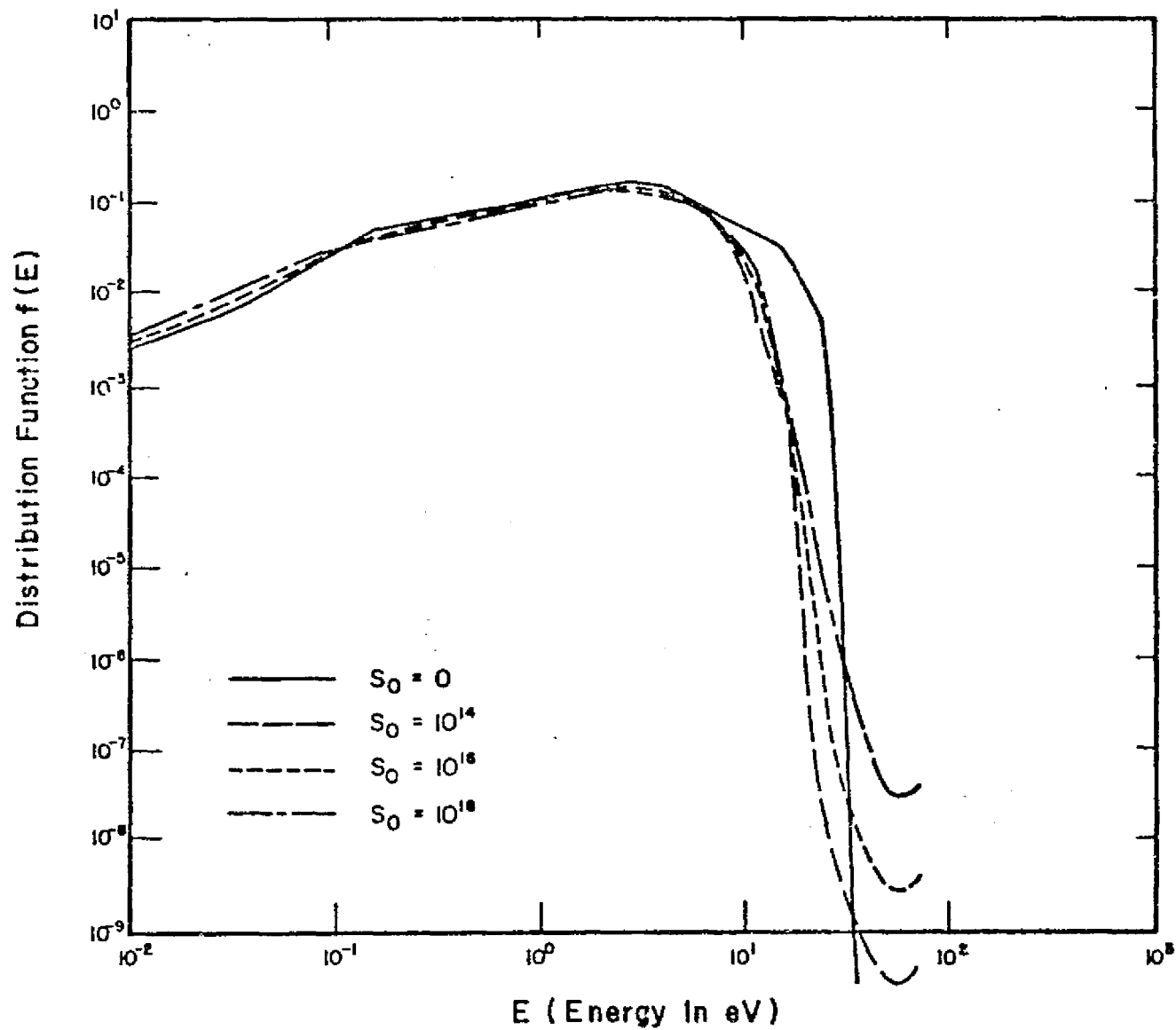


Figure 20. The electron energy distribution for He with an electric field at  $E_f/p = 10$ ,  $p = 760$  torr,  $E_0 = 70$  eV.

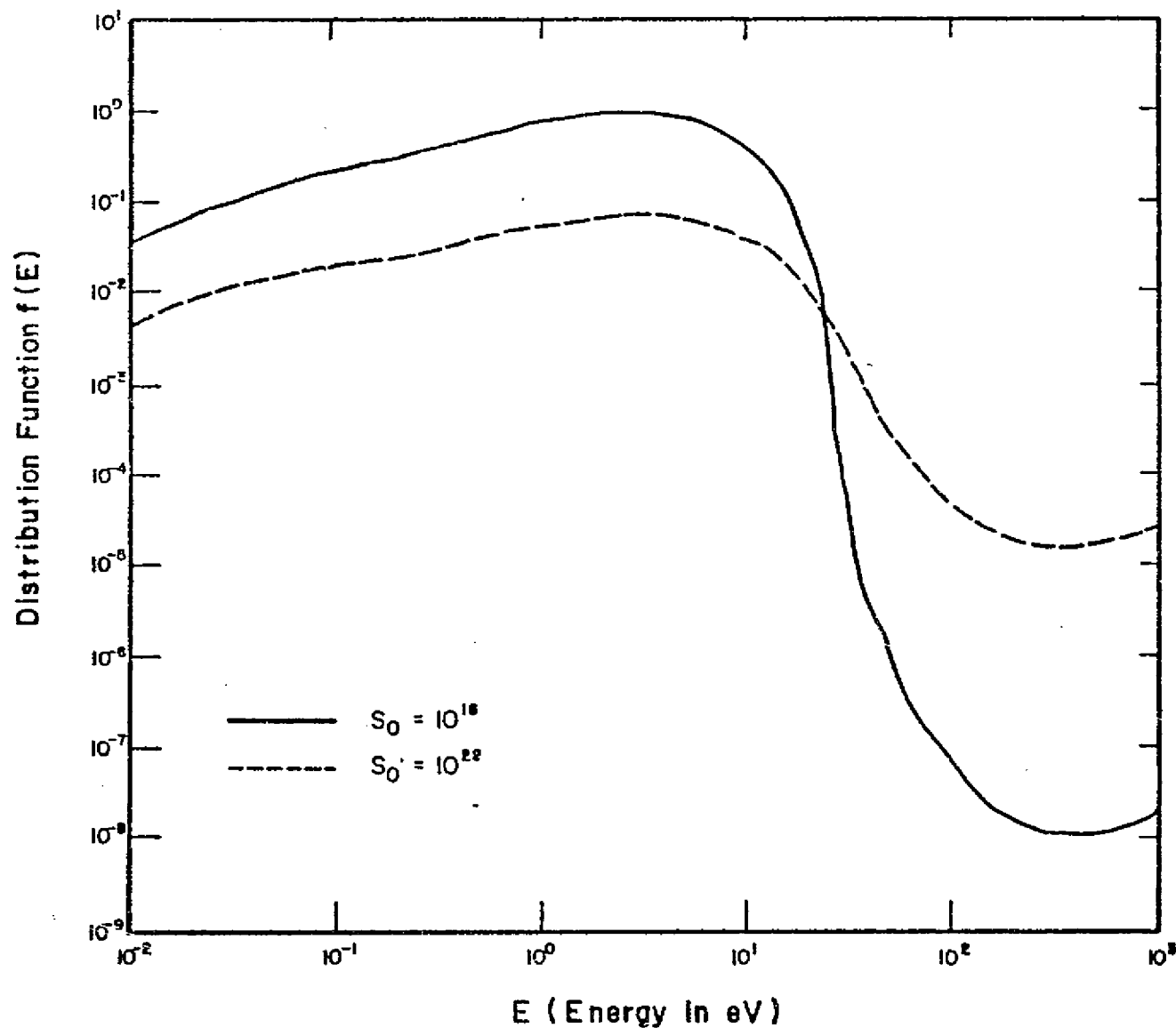


Figure 21. The electron energy distribution function at  $E_f/p = 10$ ,  $p = 760$  torr for  $S_0 = 10^{22}$  and  $S_0 = 10^{16}$  ( $\#/\text{cm}^3\text{-sec}$ ).

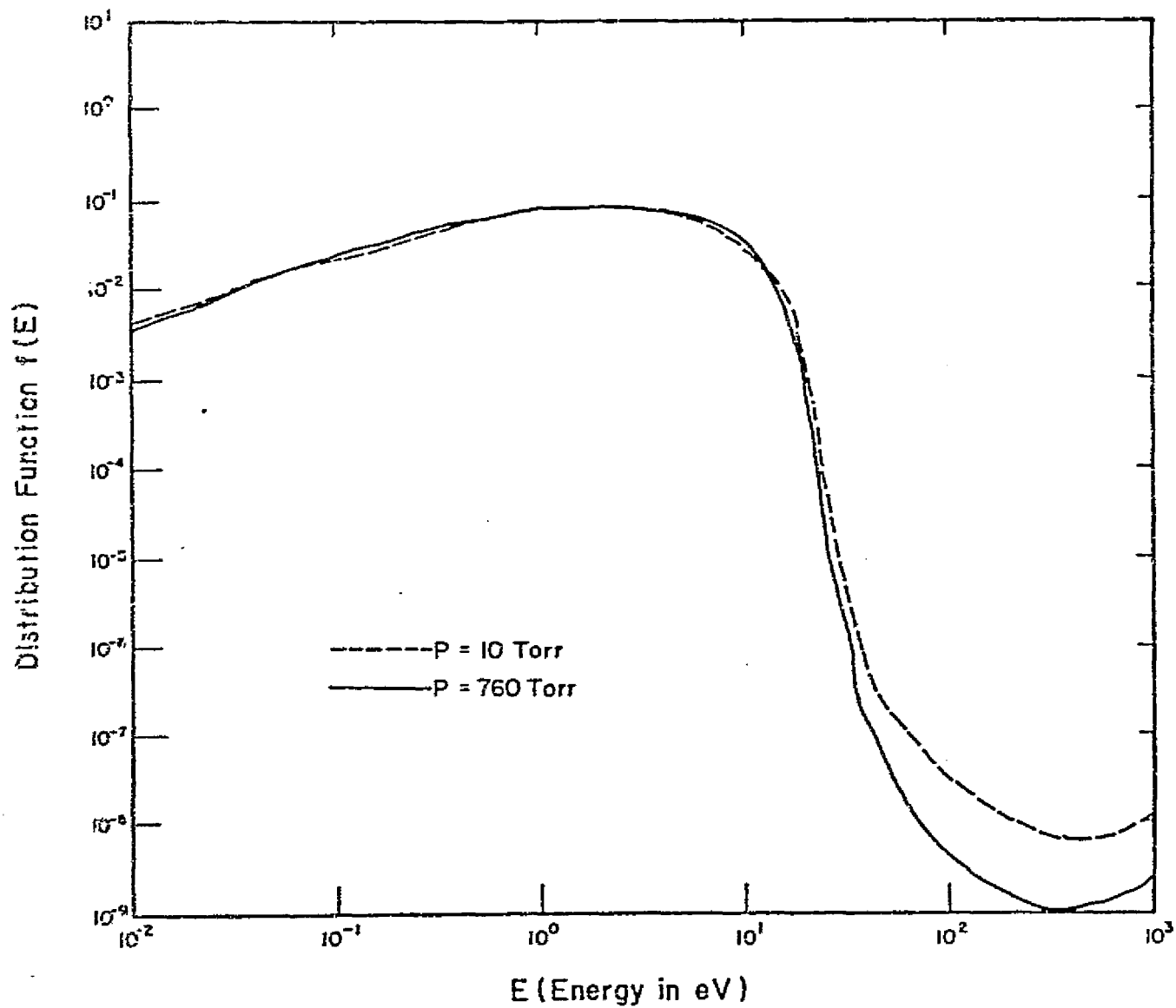


Figure 22. The electron energy distribution for He, with an electric field at  $S_0 = 10^{16}$ ,  $E_0 = 1000$  eV.

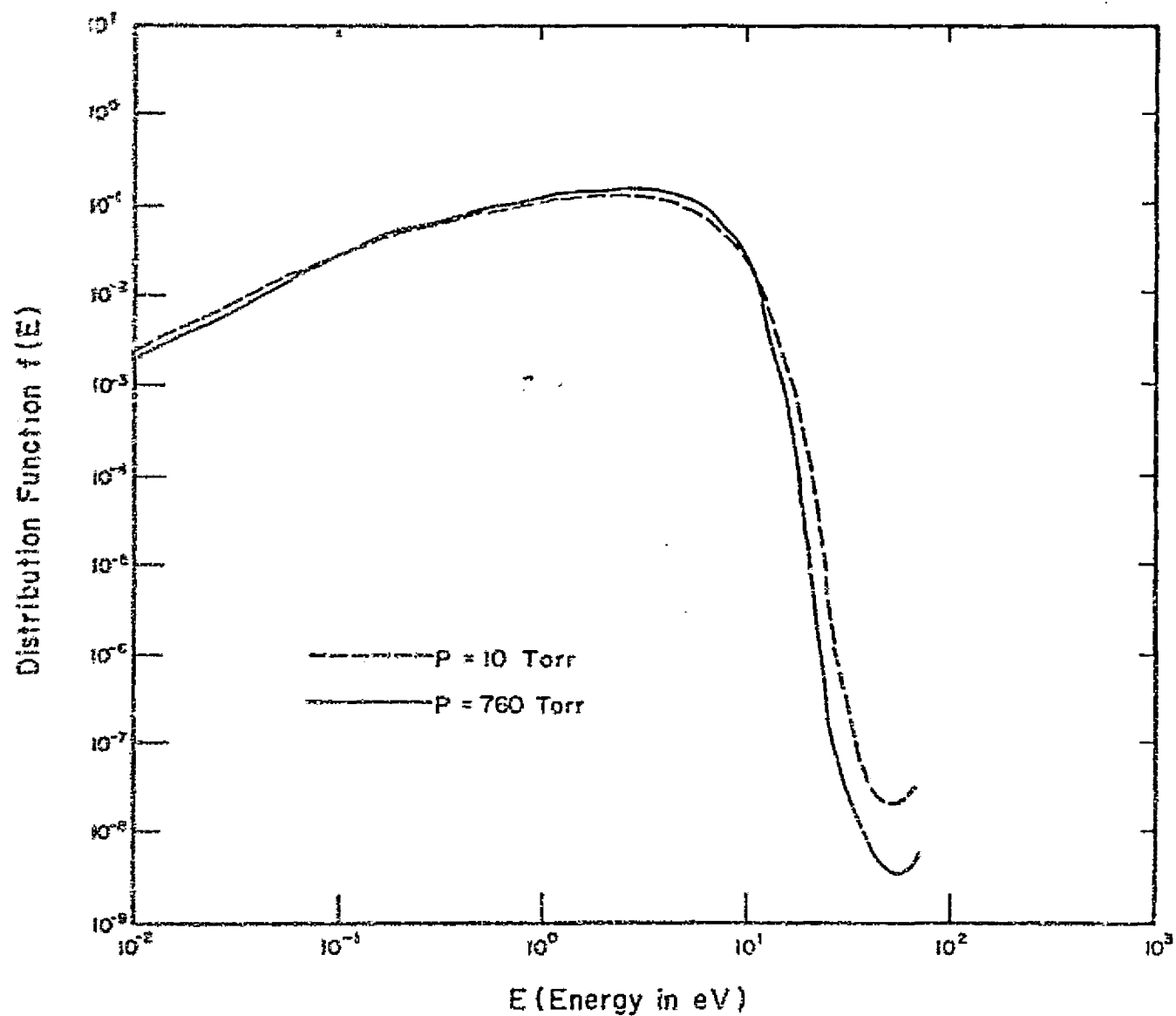


Figure 23. The electron energy distributions for He, with an electric field at  $S_0 = 10^{16}$ ,  $E_0 = 70$  eV.

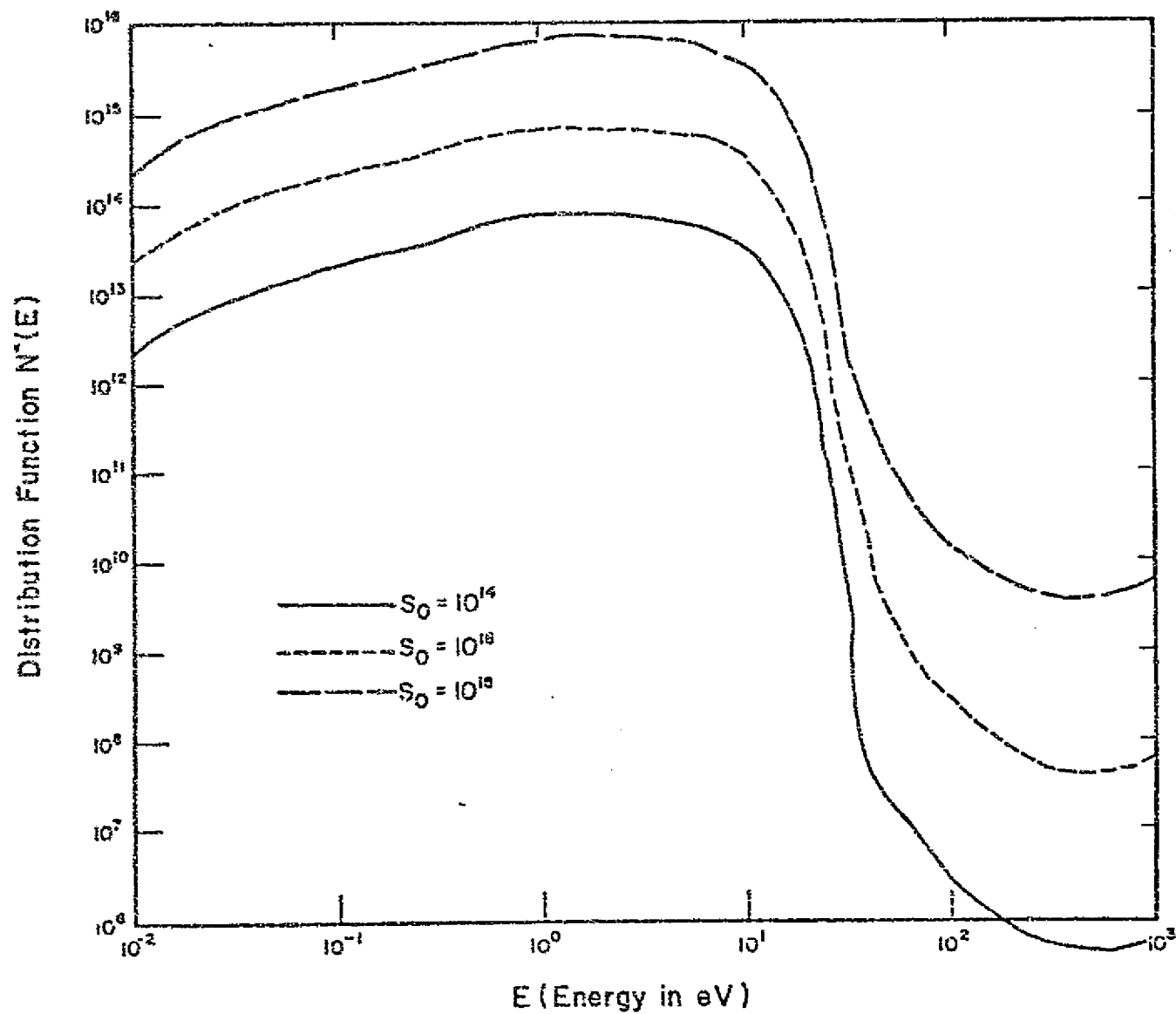


Figure 24. The unnormalized electron energy distribution for He at  $E_f/p = 10$ ,  $p = 10$  torr,  $E_0 = 1000$  eV.

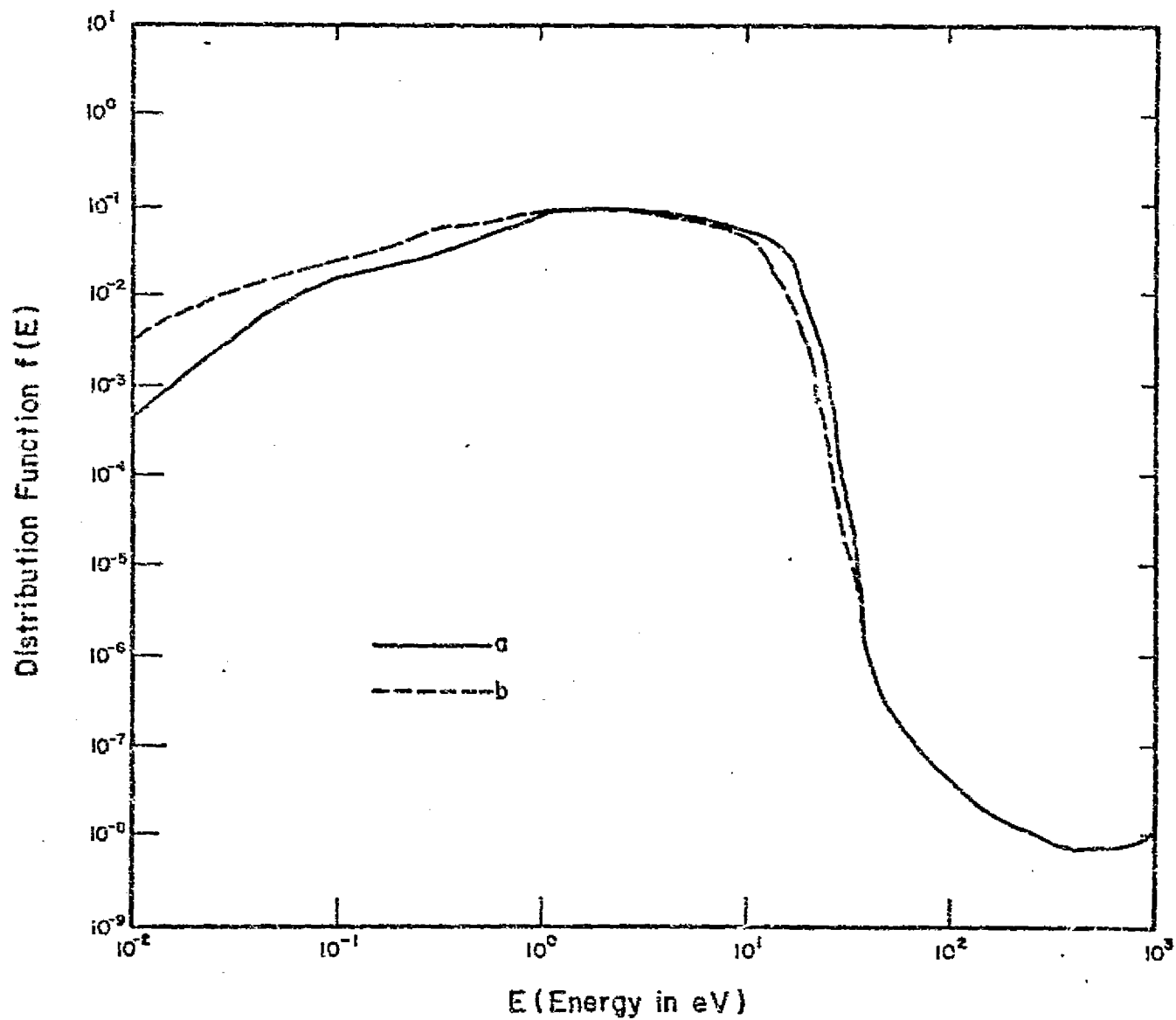


Figure 25. The effect of recombination on the electron energy distribution  
 a) with normal recombination cross section  $\times 100$   
 b) with normal recombination cross section  $\times 1$ .



and  $S_0 = 10^{14}$ , respectively.

In order to show the effect of recombination, a normal case is compared in Figure 25 to a calculation where the absorption cross section was arbitrarily increased 100 times. It is observed that the major effect comes in at low-energy region. More and more electrons recombine as they thermalize and hence, the lower energy region of the distribution is depleted. Since recombination does not occur at high energies, this region is not affected.

The calculated major parameter values for the system with an electric field are summarized in Table 2. As expected on physical grounds the values of the average energies for this case are much higher than the corresponding values with no electric field (compare to Table 1). However, other general trends are the same, e.g. the value of the average energies are slightly higher for higher source rate and/or higher source energy, but are lower for higher pressures.

For comparison the average energy (7.7 eV) found by Smit for an electric field and  $E_f/p = 10$  but with no source present is included in Table 2. It is in general higher than the present values. The reason is not entirely clear. At first thought the additional high-energy source electrons might be expected to raise the average energy. However, as found by T. Ganley (37), the presence of high-energy electrons to produce ionization allows the electrons in the Maxwellian region to decrease in energy, and as a result the overall average may actually decrease.

Table 2. Calculated Parameter Values for Plasma with  
Applied Electric Field ( $E_{f/p} = 10$ )

* P (torr)	$E_o$ (eV)	$S_o$ (#/cm <sup>3</sup> -sec)	$\bar{E}$ (eV)	$N^+$ (#)	Modified W-value range
10	1000	$10^{18}$	8.1	$0.24 \times 10^{15}$	35 ~ 44
10	1000	$10^{16}$	7.5	$0.21 \times 10^{14}$	34 ~ 54
10	1000	$10^{14}$	7.3	$0.24 \times 10^{13}$	37.5 ~ 60
760	1000	$10^{18}$	7.0	$0.23 \times 10^{17}$	37 ~ 50
760	1000	$10^{16}$	6.6	$0.204 \times 10^{16}$	30 ~ 51
760	1000	$10^{14}$	5.4	$0.24 \times 10^{15}$	34 ~ 57
10	70	$10^{18}$	5.86	$0.21 \times 10^{15}$	31 ~ 35
10	70	$10^{16}$	5.18	$0.21 \times 10^{14}$	34 ~ 50
10	70	$10^{14}$	4.46	$0.213 \times 10^{13}$	31 ~ 56
760	70	$10^{18}$	5.7	$0.284 \times 10^{16}$	35 ~ 46
760	70	$10^{16}$	5.0	$0.26 \times 10^{15}$	39 ~ 48
760	70	$10^{14}$	4.3	$0.24 \times 10^{14}$	32 ~ 49
Smit's distribution					
		0	7.7	---	---

\*p: pressure;  $E_o$  = source energy;  $S_o$  : source rate;  
 $\bar{E}$ : average energy;  $N^+$ : number of positive ions.

The values of  $N^+$  are roughly proportional to the square root of the source rate (recall Equation 33) and are independent of the pressure and source energy.

Strictly speaking, the definition of the W-value for the field case should include an allowance for energy gained to electrons in the electric field which then result in ionization. There is no way to include this energy transfer in the present case, so the values listed in Table 2 are based on the ion pair production divided by the energy input with the high-energy source electron alone (no field contribution). This is, then the same formulation (Eq. 65) as that of no field case, and we label the result as the "effective" or "modified" W-value. Again these values fluctuate considerably, so the ranges observed in the last 10 time cycles are given. As would be expected from the neglect of the field contribution, the W-values in this table are in general smaller than that of no field case. As in the earlier calculation, no general trend can be observed, as so far as the dependence of the parameters ( $E_0$ ,  $S_0$ ,  $p$ ) are concerned, and this is consistent with the earlier observations and the literature which indicate that the W-values are fairly insensitive to these parameters.

We have concluded so far that, the distribution functions for non-field case ( $E_f = 0$ ) are more or less assume a Maxwellian distribution in the low-energy region for source rates below  $10^{18}$ . However, when an electric field

is added, the distribution shifts to that found by Smit in the low-energy region. This is, in effect, the well-known Druyvesteyn distribution. It is significant that the addition of a source adds a tail to this distribution, but it does not drastically change the low-energy region. By intuition, it might be expected that Smit's distribution (with electric field or  $E_f \neq 0$ ) would change back to a Maxwellian distribution as the electric field approaches zero. In order to display the transition region, the distribution functions for smaller electric field ( $E_f/p = 1$ ) are plotted in Figure 26 together with the calculated distribution functions at  $E_f/p = 10$  and  $E_f = 0$ . The high-energy tail of the distribution curve are not significantly affected by reducing the electric field  $E_f$ , but the distribution function in the low energy region starts to approach a Maxwellian distribution as expected.

#### D. Conclusion on Monte Carlo Results

Based on the Monte Carlo solutions we obtained so far, we can make the following comments.

1. For the normal range of source rate ( $S \leq 10^{18}$ ), where the neglect of electron-electron interactions remains valid, the Maxwellian plus  $1/E$  distribution is a good initial approximate distribution for the no electric field case. Likewise the Smit (or Druyvesteyn) plus  $1/E$  distribution is a good approximate initial distribution to be used for the system with an electric field present.

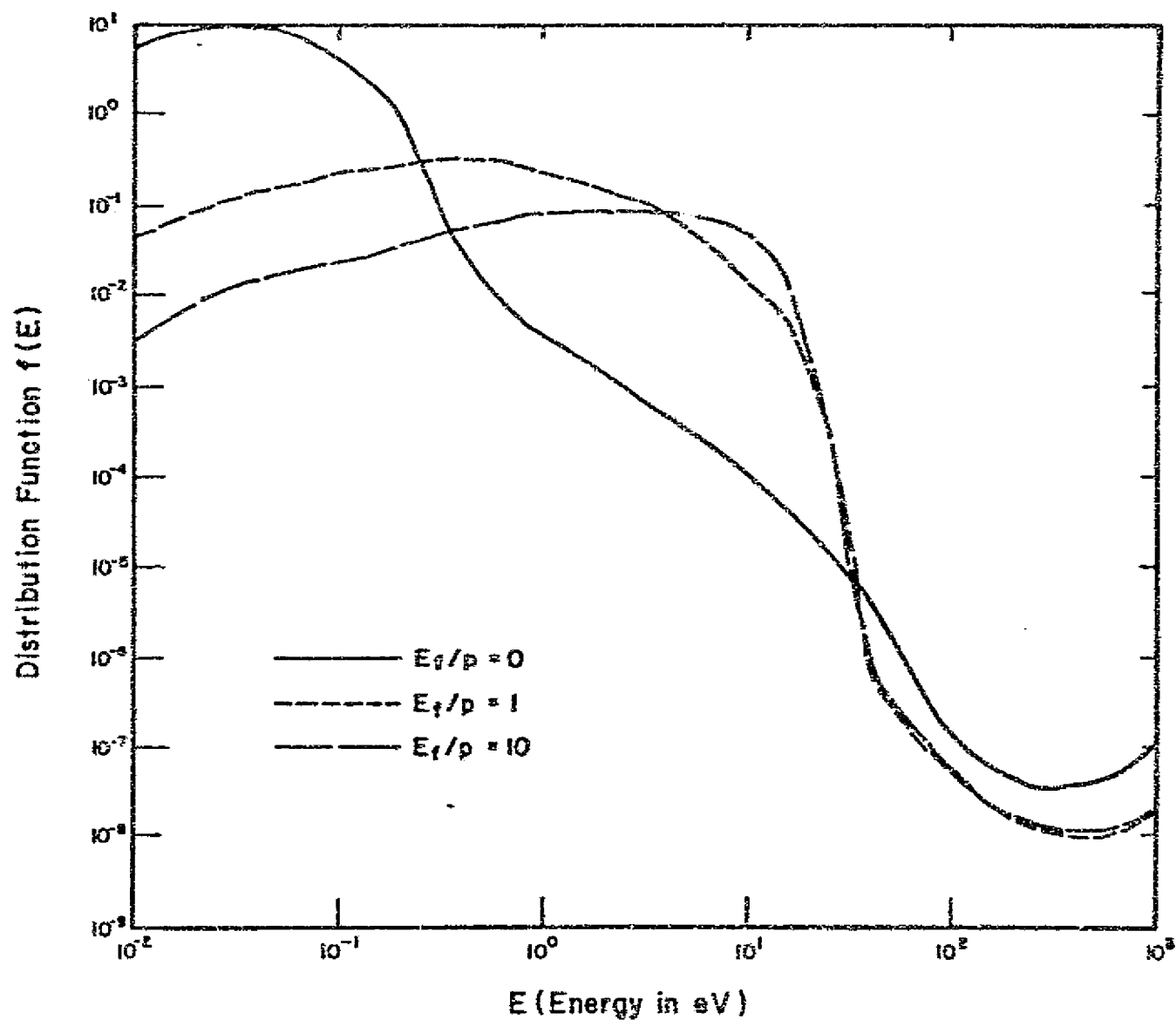


Figure 26. The electron energy distribution for He at  $p = 10$  torr,  $S_0 = 10^{16}$  for  $E_f/p = 0$ ,  $E_f/p = 1$  and  $E_f/p = 10$ .

2. The dependence of the distribution functions (with or without electric field) upon the pressure and source rate only shows up in the high-energy region, where the distribution functions have higher tails for higher source rate, and have smaller values for higher pressures.

3. The range of variations observed in the W-value calculation reflects statistical fluctuations which limit the calculational accuracy of the absolute numerical W-values. Still the results demonstrate that the W-value is more or less independent or insensitive to the changes of pressure, source rate and source energy.

4. Due to a tight budget, the primary emphasis was on increasing the computation efficiency and reducing the computation time to a possible minimum. Thus most of the calculations were carried out with 2000 to 5000 actual particles and for any particular set of parameter values, and the results were generally obtained within 10 minutes of (IBM 360/75) computer time. Compared with hours of computer time and 5 to 10 times more particles used by others for the similar calculations (e.g., see References 3, 4, and 9) the gain in computation efficiency seems to be significant. Of course, a complete comparison is not possible since all of these calculations achieved different accuracies in the final results.

5. The present results also show close agreement with the analytic solutions of Lo and Miley (32,33) on the

following points:

- a. The shape of high-energy portion of the distribution curves are in reasonable agreement as shown in Figure 9.
- b. The high-energy part of the distribution function is not changed by the presence of an electric field for field values  $E_f/p \leq 10$ .
- c. With an electric field, the low-energy part of the distribution function agrees with Smit's result; and this part is not affected appreciably by source rates  $< 10^{10}$  particles/cm<sup>3</sup>-sec.

## CHAPTER VI

## CONCLUSIONS

A. Prelude

The application of Monte Carlo techniques to solve particle transport problems is of fundamental interest in many fields of physics and engineering; especially when analytic and/or numerical solutions of the basic governing equations are too complex to be practical. Yet direct application of Monte Carlo techniques to simulate individual interactions is sometimes prohibitively costly because of the large amount of computational time required. Thus techniques for improving the efficiency are required. Various existing variance reduction techniques can be used, but they are problem dependent. Thus techniques that are efficient for one type of problem might not be as effective for other types of problems. Thus such techniques must be judiciously selected and modified for the particular problem under consideration. One basic underlying principle which applies to any Monte Carlo calculation is: Apply as much known information as one can (given in analytic and/or numerical form) to reduce the uncertainties of the problem. Whenever possible Monte Carlo experiments should be checked and replaced by exact theoretical analysis to reduce uncertainties.

The Monte Carlo experimentalist has to exercise ingenuity in distorting and modifying problems in the pursuit



of variance-reduction techniques. Although Monte Carlo methods are general devices, still a great deal of work depends upon the individual's originality to create special methods or numerical schemes to suit their needs. Unlike the physical experimentalist, the mathematical experimentalist using Monte Carlo Methods has the advantage that his experimental material consists of mathematical objects which can be distorted, controlled, and modified more easily.

The greatest successes of the Monte Carlo method have arisen where the basic mathematical problem itself consists of the investigation of some random processes. However, there are exceptions involving deterministic problem-solution of boundry value problems and partial differential equations. The solutions of these problems are closely connected with the characteristics of certain random diffusion type processes (or can be converted to such type of processes), therefore these problems are reduced to the modeling of such processes. Thus the model and the special techniques developed in present research are quite general in scope, and it should be possible to apply them to many classes of Monte Carlo calculations with a certain amount of modification.

#### B. Summary of Present Work

In this research study, we have devised a general mathematical model for the particle transport or diffusion

type plasma problems which have inherent nonlinear properties due to recombination. Many special variance reduction techniques are incorporated into the system. The model can be used for calculating both time-dependent and steady-state nonlinear transport problems provided that suitable methods exist to calculate, predict, and correct the important parameters which dominate the nonlinearities. A wide range of nonlinear problems can be formulated and handled in this way. The principle idea, the piecewise linearized predict-correct model, is a general, simple, and efficient approach for solving nonlinear problems. In principle, most nonlinear problems can be cast in terms of the basic ideas and algorithms of present work. This represents an extension of earlier works of Musha and Itoh (8), Thomas (9), and Fleck Jr. (11). As for the amount of computation time and efficiency involved, the present study has made significant improvements. Some of the improvements are a natural consequence of the large storage space available in modern computer systems. Additional improvements come from a few simple, but effective, special techniques incorporated into the present model. As described in Chapter II, these include:

- 1) The initial distribution - straightforward particle tracing procedures would require following every simulated particle from its birth (from source) with initial energy  $E_0$  at  $t = t_0$  to the end of its life time history. This type of simulation would require hundreds



of cycle times to achieve a final steady state condition. Instead, we start with an approximately steady-state system by guessing an initial particle energy distribution. Then we simulate the system by tracing particles starting with energy corresponding to this input distribution. This is equivalent to skipping over the initial transient time cycles and greatly reduces the computation time and increases the calculation efficiency.

- 2) The application of the negatively-correlated variates (antithetic variates) technique greatly reduces statistical fluctuations with a minimum number of simulating particles (machine particles).
- 3) The introduction of a weighting scheme and an artificial game of chance (using weighted cross sections) contributed some reduction in calculation time and resolved some practical difficulties.
- 4) For the system with an applied electric field, the exact formulation for a parabolic flight path was used to advance the particles in larger steps than would have been possible using the small-step linear approximation employed by Musha and Itoh (8) and also by Thomas (9). This improves the efficiency appreciably.

Throughout the entire calculation, one basic balance equation (Eq. 30) played a very important role. It merely states that at steady state, the source is balanced by losses.

By means of this equation, we predict the nonlinear parameter values and provide checks for convergence of the solution.

The need for large amounts of storage space for Monte Carlo calculations no longer creates serious problems or drawbacks on modern large computer systems, such as was the case as recent as seven or ten years ago. In addition, techniques exist for further reducing the storage requirement. In the present study, we suggest a technique to eliminate major storage requirements, and use only a small amount of storage as a temporary buffer storage (described later).

In other types of numerical solutions, e.g., in finite difference solutions of the diffusion equation, the step size  $\Delta t$  is restricted by stability requirements. However, in this present model,  $\Delta t$  is only restricted by the degree and amount of nonlinearities present in the system. This restriction is far less stringent.

### C. Practical Limitations and the Cures

It was necessary to solve several difficulties in the present simulation study in order to provide practical solutions. We shall describe them item by item.

- 1) The electron energy distribution is the quantity of major interest. The difference between the magnitudes of its maximum and minimum non-zero values is of the order  $10^9$ . This is due to high density in the low-energy region (thermal energy) and the relatively low density

in the high-energy region (near the source energy).

It is unrealistic to hope to produce reasonable numbers of particles in the low-density (high-energy) region and at the same time have  $10^9$  times as many in the low-energy region. This problem is resolved by combining the following techniques:

- a) Russian Roulette and splitting.
- b) Weight factors which produce fractional particles.
- c) Logarithmic energy intervals. The number of particles in each energy class interval  $\Delta E_i$  is directly proportional to  $f(E_i) \cdot \Delta E_i$  ( $f(E_i)$ : number density at  $E_i$ ). By using logarithmic energy intervals, we have small intervals at low energy region and large intervals at high energy so that the product  $f(E_i) \cdot \Delta E_i$  assumes a practical magnitude.

- 2) The total number of simulating particles (machine particles)  $N_T$  is limited by the storage requirement and the economy of the calculation (computing time is proportional to  $N_T$ ). We have to set an upper limit ( $N_{\max}$ ) on the total number of simulating particles in any time cycle. During each time cycle of duration  $\Delta t$ ,  $N_s$  source particles are introduced into the system, and  $N_z$  secondary particles are produced (by ionization). At the same time  $N_{es}$  particles escape from the system, and  $N_{Ab}$  particles are lost due to recombination.

There are two conditions to be satisfied at all times:

$$a) \quad N_c + N_z + N_s \leq N_T \leq N_{\max} \quad (67a)$$

$$b) \quad N_z + N_s \approx N_{cs} + N_{AB} \quad (67b)$$

where  $N_c$ : the number of census particles (background particles).

Condition a) insures that total number of particles in the system is always under the limit ( $N_{\max}$ ). Condition b) indicates that a particle balance should be satisfied at steady state.

In the program, there are two equivalent problems to be considered. One is the balance of the machine particles [corresponding to condition (a)], and the other is the balance of simulated particles [weighted particles, corresponding to condition (b)]. Both conditions must be satisfied simultaneously. Condition (a) allows a certain amount of leeway since the number of particles that survive the integration cycle (time cycle) and hence require storage will be less than  $N_c + N_z + N_s$ . Here  $N_s$  is chosen to be an input number, the weights of these particles are determined by the source rate. If condition (a) can not be met,  $N_s$  is taken to be

$$N_s = N_{\max} - N_c - N_z - JZ - 1 \quad (68)$$

where  $JZ$  is the number of spatial zones. It has been shown (12) that Equation 68 insures the stabilization of census population without overly reducing the number of source particles.

The proper selection of the energy dividing line (EDL) for Russian Roulette and splitting can insure the satisfaction of condition (b). Also as mentioned in earlier chapters, the balance factor (BF) is used to speed up convergence and force condition (b) to be satisfied sooner.

- 3) In any Monte Carlo calculation, finite samples are used to estimate or simulate the behavior of a large parent population. This is the main source of statistical fluctuations (variance or uncertainty). In the present simulating system, less than 5000 particles were used to simulate approximately  $10^{14}$  physical particles. This creates the problem of matching the values of input and estimated output quantities. Due to the huge scaling factor ( $10^{14}/5000$ ), a small amount of statistical fluctuation or estimation error in the estimated value will be amplified about  $10^{11}$  times. To avoid this, we use the estimated (or equivalent) input quantities which carry the same order of magnitude of statistical fluctuation and weight.
- 4) There are several reasons that make the logarithmic energy scale the logical choice for the group intervals in the energy axis of phase space. These include:
  - a) the wide energy span (0 to 1500 eV)
  - b) the relatively high number density at low energies combined with a low rate and small amount of energy

exchange due to collision processes. In the high-energy region, the situation is just the opposite.

- c) to produce reasonable number of particles in each energy interval [for the reason mentioned in the first section (II-B-1)]. The size and number of intervals are mainly dictated by the range of data (energy range in this case) and number of observations available. Excessive fragmentation of the data may produce many intervals with few occupants. On the other hand, insufficient divisions may obscure important dispersions. Thus an optimum number of intervals will be a compromise between these extremes.

#### D. Future Extensions

The Monte Carlo methods are applicable to the most widely diverse branches of computational mathematics including particle physics (neutron transport for instance), operations research problems such as the investigation of servicing processes, modelling the processes of information transmission, in communication theory, the evaluation of definite integrals, the solution of partial differential equations (e.g. boundary-value problems); the solution of systems of linear equations - inversion of matrices, and many other applications. Problems handled by such methods are in general of two types, the probabilistic and the deterministic, according to whether or not they are directly concerned with the behavior and outcome of random processes.



Their basis lies in simulating statistical experiments by means of computational techniques. Based on the underlying principles described earlier, various variance reduction techniques can be applied. The present work serves as an example and guide for using the important new techniques such as correlated sampling.

- 1) The code developed can readily be used to calculate electron energy distribution functions for other gaseous media by merely changing the cross sections input. The model can also be modified to include the electron-electron and electron-ion interactions although this adds more nonlinearities into the system. This would allow calculation for higher source rates and cases where the fractional ionization is large.
- 2) As for the complex transport problems encountered in areas such as plasmas physics and astrophysics, the model and techniques can readily be extended to the problems with more complex geometry and/or boundary conditions, with inhomogeneous media and multi-dimensions. Also many kinds of nonlinear features can be studied in a realistic fashion.
- 3) More elaborate error checking and correction procedure (based on the known physical phenomena or theoretical formulations) such as the correction matrix used by Berger (9) should be developed to treat the combined effect of statistical fluctuations and other types of

uncertainties. This is essential for generating good results for Monte Carlo calculation.

- 4) More elaborate algorithms or techniques for playing the artificial games of chance by means of weighted cross sections would be a good direction for future research. This must be based on the investigator's intuition and experience, but still it is a promising approach for increasing the efficiency.
- 5) The present model and techniques can be also implemented on smaller computer systems such as PDP-11 or other similar systems where large amount of fast storage space are not available. The following techniques could be used to process the calculation without storing all the particle data (namely  $E_i$ ,  $u_i$ ,  $z_i$ ,  $Wt_i$ ,  $\psi_i$ ). At the beginning of each time cycle, we generate a small number (some reasonable amount,  $\approx 100$  for instance) of particles. The corresponding particle data can be generated from the initial input distribution or from last time cycle. After these particles have been processed, only the accumulated information is stored. A new group of particles come into the system by generating a new set of data, based on the same distributions, which are stored in the same temporary storage space. The previous particle data is destroyed after they have been processed. The same process is repeated until all the necessary prespecified number of particles have been

processed, and the new distributions are obtained. This eliminates the necessity of storing all the particle data at each time cycle, and this eliminates the large arrays that demand the major amounts of storage space. Due to the random nature of the processing algorithms, and the Markovian behavior of the time cycle process (these time cycle processes are a kind of Markov process), this storage saving algorithm can accomplish the desired results.

#### E. Concluding Remarks

In developing any Monte Carlo models and performing calculations, as mentioned in Chapter V the systematic errors and outright blunders must be detected and removed by testing and comparing with known results.

The various variance reduction techniques are used to increase the relative computation efficiency, reduce inherent random errors or statistical fluctuations. For better calculational results, more samples are favorable even though not economic. In any Monte Carlo calculation there exists a minimum number of samples below which the results or statistics are no longer valid and dependable regardless of the variance reduction schemes employed. In the other words, variance reduction techniques help to reduce statistical fluctuations, but they can not be used to remove all the fluctuations caused by small samples.

The minimum sample size is, of course, problem dependent. There are no definite rules or formula for it, thus a cut and try method is necessary until the calculated variance is acceptable.

Monte Carlo calculations are perhaps best in the preliminary stages where they help to give a general idea of the situation and give hints or feelings for trends. If more accurate results are necessary, Monte Carlo calculations may not always be practical good method. In any case, the Monte Carlo method is an important tool in the inferential analysis of the mathematical-physics problems, but it should be viewed as supplementing, rather than replacing, analytic and other numerical methods.

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## APPENDIX

PROGRAM IV G LEVEL 18

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C MASTER PROGRAM OF MONTE CARLO SIMULATION OF MONOTONIC ALPHA PLASMA PAR
C TICLES INDUCED PLASMA ELECTRON ENERGY DISTRIBUTION STUDY
C DEFINITIONS AND TERMINOLOGY
C DI DIAMETER OF THE CYLINDER THE LENGTH OF THE FINITE REGION
C E0 INITIAL ENERGY OF THE SOURCE PARTICLES 50THST SOURCE RATE
C JZ TOTAL NUMBER OF ZONES JE TOTAL NUMBER OF ENERGY RANGES
C IAS1 ARRAY FOR STORING THE INDEX OF ABSORBED PARTICLES NEXT THE NO. OF THEM
C DBI DISTANCE TO BOUNDARY DCOL DISTANCE TO COLLISION DCEN DISTANCE TO CENSUS
C DI ZONE INTERVAL
C ARRAYS E, Z, U STORING ENERGY, DISTANCE, AND DIRECTION OF PARTICLES
C NTP TOTAL NUMBER OF PARTICLES TO BE PROCESSED EACH CYCLE
C NTI TOTAL NUMBER OF TIME CYCLES
C NSI NUMBER OF SOURCE PARTICLES PER UNIT TIME/DISTANCE ON SOURCE RATE
C NTINI NUMBER OF IONIZATION PROCESS OCCURRED EACH TIME CYCLE
0001 REAL NE(55), NEZ( 55,10),EZ(11),EX(11),EX(55),CEZ(11),EXZ(55)
0002 1,101, F10(100),NREZ(11),EZ(11),NTPW
0003 DIMENSION PNE(55), PNRZ(11), NRZ(11),CNE(55), CNRZ(11),TNEY(55)
COMMON D, E(5100),Z(5100),U(5100),IAS(500),IAS(500),NRR,IES(600)
K,NT(5100),SWP
0004 COMMON /AREA1/ NI2N, NO2N, N2S, DIZN(500),NSE, NNET, NEXC, ETAS,
XNNET, NNEXC, WNIZN, ELS,WNAS, WNSE
0005 COMMON /AREA2/ C1,MUO,EI,ES,SP,ANEAM,PDPO,PI,EO,NION,JE,JZ,NASH1,
INSEW,AKT,AKT,ANI,AMCI,EDI,PDNUO,PD2A,K1,K2,K3,WN2,EDZ,ED3
0006 COMMON /AREA3/ XSEL,XELC,XRCB,XNET,XEXC,XION,STOL,KELC
0007 COMMON /AREA4/ CONST,BELZ,ELT,TE,DTI,DIZ,DIZ,DTI,DEL
0008 COMMON /A5/ EG(55),EDS(55),EDD(55)
0009 COMMON /A6/ ESE(50),NUP,NDNR,NON,WSE(50)
0010 REAL MUO, NION,NO
C STATEMENT FUNCTION FOR EVALUATING MAXWELLIAN AND RELATED FONS
0011 FMIE,FNEM)=EXP(-E*AKT)*SQRT(IE)*FNEM
0012 DATA TMS,NT,OTOT,ETOT,NIN,NIZNY,NAST/0.0,1.0,0.0,0.0,0.0,0.0/
0013 DATA HSYLE,CNSIZ,CN55,CNES,CHAS,CNMET,CNEXC,CEASVC,CEASG/9*0.0/
0014 DATA CNE/ 55*0.0/,CEZ/11*0.0/,CNRZ/11*0.0/,TNE/55*0.0/,EKZ/55*0
1.0/.DSAVS/0.0/
0015 DATA NEZ/ 55*0.0/,NE/55*0.0/,EZ/11*0.0/,NREZ/11*0.0/,EEZ/55*0.0/
0016 READ(5,1) NTAPE
0017 READ(5,151) K1, K2, NTC, NTP,NTW,NRR,"KELC",JZ,TH,NCON,NDPH,NTVP
K, MC, IUSN, MUST, NBF
0018 READ(5,152) EDI,NO,EO,ANSM, O, EDP, PDPO, SR
0019 READ(5,153) FID
0020 1 FORMAT ( 110)
0021 151 FORMAT ( 16(5))
0022 152 FORMAT( 8F10.4)
0023 153 FORMAT( 8E10.3)
0024 DELE=1.0
0025 DEL1=1.0/DELE
0026 DEL2=2.0
0027 AKT=0.05
0028 AKT= 1.0/AKT
0029 PI=3.141593
0030 EMASST=1.60203E+16/9.108
0031 K3=K1*K2
0032 WN2=.0/K2

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FORTRAN IV G LEVEL 16

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0033      CJD=0.0
0034      E1=24.46
0035      EF=EDP*PDPO
0036      EIS=21.4
0037      C1=2.0*0.000549/4.003673
0038      ETV= SORT(2.0*1.60203/9.108)*1.0E+09
0039      VTE= 1.07*(ETV*ETV)
0040      CALL UNDERZ('DP')
0041      PD20=PDPO*26.0
0042      MUQ=(1.0E-09)/ETV
0043      CMUQ=1.0E-09
0044      CMUPDP=CMUQ*PDPO
0045      IF (PDPO.GT.10.0) CMUPDP=CMUQ*10.0
0046      PDHMQ=CMUPDP/ETV
0047      ELS=1.0
0048      NUP=0
0049      NDN=0
0050      NDNR=0
0051      INO=0
0052      NRR=0
0053      NAS=0
0054      NES=0
0055      NSE=0
0056      NDZN=0
0057      NIZN=0
0058      NIN=0
0059      LK=0
0060      INIY=999999
0061      NHET=0
0062      NSKC=0
0063      ETAS=0.0
0064      ETEP=0.0
0065      MTDS=0
0066      HW=0
0067      NSTP=0
0068      NCOL=0
0069      WNHET=0.0
0070      WNLZN=0.0
0071      WNEXC=0.0
0072      WNAS=0.0
0073      WNES=0.0
0074      WNSE=0.0
0075      CLOSS=0.0
0076      CSFZ=0.0
0077      CNIO2=0.0
0078      CSF1=0.0
0079      CETAS=0.0
0080      CETEP=0.0
0081      DTAVS=0.0
0082      JE=48
0083      IE=JE
0084      DEN=1.5+E+16*PDPO
0085      DETZ=0/JZ

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FORTRAN IV G LEVEL IB

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0086      DZIV= 1.07DRT1
0087      CONST=1.0/ALOG(2.0)
0088      ED2=0.44Y3I3
0089      ED3=19.8
0090      PM=EDE*AKY
0091      AM1=2.0*AKY/SQRT(PI*AKY)
0092      CALL RNSTRZ(244363211)
0093      SO=POPD*SR*1.0E+16
0094      SF=1.0
0095      AITG= CMUPON / (AKT*P())
0096      DE1=0.518
0097      DE2=0.355
0098      DE3=0.3
0099      D11=1./0.016
0100      D12=1./DE2
0101      D13=1./DE3
0102      ANC1= SQRT(2.0*2.7183*AKY)*DE2
0103      DO 11 1=6, 36
0104      EBL = (1-5)*DE2 -4.0
0105      EMV= EBL -0.1685
0106      ESD(1)= 2.0**EBL
0107      EG(1-1)= 2.0**EMV
0108      11 EDE(1)=EBD(1)*(2.0**DE2 -1.0)
0109      JE1=JE+1
0110      DO 12 1=37, JE1
0111      EBL= 7.005 +(1-36)*DE3
0112      EMV= EBL - 0.1685
0113      EBD(1)= 2.0**EBL
0114      EG(1-1)= 2.0**EMV
0115      12 EDE(1)= EBD(1)*(2.0**DE3 -1.0)
0116      DO 10 1=1, 5
0117      EDE(1)= 0.016
0118      EG(1)= EDE(1)*1 -0.006
0119      10 EBD(1)= EDE(1)*(1-1)
0120      EDE(36)=29.69
0121      EG(36)= 1737.482
0122      IF (INTAPE.NE.1) GO TO 2
0123      READ(10) EAVG, DAVG, TME, NTM, NT, NTP, NION, DDEG, NSM, SWF, E, U
0124      READ(10) Z, WT, ANIOS, SF3
0125      NS=NSM*K2
0126      NGHT=1.0
0127      REWIND 10
0128      NT=NT
0129      2 CONTINUE
0130      IF (INTAPE.NE.2) GO TO 3
0131      READ(10) EAVG, JAVG, TME, NTM, NT, NTP, NION, DDEG, NSM, SWF, E, U
0132      READ(10) Z, WT, ANIOS, SF3
0133      NS=NSM*K2
0134      NGHT=1.0
0135      REWIND 9
0136      NT=NT+1
0137      3 CONTINUE
0138      IF (INTAPE.NE.0) GO TO 122

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FORTRAN IV, G LEVEL 18          MAIN          DATE = 72126          05/31/62

0139      NION= SQRT(SO*10/AITG)
0140      ANEAM=NION*0.33333/DEN
0141      CALL SOURCE2(NTP, EAVG, NTW, PM, FID)
0142      AL=ALDALEAVG)
0143      DAVG=AL
0144      DDEG=DAVG*ND/SQRT(EAVG)
0145      SF2=1.0
0146      IF (NION.NE.0) SF2=NTW/NION
0147      SN=SO*AL*ND*SF2/(ETV*SQRT(EAVG))
0148      NSW=ANSW+0.5
0149      IF (SN.GT.ANSW) NSW=SN+0.5
0150      NS=NSW*K2
0151      SWF=NSW/SN
0152      SF=SWF
0153      SF1=SF2
0154      SF3=SF2
0155      CALL SOURCE(NS,NTP)
0156      TME=0.0
0157      ANIOS=NION
0158      WGHY=1.0
0159      *122 CONTINUE
C GENERATE SOURCE PARTICLES TO INITIALIZE THE PROGRAM
0160      ANEAM=NION*0.33333/DEN
0161      WRITE(6,70)
0162      70 FORMAT('1','PROGRAM INITIALIZATION AND THE INITIAL PARAMETERS',//)
0163      WRITE(6,71)AITG, NION, NTP, AL,EAVG
0164      71 FORMAT('1','AITG =',E14.6, 3X,'NION =',E14.6,3X,'NTP =',I12,3X,
0165      1'DCEN(E0) =',E14.6,3X,'EAVG =',E14.4,/)
0166      72 FORMAT('1',8X,'DCOL',13X,'DCEN',13X,'ENERGY',11X,'2',14X,'U',10X,
0167      1'WEIGHT',/)
0168      75 FORMAT('1',F14.6,3X, F14.6, 3X, F14.5,3X, F12.4,3X,F12.4,3X,E13.4)
0169      WRITE(6,72)
0170      RIP=RAN32(0)
0171      I=1
C WITHIN THE OUTER I LOOP THE I TH PARTICLE HISTORY
C GEOMETRIC ROUTINE ENTERING HERE
0170      7777 IF ((E(I).LE.0.01.OR.(E(I).GT.E0*1.5)) GO TO 9008
0171      TCEN=DDEG/ETV
0172      7778 IF ((NIZN.GT.0).AND.(I.GT.NTP)) TCEN= DIZNINOZN+1)
0173      U(I)=U(I)*IUSN
0174      LFG2=0
0175      ETAG1=1.0
0176      IF (E(I).GT.E01) ETAG1=-1.0
0177      SQE= SQRT(E(I))
0178      OVA=EF*TCEN*EMASSI*0.5
0179      VZA=SQE*ETV*U(I) + OVA
0180      VYS=E(I)*ETV*(1.0-U(I))*U(I)*ETV
0181      DCEN= SQRT(VZA**2+VYS)*TCEN
0182      NPLG=0
0183      900 IF (E(I).LE.0.01 GO TO 9008
0184      IF (NPLG.EQ.0) GO TO 809
0185      SQE= SQRT(E(I))
0186      OVA=EF*TCEN*EMASSI*0.5

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0187      VZA=SQE*E1V*U11) + DVA
0188      809 DZ=VZA*TCEN
0189      ENEW= E11) + EF*DZ
0190      IF (ENEW.LE.0.0) GO TO 9008
0191      SQEP=SQRT(ENEW)
0192      UNEW= SQE*U11)/SQEP +DVA*2./((SQEP*ETV)
0193      AL1=ALDA(E11)
0194      AL2=ALDA(ENEW)
0195      AL1=1.0/AL1 + 1.0/AL2
0196      ALD=2.00/AL1
0197      RI=RAN32(0)
0198      ALR=ABS(ALOG(RI))
0199      DCOL= ALR*ALD
0200      IF (INFLG.EQ.1) GO TO 901
0201      IF (MOD(I,INT.EQ.0) WRITE(6,75) DCOL,DCEN,E11),Z11),U11),WY1)
0202      DCOL2=ABS(ALOG(1.-RI))*ALD
0203      DTOT=DTOT +DCOL*DCOL2
0204      ETOT=ETOT+E11)
0205      IF (NDPN.LY.2) GO TO 901
0206      TCEN2=TCEN
0207      DCEN2=DCEN
0208      E10=E11)
0209      U10= U11)*TUSN
0210      Z10=Z11)
0211      ENW0=ENEW
0212      UNW0=UNEW
0213      DZ0=DZ
0214      901 NPLG=1
0215      902 IF (DCEN.LY.DCOL) GO TO 905
0216      TCEN=TCEN - DCOL*TCEN/DCEN
0217      DZC=DZ*DCOL/DCEN
0218      E11)=E11) + EF*DZC
0219      Z11)=Z11) +DZC
0220      DCEN= DCEN -DCOL
0221      NCOL=NCOL+1
0222      IF (MOD(NCOL,MC).EQ.0) RIP=RAN32(0)
0223      IF (MOD(NCOL,MC).NE.0) RIP=1.0-RIP
0224      CALL COLPRO(I, TCEN, NTP, RIP, MUST, 69008, 69045)
0225      IF (E11).LE.0.0).AND.(LFG2.EQ.1)) GO TO 9008
0226      GO TO 900
0227      905 E11)=ENEW
0228      U11)= UNW0
0229      Z11)= Z11) + DZ
0230      IF ((Z11).LE.0.0).OR.(Z11).GT. 0 )) GO TO 9008
0231      J=INT(Z11)*DZ1V)+1
0232      IF (J.GT.J2) J=J2
0233      K= INDEXE(E11)
0234      IF (K.GT.JE) K=JE
0235      IF (K3.EQ.1) GO TO 904
0236      ETAG2=1.0
0237      IF (E11).GT.ED1) ETAG2=-1.0
0238      IF ((ETAG1*ETAG2).GT.0.0) GO TO 904
0239      IF (ETAG2.GT.0.0) GO TO 903

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0240      NUP=NUP+1
0241      ESE(NUP)=E(I)
0242      WT(I)=WT(I)/K3
0243      WSE(NUP)=WT(I)
0244      GO TO 904
0245      903 IF((LFG2.EQ.1).OR.(NDPN.LT.2)) NDNR=NDNR+1
0246      IF((LFG2.EQ.1).OR.(NDPN.LT.2)) IAA(NDNR)=1
0247      NDN=NDN+1
0248      IF (MOD(NDN,K3).NE.0) GO TO 9045
0249      WT(I)=WT(I)*K3
0250      904 NEZ(K,J)= NEZ(K,J) + WT(I)
0251      EKZ(K,J)=EKZ(K,J)+E(I)*WT(I)
0252      9045 IF((LFG2.EQ.1).OR.(NDPN.LT.2)) GO TO 9009
0253      LFG2=1
0254      E(I)=EIO
0255      U(I)=UIO
0256      Z(I)=ZIO
0257      TCEN=TCEN2
0258      DCEN=DCEN2
0259      DCOL=DCOL2
0260      ENEN=ENNO
0261      UNEN=UNNO
0262      DZ=DZO
0263      IF (IUSN.LT.0) GO TO 900
0264      GO TO 902
0265      9008 NES=NES+1
0266      IF (E(I).LT.ED1) WT(I)=WT(I)/K3
0267      IF (WT(I).GT.1) WT(I)=1.0
0268      WNES=WNES +WT(I)
0269      IES(NES)=1
0270      IF (E(I).GT.0.0) ETEP=ETEP+E(I)*WT(I)
0271      9009 I=I+1
0272      LFG2=0
0273      IF (I.LF.NTP) GO TO 7777
0274      IF ((I.GT.NTP).AND.(NIZN.GT.0)) NOZN=NOZN+1
0275      IF (NOZN.LT.0) NIN=1
0276      IF ((NIZN.GT.0).AND.(NIN.EQ.0)) GO TO 7778
0277      NIZN=WNIZN+0.5
0278      WNS=WNES+0.5
0279      NASW=WNAS+0.5
0280      WNAS=WNAS/KELC
0281      WNSE=WNSE/KELC
0282      NSEW=WNSE+0.5
0283      DO 191 M=1,J
0284      DO 191 N=1,JZ
0285      EE(M)=EE(M) + EKZ(M,N)
0286      191 NE(M)= NE(M) + NEZ(M,N)
0287      IF ((K3.EQ.1).OR.(NUP.EQ.0)) GO TO 24
0288      N3=K3/JZ
0289      IF (N3.LT.1) M3=1
0290      NK3=K3-1
0291      IF (NK3.LT.1) IZ=NK3
0292      IF (N3.GE.1) M3=N3

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0293      DO 23 N=1,NUP
0294          K=INDEXE(ESE(N))
0295          EE(K)=EE(K) + (K3-1)*ESE(N)*WSE(N)
0296          NE(K)=NE(K) + (K3-1)*WSE(N)
0297      22 IF (NK3.GE.JZ) IZ=JZ
0298          DO 21 L=1,IZ
0299          NREZ(L)=NREZ(L) + M3*WSE(N)
0300      21 EZ(L)=EZ(L) + ESE(N)*M3*WSE(N)
0301          NK3=NK3 -JZ*M3
0302          IF (NK3.LT.JZ) IZ=NK3
0303          M3=1
0304          IF (NK3.GT.0) GO TO 22
0305      23 CONTINUE
0306      24 CONTINUE
0307          DO 192 M=1, JZ
0308          DO 1915 N=1,JE
0309          NREZ(M)=NREZ(M) + NEZ(N,M)
0310      1915 EZ(M)=EZ(M) + EKZ(N,M)
0311      192 CONTINUE
0312          NTPN=NTP*MIZN-NES-NAS
0313          NTPW=0
0314          ETOL=0.0
0315          DO 14 M=1,JE
0316          ETOL=ETOL+EE(M)
0317      14 NTPW=NTPW+NE(M)
0318          EWTOL=ETOL
0319          NTP=NTPN
0320          DAVG=DAVG*ND
0321          DAVS=(DAVG*1000.)*NTPW
0322          DSAYS=DSAYS+DAVS
0323          DTAYS=DTAYS+DAVS
0324          DISG=1.0/DSAYS
0325          ANTP=1.0/NTPW
0326          NSTP=NSTP + NTPW*0.5
0327          NTPS=NTPS+ NTPN*DAVS
0328          NTP=NTP
0329          EAVG=EWTOL*ANTP
0330          NSQ=NS
0331          NSQ=NSQ*HNZ*0.5
0332          NTP=NTP*MIZN
0333          NTPP=NTP
0334          DAVG=DTOT*0.5/NTP
0335          EAVG=ETOT/NTP
0336          DEG=ODEG/ETV
0337          TME=TME+DEG
0338          TMS=TMS+DEG
0339      C IF SOURCE PARTICLES SO(N) NES THEN ELIMINATE AS MANY RESIDUE PARTICLES AT TH
0340          NTPS=NTP+ NWS +NASW
0341          NTP=NTP+0.5
0342          OCSN=ALDAI EAVG
0343          DDEG=DAVG*ND/SQRT(EAVG)
0344          BF=1.0
0345          IF ((HNES+WNAS).GT.0.0) BF=(WNIZ+SN)/(WNES+WNAS)

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0345      SF2=1.0
0346      IF (ANIOS.NE.0) SF2=NTPS/ANIOS
0347      ESAVG=0.0
0348      IF (WNES.NE.0.01) ESAVG= ETEP/WNES
0349      EASAVG=0.0
0350      IF (MNAS.NE.0) EASAVG=ETAS/MNAS
0351      SWFO=SWF
0352      CN55=CN55+ SN=DAVS*WGHT
0353      CSF1= CSF1 +(ETEP+ETAS+40.*WNIZM)/TMS
0354      SF1=CSF1*0.15G/150*EO1
0355      CSF2= CSF1 + SF2*DAVS
0356      IF ((MOD(INT,9),EQ,0),AND,(INT.NE.0)) SF3=CSF2*0.15G
0357      WGHT=SF1/SF3
0358      SN=SQDDEG*SF3/ETV
0359      IF (INT.GT.+920) ANSW=2
0360      NSN=ANSW+0.5
0361      IF (SN.GT.ANSW) NSN=SN+0.5
0362      SNF=NSN/SN
0363      SF=NSN*K2
0364      SF=SWF
0365      NMEN=NMEN+0.5
0366      NEXN=NEXC+0.5
0367      WSVLE=WSVLE+ NVALUE*DAVS
0368      CETEP=CETEP + ETEP*DAVS
0369      CETAS=CETAS + ETAS*DAVS
0370      CN5IZ=CN5IZ+WNIZN*DAVS
0371      CNES=CNES+WNES*DAVS
0372      CNAS=CNAS+MNAS*DAVS
0373      CNMET=CNMET+NMEN*DAVS
0374      CNEXC=CNEXC+NEXC*DAVS
0375      CESAVG=CESAVG+ ESAVG*DAVS
0376      CEASG=CEASG+EASAVG*DAVS
0377      CLOSS= CLOSS+(WNIZN-WNES)*DAVS/(TMS*SF3)
0378      WVLE1=WSVLE*0.15G
0379      NSAVG=CN5S*0.15G+0.5
0380      NIZAVG=CN5IZ*0.15G+0.5
0381      ANIZVG=CN5IZ*0.15G
0382      ANSVG=CN5S*0.15G
0383      AETEP=CETEP/DTAVS
0384      AETAS=CETAS/DTAVS
0385      IF (ANIZVG.NE.0) WVLE2=(EO*ANSVG-AETEP-AETAS)/ANIZVG
0386      ALOSS=CLOSS*0.15G
0387      ANES=CNES/DTAVS
0388      ANAS=CNAS/DTAVS
0389      AESAVG=CESAVG*0.15G
0390      AEASG=CEASG*0.15G
0391      ANMET=CNMET*0.15G
0392      ANEXC=CNEXC*0.15G
0393      ANS=ANES+ ANAS -ANIZVG
0394      IF (ANS.LE.0.01) ANS=NSW/SWF
0395      IF (WNIZN.NE.0.0) NVALUE= (SN*EO*WGHT-ETEP- ETAS)/WNIZN
0396      IF (ANIZVG.NE.0) WVC=(ANS*EO-AETEP-AETAS)/ANIZVG
0397      IF (INS.GE.1) TMS=0.0

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0398      NET=NE5+NAS
0399      CALL SOURCE(NS,NTP)
0400      IF (NET.GT.0) CALL FILHOL(NTP,NE5, NAS)
0401      90 NIZNT= NIZNT +NIZW
0402      IF (NUP .GT.0) CALL SEPROC(NTP, K3)
0403      IF (NDNR.GT.0) CALL RRPROC(NTP,K3)
0404      NAST= NAST +NWS
      C INTRODUCE THE SOURCE PARTICLES FOR THE NEXT CYCLE
      C EVERY FIFTH CYCLE RENORMALIZE NION AND OUTPUT QUANTITIES ARE PRINTED
      C RENORMALIZATION PROCESS OF NION
0405      ANSTP= 1.0/NSTP
0406      ANTP5=DSAVS/NTP5
0407      NSTPN=NSTP
0408      DO 16 M=1,JZ
0409      CNREZ(M)= CNREZ(M) + NREZ(M)*DAVS
0410      16 CEZ(M)= CEZ(M) + EZ(M)*DAVS
0411      DO 18 M=1,JZ
0412      FNREZ(M)=NREZ(M)*ANTP5
0413      18 EZ(M)= EZ(M)/ NREZ(M)
0414      SUMNE=0.0
0415      DO 193 M=1,JE
0416      SUMNE= SUMNE + NE(M)
0417      TNE(M)= TNE(M) + NE(M)/EDE(M)
0418      193 FNE(M)=NE(M)*ANTP5/EDE(M)
0419      EAV3=0.
0420      DO 1755 M=1, JE
0421      EAV3= EG(M)*FNE(M)*EDE(M) +EAV3
0422      1755 CNE(M)= TNE(M)*ANSTP
0423      CNION=CMUPDP*2.0*AKT/SQRT(AKT*PI)
0424      ANION=0.0
0425      ANIO2=0.0
0426      DO 20 M=1,JE
0427      ANIO2= ANIO2 + FM1(EG(M),CNE(M))*EDE(M)
0428      ANION= ANION + FM1(EG(M),FNE(M))*EDE(M)
0429      20 CONTINUE
0430      DNION=ANION*CNION
0431      DNIO2=ANIO2*CNION
0432      CJD=ALOSS
0433      IF (ALOSS.LT.0.0) CJD=0.0
0434      IF (ABS(CJD).GT.110.*50) CJD=110.*50
0435      TNION=ABS(SO+CJD)/DNION
0436      TNIO2=ABS(SO+CJD)/DNIO2
0437      TNIO2=SQRT(TNIO2)
0438      TNION=SQRT(TNION)
0439      CNIO2=CNIO2 + TNIO2*DAVS
0440      ANIOS=CNIO2*0.156
0441      WF=1.0
0442      IF (ANES+ANAS,NE.0.) WF=(ANIZVG+ANSVG)/(ANES+ANAS)
0443      IF (MOD(INT,NPM),NE.0) GO TO 1790
0444      NION=ANIOS
0445      IF ((WF.GT.0.005).AND.(WF.LT. 200.)) NION=ANIOS*WF**NBF
0446      1790 CONTINUE
0447      DAVT=DAVG*ND

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0448      ALPHA=MNIZM/(NTPW*DAVT)
0449      ANEAM=MION*0.33333/DEEN
0450      1785 FORMAT('0', 'ANION=', E13.4, 3X, 'ONION=', E13.4, 3X, 'TNION=', E13.4, 3X,
      X'ALPHA=', F12.5, 4X, 'LEAK=', E13.4, /)
0451      WRITE(6,1785) ANION, ONION, TNION, ALPHA, ALOSS
0452      WRITE(6,171) EO, POPO, D, DET2, DEG, NDPN
0453      171 FORMAT('1', 'GAS: HE', 2X, 'SOURCE ENERGY', F9.0, 3X, 'PRESSURE', F4.0,
      X'TORR', 3X, 'SLAB THICKNESS', F4.1, ' CM', ' DX', F4.1, ' CM', 4X, 'DELT
      X=', E12.5, 3X, 'PROC=', I2, /)
0454      WRITE(6,180) NT, TME, NS, NSH
0455      180 FORMAT(' ', 'NUMBER OF TIME CYCLES=', I6, 3X, 'TIME =', E12.5, ' SEC',
      13X, 'GENERATED SOURCE PART =', I5, 3X, 'MGTD # OF SOURCE PART.=',
      I15)
0456      WRITE(6,1805) DAVT, OCEN1, EAVG, EAVG, EAV3
0457      1805 FORMAT(' ', 'DAVG =', F12.5, 3X, 'OCEN =', F12.5, 3X, 'OVERALL EAVG=',
      F10.5, 3X, 'MGTD EAVG=', F10.5, 3X, 'EAVG CHECK=', F10.5, /)
0458      WRITE(6,181)
0459      181 FORMAT(' ', 1X, 'NUMBER DISTRIBUTION', 2X, 'FLUX DENSITY', 4X, 'AVERAGE
      1E ENERGY', 3X, 'CUMULATED # DISTR', 3X, 'ENERGY INTERVAL', 2X, 'CUM FLUX
      XDEN', 4X, 'SDEN-DEVIATION', /)
0460      CESUM=0.0
0461      DO 281 M=1, JF
0462      CESUM=CESUM + CNE(M)*EDE(M)
0463      IF (NE(M).GT.0) DEAVG=EG(M)-EE(M)/NE(M)
0464      IF (NE(M).LE.0) DEAVG=0.0
0465      IF (NE(M).EQ.0) DNAGV=0.
0466      FNOV=CNE(M)-FNE(M)
0467      VFE=EG(M)-ETV
0468      FLX1=FNE(M)*VFE
0469      FLX2=CNE(M)*VFE
0470      IF (NE(M).EQ.0) GO TO 281
0471      IF ((M.EQ.1).AND.(DEAVG.GT.0.0)).OR.((M.EQ.JF).AND.(DEAVG.LT.0.0))
      X) DNAGV=-DEAVG*(FNE(M)+0.5/EDE(M)
0472      IF ((DEAVG.LT.0.0).AND.((M+1).LE.JF)) DNAGV=-DEAVG*(FNE(M+1)-FNE(M)
      X)/((EG(M+1)-EG(M))
0473      IF ((DEAVG.GT.0.0).AND.(M.GT.1)) DNAGV=DEAVG*(FNE(M)-FNE(M-1))/((EG(
      X)M-EG(M-1))
0474      281 WRITE(6,182) M, FNE(M), FLX1, EG(M), CNE(M), EDE(M), FLX2, DNAGV
0475      182 FORMAT(' ', I3, E12.5, 3X, E12.4, 5X, F12.6, 8X, E12.5, 6X, F12.6, 4X, E12
      X5.4X, E12.4)
0476      DET=DDEG/ETV
0477      WRITE(6,183) NTPP, NSTPN, NTPW, DET, DAVGO, NCOL
0478      183 FORMAT(' ', //, 3X, 'TOTAL # OF PARTICLES PROCESSED =', I5, 3X, 'CUMULAT
      1ED # OF PARTICLES =', I10, 3X, 'MGTD TOTL # OF PARTICLES=', F12.3, //,
      X' NEW DT=', E12.5, 4X, 'OLD DAVG=', F12.5, 4X, ' # OF COLLISIONS=', I5, //)
0479      WRITE(6,187)
0480      187 FORMAT(' ', 1X, 'ZONE NUMBER', 5X, 'AVERAGE ENERGY', 5X, 'NUMBER O
      1F ELECTRONS', 3X, 'CUMULATIVE # DEN(Z)', 3X, 'CUMU. AVG ENERGY', 3X,
      1'ZONE # DENSITY', /)
0481      DO 282 M=1, JZ
0482      ANZ=CNREZ(M)*ANSTP*DISG
0483      PCZ=CEZ(M)/CHREZ(M)
0484      282 WRITE(6,184) M, EZ(M), NREZ(M), ANZ, PCZ, PNREZ(M)

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0485 184 FORMAT(1, 5X, 15, 5X, F14.6, 5X, F14.6, 5X, F14.6, 7X, F14.6, 7X,
      1F12.5)
0486 WRITE(6,185) NION,WNIZN,NIZN, ANIZVG,WNES, NES,ANES, WNAS,NAS,
      XANAS, NTP, K1, K2, WF,BF
0487 185 FORMAT('0', 'NUMBER OF POSITIVE ION=',E14.6,/, 'NUMBER OF IONIZAT
      IONS OCCURRED=',F9.3,5X, 'ACTUAL #=',15,3X, 'CUMULATIVE AVG=', F10.3,
      X/, ' # OF ESCAPED PARTICLES=',F9.3,5X, 'ACTUAL #=',15,3X, 'CUMULATIVE
      X AVG=',F9.2,/,/, ' # OF ABSORBED PARTIC=',E12.4,3X, 'ACTUAL #=',15,3X
      X, 'CUMULATIVE AVG=',F9.2,/,/, ' TOTAL # OF PARTICLES IN THE NEW CYCL
      E =',15,/,/, ' K1=',12, 3X, 'K2=',12, 4X, ' BALANCE FACTOR WF=',
      XF10.5,3X, 'BF2=', F10.5,/)
0488 WRITE(6,186) ED1, NPM, ND, KELC
0489 186 FORMAT(' ENERGY DIVIDING LINE FOR RR PLAY E=',F10.4,4X, 'EVERY',12
      X, 'TH CYCLE CHANGE N=',/,/, ' TIME INTERVAL MULTIPLY FACTOR=',F9.4,
      X3X, 'ELASTIC XSEC / DV',12,/)
0490 NDR=NDN-NDN/K3
0491 WRITE(6,186) NSO, NSW, ANSVG,NOR,WNSE,WNNET,ANNET,WNEXC,ANEXC,NUP,
      XWGHT
0492 186 FORMAT(' ', ' SOURCE PARTICLES GENERATED OLD =',110,/,/,
      1' SOURCE PARTICLES (WEIGHTED) IN THE NEW CYCLE =',110,3X, 'CUMULAT
      XIVE AVG=',F10.3,/,/
      1' # OF PARTICLES KILLED OFF IN RR PROC=',110,/,/,
      1' # OF SUPER ELASTIC COLLISION=',F10.4,/,/,
      1' # OF EXCITATIONS TO META STABLE LEVEL='
      X,F10.4, 3X, 'CUMULATIVE AVG=', F9.3,
      1' //, ' # OF EXCITATIONS TO OTHER LEVELS=',F10.4,
      X4X, 'CUMULATIVE AVG=',F9.2,/,/, ' # OF PARTICLES CROSS OVER TO HIGH
      XENERGY REGION NUP=',15,4X, 'OVERALL WT=',E12.4,/)
0493 WRITE(6,195) EF, EDP, PDPD, SF1, SF2, SF3
0494 195 FORMAT(' E-FIELD=', F9.2, 'V/CM', 4X, 'E/P=',F9.2,4X, 'P/PO =',F7.1,
      X3X, 'SF1=', E12.4,3X, 'SF2=',E12.4,3X, 'SF3=', E12.4,/)
0495 WRITE(6,170) SO,WVC,SWFO,WVALUE,WVLE1, WVLE2,ESAVG, AESAVG,EASAVG,
      XAEASG
0496 170 FORMAT(' SOURCE RATE =',E11.3, ' PER SEC',4X, 'W-VALUE(LOCAL )=',
      XF10.3, 3X, 'SOURCE RED.FACT=',E13.4,/,/, ' W-VALUE=',
      XF10.3,4X, 'CUM AVG WV1=',F9.3,3X, 'CUM AVG WV2=',F9.3,
      1' //, ' AVERAGE ENERGY OF ESCAPED PARTICLES=',F12.5,3X, 'CUMULAT
      XIVE AVG=',F11.4,/,/, ' AVERAGE ENERGY OF ABSORBED PARTICLES=',F12.5,
      X3X, 'CUMULATIVE AVG=',F11.4,/)
0497 CKFNE=SUMNE*ANTPO
0498 WRITE(6,189)CKFNE, CESUM, TNION, TNIO2
0499 189 FORMAT(' ', 'THE SN # CHECK =',F12.4,3X, 'THE CN # CHECK=',F12.4,/,/,
      1' CALCULATED N+ USE F=',E14.6,3X, 'CALC. N+ USE CN =',E14.6,3X, 'AR
      XTIF CN=',E14.6,/)
0500 IF ((MOD(INT,51.NE.0).OR.(MOD(INT,10).EQ.0)) GO TO 3
0501 WRITE(9) EAVG, DAVG, TME, NTW, NT, NTP, NION, DDEG, NSW,SWF, E, U
0502 WRITE(9) Z, WT, ANIOS, SF3
0503 ENDFILE 9
0504 REWIND 9
0505 3 CONTINUE
0506 IF (MOD(INT,10).NE.0) GO TO 4
0507 WRITE(10) EAVG, DAVG, TME, NTW, NT, NTP, NION, DDEG, NSW,SWF, E, U
0508 WRITE(10) Z, WT, ANIOS, SF3

```

```

FORTRAN IV G LEVEL 18          MAIN          DATE = 72126          09/31/02

0509          ENDFILE 10
0510          REWIND 10
0511          * CONTINUE
0512          IF (MOD(INT,NCUM1,NE,0) GO TO 6666
0513          NSTP=0
0514          NTP=0
0515          OSAVS=0.0
0516          NAST=0
0517          NIZNT=0
0518          NSVLE=0.0
0519          CN5IZ=0.0
0520          CN55=0.0
0521          CNMET=0.0
0522          CHEXC=0.0
0523          CESAVG=0.0
0524          CEASG=0.0
0525          CSF2=0.0
0526          CLOSS=0.0
0527          CNIO2=0.0
0528          CSF1=0.0
0529          DO 54 M=1,JZ
0530          CNREZ(M)=0.0
0531          54 CEZ(M)=0.0
0532          DO 55 M=1,JE
0533          TNE(M)=0.
0534          55 CNE(M)=0.0
0535          CNES=0.0
0536          CNAS=0.0
0537          OSAVS=0.0
0538          CETER=0.0
0539          CETAS=0.0
0540          IF (MOD(INT,10),NE,0) GO TO 6666
0541          6666 CONTINUE
0542          NT= NT + 1
0543          DO 56 M=1,500
0544          DIZN(M)=0.0
0545          IES(M)=0
0546          IRR(M)=0
0547          56 IAS(M)=0
0548          DO 51 M=1,50
0549          WSE(N)=0.0
0550          51 ESE(N)=0.0
0551          I=1
          C. REINITIALIZE THE OUTPUT QUANTITIES
0552          DO 52 M=1, JE
0553          FE(M)=0.
0554          52 NE(M)= 0.0
0555          DO 53 M=1,JZ
0556          NREZ(M)= 0.0
0557          EZ(M)=0.0
0558          DO 53 N=1,JE
0559          EKZ(N,M)=0.0
0560          53 NEZ(N,M)=0.0

```

FORTRAN IV G LEVEL	18	MAIN	DATE = 72126	03731702
0561	NTZN=0			
0562	NDZN=0			
0563	NTN=0			
0564	NES=0			
0565	NSE=0			
0566	NCOL=0			
0567	NAS=0			
0568	NASH1=0			
0569	DTOT=0.0			
0570	NMET=0			
0571	NEXC=0			
0572	WNES=0.0			
0573	WNAS=0.0			
0574	WNEXC=0.0			
0575	WNIZN=0.0			
0576	WNMET=0.0			
0577	WNSE=0.0			
0578	ETAS=0.0			
0579	ETEP=0.0			
0580	ETOT=0.0			
0581	NW=0			
0582	NUP=0			
0583	NDN=0			
0584	NDNR=0			
0585	INIV=INIV+1100			
0586	CALL RN3INZ(INIV)			
0587	WRITE(6,72)			
0588	IF (NT.LE.NTC) GO TO 7777			
0589	9999 STOP			
0590	DEBUG SUBCHK			
0591	END			

```

FORTRAN IV G LEVEL 18      INDEXE      DATE = 72126      05/31/02
0001      FUNCTION INDEXE(E)
0002      COMMON /AREA4/ CONST,DEL2, ELT,IE, U11, O12, O13 ,DELI
0003      IF (E.LT.0.001) GO TO 11
0004      IF (E.GT.150.1333) GO TO 13
0005      12 INDEXE= 5*(ALOG10E)*CONST +4.01*O12
0006      RETURN
0007      11 INDEXE=1 + E*O11
0008      RETURN
0009      13 INDEXE=36 + (ALOG10E)*CONST-7.003*O13
0010      RETURN
0011      END

```

FORTRAN IV G LEVEL 18

SOURCE2

DATE = 7/12/6

05/31/72

```

0001      SUBROUTINE SOURCE2(NSD, EAVG, NTW, PM, FID)
0002      REAL N(55), FID(50), WI(55)
0003      COMMON D, E(5100), Z(5100), U(5100), IAS(500), IAN(500), NRR, IFS(600)
0004      A, WT(5100), SWF
0005      COMMON /AREA2/ C1, MUO, E1, F15, SF, ANEAM, PDPC, P1, PC, NION, JF, JZ, NASW1,
0006      IJSEW, AKT, AIKT, AM1, AMC1, ED1, PDMUO, PD26, K1, K2, K3, WN2, ED2, ED3
0007      COMMON /AS/ EG(55), EDE(55), EDD(55)
0008      F41(F) = EXP(-E*AIKT)*SQRT(E)*AM1
0009      F42(F) = AA/F
0010      F43(F) = EXP(-E*AIKT)*SQRT(E)*AMC1
0011      REAL MUO, NION
0012      404 FORMAT(' ', F12.6, 3X, I3, 3X, E12.4, 2X, F12.6, 2X, F12.5, 3X, E12.4)
0013      405 FORMAT(' ', 'J1=' , I3, 3X, 'ED1=' , F9.3, 3X, 'NSD=' , I3, 3X, 'K1=' ,
0014      I3, 3X, 'K2=' , I3, //, ' INITIAL DISTRIBUTION' , //, ' # OF PTS' , 7X, 'INDEX
0015      I' , 5X, ' DENSITY' , 4X, 'AVG ENERGY' , 6X, 'DEL-E' , 4X, 'WEIGHT' , /)
0016      606 FORMAT(' ', 'NS1=' , I5, 3X, 'NS2=' , I5, 3X, 'NSD=' , I5, 3X, 'NTW=' , I5, 3X,
0017      'X'NTW(CALC)' , ' , I5, //)
0018      607 FORMAT(' EAVG=' , F12.5, 3X, 'WGHTED EAVG=' , F12.5, //)
0019      AA = 2.0*PM*SQRT(PM/P1)*EXP(-PM)
0020      J1 = INDEXC(ED1-0.5)
0021      J2 = J1+1
0022      ISUM = 0
0023      WRITE(6, 606) J1, ED1, NSD, K1, K2
0024      DO 100 I = 1, J1
0025      DNI = FID(I)*EDE(I)*NTW/K1
0026      NI(I) = DNI
0027      WII(I) = 1.0
0028      IF (FID(I).LE.0.0) GO TO 112
0029      IF (DNI.GT.200.0) NI(I) = 200
0030      IF (DNI.GT.200.0) WII(I) = DNI/200.0
0031      IF (DNI.LT.12) NI(I) = 12
0032      IF (DNI.LT.12) WII(I) = DNI/12.0
0033      112 WRITE(6, 604) NI(I), I, FID(I), EG(I), EDE(I), WII(I)
0034      100 ISUM = ISUM + NI(I)
0035      JSUM = 0
0036      DO 200 I = J2, JF
0037      DNI = FID(I)*EDE(I)*NTW*K2
0038      NI(I) = DNI
0039      WII(I) = 1.0
0040      IF (FID(I).LE.0.0) GO TO 202
0041      IF (DNI.GT.200.0) NI(I) = 200
0042      IF (DNI.GT.200.0) WII(I) = DNI/200.0
0043      IF (DNI.LT.12) NI(I) = 12
0044      IF (DNI.LT.12) WII(I) = DNI/12.0
0045      202 WRITE(6, 604) NI(I), I, FID(I), EG(I), EDE(I), WII(I)
0046      200 JSUM = JSUM + NI(I)
0047      NSUM = ISUM + JSUM
0048      NTW2 = (ISUM*K1 + JSUM*WN2 + 0.5)
0049      WRITE(6, 606) ISUM, JSUM, NSUM, NTW, NTW2
0050      NTW = NTW2
0051      IE = 1
0052      EAG = 0.0
0053      DO 300 I = 1, JE

```

FORTRAN IV G LEVEL 1d

SOURCEZ

DATE = 72126

05/31/02

```

0049      IF (NI(1).LT.1) GO TO 300
0050      JI=NI(1) +IE -1
0051      DO 301 J=IE, JI, 2
0052      R1=RAN3Z(0)
0053      R2=1.0-R1
0054      E(J)= ERD(1) + EDE(1)*R1
0055      E(J+1)= FRD(1) + FDE(1)*R2
0056      WT(J)=W(1)
0057      WT(J+1)=WI(1)
0058      EAG =EAG + E(J) +E(J+1)
0059      IF((MOD((J)-IE),2).EQ.0).AND.(J.GE.JI)) EAG=EAG-E(J+1)
0060 301 CONTINUE
0061      IE=IE+NI(1)
0062 300 CONTINUE
0063      NSQ=NSUM
0064      EAVG=EAG/NSQ
0065      I2=NSQ
0066      ETW=0.0
0067      DO 400 I=1,NSQ
0068      IF (E(I).LE.ED1) WT(I)=K1 *WT(I)
0069      IF (E(I).GT.ED1) WT(I)=WN2*WT(I)
0070      IF (E(I).LE.0.0) E(I)=E(I-1)
0071      FTW=ETW + E(I)*WT(I)
0072 400 CONTINUE
0073      EWAVG=ETW/NTW2
0074      WRITE(6,607) EAVG, EWAVG
0075      QDZ=D*0.25
0076      I=1
0077 101 RZ=RAN3Z(0)
0078      RU=RAN3Z(0)
0079      Z(I)=D*RZ
0080      U(I)=2.0*RU -1.0
0081      I= I+ 1
0082 102 Z(I)=D*RZ
0083      U(I)=1.0- 2.0*RU
0084      I= I +1
0085 103 Z(I)=D*(1.0-RZ)
0086      U(I)= 2.0*RZ-1.0
0087      I= I +1
0088 104 Z(I)=D*(1.0-RZ)
0089      U(I)= 1.-2.*RZ
0090      I= I+ 1
0091      IF((I2 -I-3).GE.0) GO TO 101
0092      IF (I2.LT.1) GO TO 105
0093      RZ=RAN3Z(0)
0094      RU=RAN3Z(0)
0095      IF((I2 -I).EQ.2) GO TO 102
0096      IF((I2 -I).EQ.1) GO TO 103
0097      IF((I2 -I).EQ.0) GO TO 104
0098 105 RETURN
0099      END

```



```

FORTRAN IV G LEVEL 18          SOURCE          DATE = 72126          05/31/02

0001      SUBROUTINE SOURCE(NS,NTP)
0002      COMMON D, E(5100),Z(5100), U(5100) ,IAS(500),IRR(850),NRR,IES(600)
           X,WT(5100),SWF
0003      COMMON /AREA2/ C1,MUO,EI,EIS,SF,ANEAM,PDPO,P1,E0,NION,JE,J2,MASW1,
           INSEW ,AKT, AIKT ,AHI ,AMCI,ED1,PDHMO, PD26,K1, K2, K3, MN2,ED2,E03
0004      REAL MUO, NION
           C SOURCE PARTICLE PARAMETERS GENERATING PROCEDURE
           C I1: THE INITIAL INDEX VALUE FOR THE SOURCE PARTICLE
0005      I=NTP+1
0006      DO 100 K=1, NS, 2
0007      E(I)=E0
0008      RZ=РАН3Z(0)
0009      RU=РАН3Z(0)
0010      Z(I)=D*RZ
0011      U(I)=RU+RU-1.0
0012      WT(I)=MN2/SWF
0013      IF ((NS-K).LE.3) GO TO 100
0014      E(I+1)=E0
0015      Z(I+1)= RZ*D
0016      U(I+1)=-U(I)
0017      WT(I+1)=MN2/SWF
0018      I=I+2
0019      100 CONTINUE
0020      NTP=NTP+NS
0021      105 RETURN
0022      END

```

FORTRAN IV G LEVEL 18

MAIN

DATE = 72126

05/31/02

C WHENEVER THERE ARE MORE HOLES THAN SOURCE PARTICLES, FILL IN THE EXTRA HOLES W  
C WITH PARTICLES AT THE END

```

0001      SUBROUTINE FILHOL(NTP, NES, NAS)
0002      COMMON D, E(5100), Z(5100), U(5100), IAS(500), IRR(850), NRR, IES(600)
           X, WT(5100), SWF
0003      K=NTP
0004      IF (NES.LE.0) GO TO 105
0005      DO 100 I=1, NES
0006          J=IES(I)
0007          E(J)=E(K)
0008          Z(J)=Z(K)
0009          U(J)=U(K)
0010          WT(J)=WT(K)
0011      100 K=K-1
0012          NTP=NTP-NES
0013      105 IF (NAS.LE.0) RETURN
0014          K=NTP
0015          DO 300 I=1, NAS
0016              J=IAS(I)
0017              E(J)=E(K)
0018              Z(J)=Z(K)
0019              U(J)=U(K)
0020              WT(J)=WT(K)
0021      300 K=K-1
0022          NTP=NTP-NAS
0023      200 RETURN
0024      END

```

FORTRAN IV G LEVEL 18

RRPROC

DATE = 72126

05/31/02

```

0001      SUBROUTINE RRPROC(NTP,K3)
0002      COMMON D, E(5100),Z(5100), U(5100),IAS(500),IRR(850),NR,IES(600)
          X,WT(5100),SWF
0003      COMMON /A6/ ESE(50), NUP,NDNR,NDN, WSE(50)
0004      NR=NDNR
0005      K=NTP
0006      IF (K3.LE.1) RETURN
0007      100 CONTINUE
0008      I=IRR(NR)
0009      E(I)=E(K)
0010      U(I)=U(K)
0011      Z(I)=Z(K)
0012      WT(I)=WT(K)
0013      K=K-1
0014      200 NR=NR-1
0015      IF (NR.GT.0) GO TO 100
0016      NTP=K+1
0017      300 RETURN
0018      END

```

FORTRAN IV G LEVEL 18

ALDA

DATE = 72176

05/31/02

```

0001      FUNCTION ALDA(E)
0002      COMMON /AREA2/ C1,MUO,EI,EIS,SF,ANEAM,PDPO,P1,E0,NION,JE,JZ,NASW1,
INSEW ,AKT, AIKT ,AM1 ,AMC1,ED1,PMUO, PD26,K1, K2, K3, WM2,ED2,ED3
0003      COMMON /AREA3/ XSEL,XELC,XRCB,XMET,XEXC,XION,XTOL,KELC
0004      REAL MUO, NION
0005      WELC=1.0/KELC
0006      IF (E.LE.5.) XELC=23.0*PDPO
0007      IF((E.GT.5.).AND.(E.LE.19.8)) XELC=23.0*PDPO*SQRT(5./E)
0008      IF (E.GT.19.8) XELC=EXP(-0.02*E)*16.76*PDPO
0009      XELC=XELC*WELC
0010      IF (E.LE.ED3 ) GO TO 5001
0011      4001 AE1= 0.25 +(E - 19.8)**2
0012      XMET= 0.109*(E - 19.807*AE1 + 0.0311
0013      XMET=XMET*PDPO
0014      E19= E - 21.4
0015      AE2= 1850. + E19**2
0016      XEXC= 68.80*E19*PDPO/AE2
0017      IF (E.LT.21.4) XEXC=0.0
0018      E24= E - E1
0019      AE3= 4900. + E24**2
0020      XION= 172.*E24/AE3
0021      XION=XION*PDPO
0022      IF (E.LT.24.46) XION=0.0
0023      XTOL= XMET +XEXC +XION +XELC
0024      XRCB=0.0
0025      XSEL=0.0
0026      GO TO 6001
0027      5001 BE1=0.250+ (E-19.8)**2
0028      XSEL=0.109*(19.8 -E )/BE1
0029      XSEL=XSEL*PDPO*ANEAM
0030      XRCB=PMUO*NION/SQRT(E)
0031      XION=0.0
0032      XEXC=0.0
0033      XMET=0.0
0034      XTOL= XSEL +XELC +XRCB
0035      6001 ALDA=1.0/XTOL
0036      RETURN
0037      END

```

FORTRAN IV G LEVEL 10

COLPRO

DATE = 72126

09/31/02

```

0001      SUBROUTINE COLPRO(I, DCEN, NTP, RTP, NUST, *, *)
0002      COMMON D, E(5100), Z(5100), U(5100), IAS(500), IRR(850), HRR, IES(600)
0003      X, HT(5100), SWF
0004      COMMON /AREA1/ NIZN, NOZN, NAS, DIZN(500), NSE, NMET, NEXC, ETAS,
0005      XNMET, WNEXC, WNIZN, ELS, WNAS, WNSE
0006      COMMON /AREA2/ C1, MUO, E1, EIS, SF, ANEAM, PDPO, PI, EO, NION, JE, JZ, NASW1,
0007      INSEW, AKT, AIT, AMI, AMCI, EDI, PDHUU, PDZ6, KI, K2, K3, WN2, ED2, ED3
0008      COMMON /AREA3/ XSEL, XELC, XRCB, XMET, XEXC, XION, XTOL, KELC
0009      COMMON /A6/ ESE(50), NUP, HDNR, HON, WSE(50)
0010      DIMENSION PX(4)
0011      FM3(E)=EXP(-E*AKT)*SQRT(E)*AMCI
0012      REAL MUO, NION
0013      C DECIDE WHICH KIND OF COLLISION
0014      XTI=ALDA(E(I))
0015      IF (E(I).LE.ED3) GO TO 5000
0016      C ENERGY REGION 1 19.0EV TO 1 KEV
0017      4000 PMET=XMET*XTI
0018      PEXC=XEXC*XTI
0019      PION=XION*XTI
0020      PELC=XELC*XTI
0021      PX(1)=PEXC
0022      PX(2)=PEXC+PMET
0023      PX(3)=PX(2)+PION
0024      IP=1
0025      DO 45 IX=1,3
0026      45 IF(RIP.GT.PX(IX)) IP=IX+1
0027      GO TO (4015,4013,4017,4011), IP
0028      4011 R1=РАН32(0)
0029      U(I)=2.0*R1-1.0
0030      RETURN
0031      C INELASTIC COLLISION PROCESS EXCITE TO META SYABLE STATE
0032      C EXCITAYION TO METASTABLE LEVEL
0033      4013 NMET=NMET+WT(I)
0034      E(I)=E(I)-19.80
0035      IF (E(I).LE.0.0) RETURN1
0036      NMET=NMET+1
0037      R11=РАН32(0)
0038      U(I)=R11+R11-1.0
0039      RETURN
0040      C EXCITATION TO OTHER LEVELS
0041      4015 WNEXC=WNEXC+WT(I)
0042      E(I)=E(I)-21.40
0043      IF (E(I).LE.0.0) RETURN1
0044      NEXC=NEXC+1
0045      R11=РАН32(0)
0046      U(I)=R11+R11-1.0
0047      RETURN
0048      C IONIZATION PROCESS PRODUCE SECONDARY ELECTIONS
0049      4017 WNIZN=WNIZN+WT(I)
0050      E(I)=E(I)-E1
0051      IF (E(I).LE.0.0) RETURN1
0052      NIZN=NIZN+1
0053      NOZN=NOZN+1

```

FORTRAN IV G LEVEL 18

COLPRO

DATE = 72126

05/31/02

## C IONIZATION PRODUCED PARTICLES ALWAYS PUT TO THE END OF RESIDUES

```

0043      OIZN(NOZN)= OCEN
0044      RI=RAN3Z(0)
0047      U(I)=RI+RI-1.0
0048      NNTP =NTP +NIZN
0049      E(NNTP)= E(I)*0.1
0050      Z(NNTP)=Z(I)
0051      U(NNTP)=-U(I)
0052      E(I)= E(I)*0.9
0053      WT(NNTP)= WT(I)
0054      IF ((E(NNTP) .GT.ED1).OR.((E(I)*1.111+E(I).LT.FD1)) GO TO 4018
0055      NON=NON+1
0056      NONR=NONR+1
0057      WT(NNTP)=WT(I)*K3
0058      IRR(NONR)=NTP+NIZN
0059      4018 CONTINUE
0060      RETURN

```

## C ENERGY REGION 2 LOW ENERGY RANGE DOWN TO THERMAL REGION

```

0061      5000 PELC= XELC*XTI
0062      PSEL= XSEL*XTI
0063      PRCB= XRCB*XTI
0064      PX(1)= PRCB
0065      PX(2)= PSEL +PRCB
0066      IP=1
0067      DO 55 IX=1,2
0068      55 IF(RIP.GT.PX(IX)) IP=IX+1
0069      GO TO (5013,5015,5011),IP

```

## C ELASTIC COLLISION PROCESS

```

0070      5011 IF (ELS.LE.0.0) GO TO 50125
0071      IF ((INUST.EQ.0).OR.(E(I).GT.2.5)) GO TO 50124
0072      50115 RE=RAN3Z(0)
0073      R2=RAN3Z(0)
0074      E1=ED2*RE
0075      E2=ED2*R2
0076      FR1=FW3(E1)
0077      FR2=FW3(E2)
0078      IF (RE.LE.FR2) RS=RE
0079      IF (R2.LE.FR1) RS=R2
0080      IF ((RE.LE.FR2).OR.(R2.LE.FR1)) GO TO 5012
0081      GO TO 50115
0082      5012 ENN=RS*ED2
0083      50124 IF ((INUST.EQ.0).OR.(E(I).GT.2.5)) ENN=0.0
0084      R1=RAN3Z(0)
0085      U1=2.0*R1-1.0
0086      UM=SQRT((1.-U1**2)*(1.-U(I)**2))
0087      US=2.0*R1P-1.0
0088      COSST=U1*U(I) +SIGN(UM,US)
0089      DELE=C*(E(I)-ENN)*(1.0-COSST)*ELS*KELC
0090      E(I)= E(I) -DELE
0091      U(I)=U1
0092      50125 IF (ELS.LE.0.0) U(I)= 2.0*RAN3Z(0) -1.0
0093      RETURN

```

## C RECOMBINATION PROCESS

FORTRAN IV G LEVEL 18

COLPRO

DATE = 72126

05/31/02

```

0094      5013 NAS=NAS+1
0095          IAS(NAS)=2
0096          WNAS=WNAS+WT(I)
0097          IF (E(I).GT.0.0) ETAS=ETAS+E(I)*WT(I)
0098          RETURN2
C SUPERELASTIC COLLISION PROCESS
0099      5015 E(I)=E(I)+EIS
0100          WNSE=WNSE+WT(I)
0101          NSE=NSE+1
0102          RP= RAN32(0)
0103          U(I)= RP +RP -1.0
0104          RETURN
0105          END

```

FORTHAN IV G LEVEL 10

MAIN

DATE = 72125

01/28/29

```

C MASTER PROGRAM OF MONTE CARLO SIMULATION OF NONLINEAR ALPHA PLASMAS (ALPHA PIR
C TICLES INDUCED PLASMA ELECTRON ENERGY DISTRIBUTION STUDY
C DEFINITIONS AND TERMINOLOGY
C D: DIAMETER OF THE CYLINDER THE LENGTH OF THE FINITE REGION
C E: INITIAL ENERGY OF THE SOURCE PARTICLES SCINSI: SOURCE RATE
C JZ: INITIAL NUMBER OF ZONES JE: TOTAL NUMBER OF ENERGY RANGES
C IAS: ARRAY FOR STORING THE INDEX OF ABSORBED PARTICLES NASI: THE NO. OF THEM
C DRI: DISTANCE TO BOUNDARY DCOL: DISTANCE TO COLLISION DCEN: DISTANCE TO CENSUS
C DZ: ZONE INTERVAL
C ARRAYS E, Z, U STOPPING ENERGY, DISTANCE, AND DIRECTION OF PARTICLES
C NTP: TOTAL NUMBER OF PARTICLES TO BE PROCESSED EACH CYCLE
C NT: TOTAL NUMBER OF TIME CYCLES
C NS: NUMBER OF SOURCE PARTICLES PER UNIT TIME-DISTANCE OR SOURCE RATE
C NIZN: NUMBER OF IONIZATION PROCESS OCCURRED EACH TIME CYCLE
0001      REAL NF(55), NEZ( 55,10), FZ(11), DX(3), EE(55), CEZ(11), EKZ(55
          1,10), FIC(60), NRZ(11), EZ(11), NTPW
0002      DIMENSION FNF(55), FNREZ(11), NRZ(11), CNE(55), CNREZ(11), TNE(55)
0003      COMMON D, F(5100), Z(5100), U(5100), IAS(500), IRR(950), NRP, IFS(600)
          X, WT(5100), SWF
0004      COMMON /ARFA1/ NIZN, NOZN, NAS, DIZN(500), INSE, NMET, NEXC, ETAS ,
          XNMET, WNFXC, WNIZN, ELS, WNAS, WNSE
0005      COMMON /AREA2/ C1, MU0, EI, EIS, SF, ANEAM, POPC, PI, EC, NION, JE, JZ, NASW1,
          INSEW, AKT, AIKT, AM1, AMCI, ED1, PMUO, PD26, K1, K2, K3, WN2, ED2, ED3
0006      COMMON /AREA3/ XSEL, XELC, XRCR, XMET, XEXC, XTOW, XTOL, KELC
0007      COMMON /ARFA4/ CONST, DEL2, ELT, IE, D11, D12, D13, DEL1
0008      COMMON /A5/ EG(55), EDE(55), EBD(55)
0009      COMMON /A6/ ESE(50), NUP, NONR, NDR, WSE(50)
0010      REAL MUO, NION, ND
C STATEMENT FUNCTION FOR EVALUATING MAXWELLIAN AND RELATED FCNS
0011      FMI(E, FNEH) = EXP(-E*AIKT)*SQRT(E)*FNEH
0012      DATA TMS, NT, DTOT, ETOT, NIN, NIZNT, NAST/0.0, 1.0, 0.0, 0.0, 0.0, 0.0/
0013      DATA WSVLE, CNS1Z, CNS5, CNES, CNAS, CNMET, CNEKC, CESAVG, CEASG/9*0.7/
0014      DATA CNF/ 55*0.0/, CEZ/11*0.0/, CNREZ/11*0.0/, TNE/55*0.0/, EKZ/55*0
          1.0/, DSAVS/0.0/
0015      DATA NEZ/ 55*0.0/, NE/55*0.0/, EZ/11*0.0/, NREZ/11*0.0/, EE/55*0.0/
0016      READ(5,1) NTAPE
0017      READ(5,151) K1, K2, NTC, NTP, NTW, NPH, KELC, JZ, IM, NCUM, NDPN, NTMP
          X, MC, IUSN, MUST, NBF
0018      READ(5,152) ED1, ND, EQ, ELS, D, ANSW, POPO, SR
0019      READ(5,153) FID
0020      1 FORMAT( 110)
0021      151 FORMAT( 1615)
0022      152 FORMAT( 8F10.4)
0023      153 FORMAT( 8E10.3)
0024      AKT=0.05
0025      AIKT= 1.0/AKT
0026      PI=3.141593
0027      EMAS1=1.60203E+16/9.108
0028      K3=K1*K2
0029      WN2=1.0/K2
0030      CJD=0.0
0031      EI=24.46
0032      EIS=21.4

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0033      C1=2.0*0.00549/4.003873
0034      ETV= SORT(2.0*1.60203/9.108)*1.0E+08
0035      VTF= 1.0/(ETV*ETV)
0036      CALL UNDER2('OFF')
0037      PD26=PDPU*26.0
0038      MUJ=(1.0E+C9)/ETV
0039      CMUQ=1.0E-09
0040      CMUPDP=CMUJ*PDPU
0041      IF (PDPU.GT.10.0) CMUPDP=CMUQ*10.0
0042      PMUQ=CMUPDP/ETV
0043      CLOUS=0.0
0044      NUP=0
0045      NDN=0
0046      NDNF=0
0047      INC=0
0048      NKR=0
0049      NAS=0
0050      NES=0
0051      NSE=0
0052      NOZN=0
0053      NI7N=0
0054      NIN=0
0055      LK=0
0056      INIV=999999
0057      NMET=0
0058      NEXC=0
0059      ETAS=0.0
0060      ETEP=0.0
0061      NIP5=0
0062      NW=0
0063      N5TP=0
0064      DTAVS=0.0
0065      CSF2=0.0
0066      CNIO2=0.0
0067      CSF1=0.0
0068      CETAS=0.0
0069      CETEP=0.0
0070      NCOL=0
0071      WNMET=0.0
0072      WNSE=0.0
0073      WNEXC=0.0
0074      WNES=0.0
0075      WNAS=0.0
0076      WNIZN=0.0
0077      IE=4.8
0078      IE=JE
0079      DEN=3.56E+16*PDPU
0080      DETZ=D/JZ
0081      DZIV= 1.0/DETZ
0082      CONST=1.0/ALOG(2.0)
0083      ED2=0.447513
0084      ED3=19.8
0085      PM=ED2*AIKT

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0086      AM1=2.0*AIKT/SQRT(PI*AKT)
0087      CALL RNSINZ(994863211)
0088      SQ=PDPO*SR*1.0E+16
0089      AITG= CHUPDP/(AKT*PI)
0090      DE1=0.016
0091      DE2=0.355
0092      DE3=0.3
0093      D11=1.70.016
0094      D12=1./DE2
0095      D13=1./DE3
0096      AMC1= SORT(2.0*2.7183*AIKT)*ED2
0097      DO 11 I=6, 36
0098      ERL=(1-5)*DE2-4.0
0099      EMV= ERL -0.1685
0100      EBD(I)= 2.0**ERL
0101      EG(I-1)= 2.0**EMV
0102      11 EFD(I)=EBD(I)*(2.0**DE2-1.0)
0103      JE1=JE+1
0104      DO 12 I=37, JE1
0105      EAL= 7.005 + (I-36)*DE3
0106      EMV= ERL - 0.1405
0107      EBD(I)= 2.0**ERL
0108      EG(I-1)= 2.0**EMV
0109      12 EDE(I)= EBD(I)*(2.0**DE3-1.0)
0110      DO 10 I=1, 5
0111      EDE(I)= 0.016
0112      EG(I)= EDE(I)*(-0.008
0113      10 EBD(I)= EDE(I)*(I-1)
0114      EDE(36)=29.66
0115      EG(48)= 1737.482
0116      IF (NTAPE.NE.1) GO TO 2
0117      READ(1) EAVG, DAVG, TME, NTW, NT, NTP, NION, DDEG, NSW, SWF, E, U
0118      READ(10) Z, WT, ANIOS, SF3
0119      NS=NSW*K2
0120      REWIND 10
0121      NT=NT+1
0122      2 CONTINUE
0123      IF (NTAPE.NE.2) GO TO 5
0124      READ(9) EAVG, DAVG, TME, NTW, NT, NTP, NION, DDEG, NSW, SWF, E, U
0125      READ(9) Z, WT, ANIOS, SF3
0126      NS=NSW*K2
0127      REWIND 9
0128      NT=NT+1
0129      5 CONTINUE
0130      IF (NTAPE.NE.0) GO TO 122
0131      NION= SORT(SQ*10/AITG)
0132      ANEAM=NION*0.33333/DEN
0133      CALL SORCE2(NTP, EAVG, NTW, PH, FID)
0134      AL=ALDA(EAVG)
0135      DAVG=AL
0136      DDEG=DAVG*ND/SQRT(EAVG)
0137      SF2=1.0
0138      IF (NION.NE.0) SF2=NTW/NION

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0139      SN=SQ*AL*ND*SF2/(ETV*SQRT(EAVG))
0140      NSW=ANSW*0.5
0141      IF (SN.GT.ANSW) NSW=SN*0.5
0142      NS=NSW**K2
0143      SWF=NSW/SN
0144      SF=SWF
0145      ANIQ=NIEN
0146      CALL SOURCE(NS,NTP)
0147      SF1=SF2
0148      SF3=SF2
0149      TME=0.C
0150      122 CONTINUE
0151      SF=SWF
C GENERATE SOURCE PARTICLES TO INITIALIZE THE PROGRAM
0152      ANEAM=NIEN*0.33333/DEN
0153      WRITE(6,70)
0154      70 FORMAT('1','PROGRAM INITIALIZATION AND THE INITIAL PARAMETERS',//)
0155      WRITE(6,71) AITG, NIEN, NTP, AL,EAVG
0156      71 FORMAT('1','AITG=',F14.6,3X,'NIEN=',F14.6,3X,'NTP=',I12,3X,
0157      1'DCEN(EO)=',F14.6,3X,'EAVG=',F14.4,/)
0158      72 FORMAT('1',8X,'DCOL',13X,'DCEN',13X,'ENERGY',11X,'7',14X,'U',10X,
0159      1'WEIGHT',//)
0160      75 FORMAT('1',F14.6,3X,F14.6,3X,F14.5,3X,F12.4,3X,F12.4,3X,F13.4)
0161      WRITE(6,72)
0162      R1=RBAN32(0)
0163      1=1
C WITHIN THE OUTER I LOOP THE I TH PARTICLE HISTORY
C GEOMETRIC ROUTINE ENTERING HERE
0164      7777 IF ((E(1).LE.0.01.OR.(E(1).GT.EQ1)) GO TO 9008
0165      DCEN=DCEN*SQRT(E(1))
0166      7778 IF ((NIEN.GT.01.AND.(1.GT.NTP)) DCEN=012N(NQ2N+1)
0167      NPLG=0
0168      U(1)=U(1)*IUSN
0169      LF62=0
0170      ETAG1=1.0
0171      IF (E(1).GT.EQ1) ETAG1=-1.0
0172      3000 IF ((Z(1).LE.0.01.OR.(Z(1).GT.01.OR.(E(1).LE.0.01)) GO TO 9009
0173      J=INT(Z(1)*DZIV)+1
0174      R1=RBAN32(0)
0175      ALD=ALD*(E(1))
0176      DCOL=ALD*ABS(ALOG(R1))
0177      IF (NPLG.EQ.1) GO TO 809
0178      IF (MOD(J,IM).EQ.0) WRITE(6,75) (DCOL,DCEN,E(1),Z(1),U(1),WT(1))
0179      DCOL2=ABS(ALOG(1.-R1))*ALD
0180      DINT=DINT+(DCOL+DCOL2)
0181      ETOT=ETOT+E(1)
0182      IF (NDPN.LY.2) GO TO 809
0183      DCEN2=DCEN
0184      EIQ=E(1)
0185      U10=U(1)*IUSN
0186      Z10=Z(1)
0187      809 NPLG=1
0188      900 IF (DCEN.LT.DCOL) GO TO 902

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0187      901 DCEN=DCEN- DCOL
0188      Z(1)=Z(1)+DCOL*U(1)
0189      NCOL=NCOL+1
0190      IF (MOD(NCOL,MC).EQ.0) RIP=KANSZ(0)
0191      IF (MOD(NCOL,MC).NE.0) RIP=1.0-RIP
0192      CALL COLPRO(1, DCEN, NTP, RIP, NUST, 89008, 89009)
0193      GO TO 300
0194      902 Z(1)= Z(1)+DCEN*U(1)
0195      IF ((Z(1),LE.C.0).OR.(Z(1).GT.0)) GO TO 9008
0196      J=INT(Z(1)*DZIVI)+1
0197      IF (J.GT.JZ) J=JZ
0198      K= INDEX(E(1))
0199      IF (K.GT.JE) K=JE
0200      IF (K3.EQ.1) GO TO 904
0201      ETAG2=1.0
0202      IF (E(1).GT.ED1) ETAG2=-1.0
0203      IF ((ETAG1*ETAG2).GT.0.0) GO TO 904
0204      IF (ETAG2.GT.0.0) GO TO 903
0205      NUP=NUP+1
0206      FSE(NUP)=E(1)
0207      WT(1)=WT(1)/K3
0208      WSE(NUP)=WT(1)
0209      GO TO 904
0210      903 IF ((LFG2.EQ.1).OR.(NDPN.LT.2)) NDNR=NDNR+1
0211      IF ((LFG2.EQ.1).OR.(NDPN.LT.2)) IRR(NDNR)=1
0212      NON=NON+1
0213      IF (MOD(NON,K3).NE.0) GO TO 9065
0214      WT(1)=WT(1)*K3
0215      904 NEZ(K,J)= NEZ(K,J) + WT(1)
0216      FKZ(K,J)=EKZ(K,J)+E(1)*WT(1)
0217      9045 IF ((LFG2.EQ.1).OR.(NDPN.LT.2)) GO TO 9009
0218      LFG2=1
0219      E(1)=E10
0220      U(1)=U10
0221      Z(1)=Z10
0222      DCEN=DCFN2
0223      DCOL=DCOL2
0224      GO TO 900
0225      9008 NES=NES+1
0226      IF (E(1).LT.E01) WT(1)=WT(1)/K3
0227      WNES=WNES+WT(1)
0228      TES(NES)=1
0229      IF (E(1).GT.0.0) ETEP=ETEP+E(1)*WT(1)
0230      9009 I= I+1
0231      LFG2=0
0232      IF (I.LE.NTP) GO TO 7777
0233      IF ((I.GT.NTP).AND.(N1ZN.GT.0)) NOZN= NOZN-1
0234      IF (NOZN.LT.0) N1N=1
0235      IF ((N1ZN.GT.0).AND.(N1N.EQ.0)) GO TO 7778
0236      N1ZN=WN1ZN +0.5
0237      NWS=WNES +0.5
0238      NASW=WNAS +0.5
0239      WNAS=WNAS/KELC

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0240      WNSE=WNSE/KELC
0241      WNSE=WNSE+0.5
0242      DO 191 M=1,JE
0243      DO 191 N=1,JZ
0244      EE(M)=EE(M)+EKZ(M,N)
0245      191 NE(M)=NE(M)+NEZ(M,N)
0246      IF ((K3.EQ.1).OR.(NUP.EQ.0)) GO TO 24
0247      M3=K3/JZ
0248      IF (N3.LT.1) M3=1
0249      NK3=K3-1
0250      IF (NK3.LT.JZ) IZ=NK3
0251      IF (N3.GE.1) M3=N3
0252      DO 23 L=1,NUP
0253      K=INDX(ESE(N))
0254      EF(K)=EF(K)+(K3-1)*ESE(N)*WSE(N)
0255      NE(K)=NE(K)+(K3-1)*WSE(N)
0256      22 IF (NK3.GE.JZ) IZ=JZ
0257      DO 21 L=1,IZ
0258      NREZ(L)=NREZ(L)+M3*WSE(N)
0259      21 FZ(L)=FZ(L)+ESE(N)*M3*WSE(N)
0260      NK3=NK3-JZ*M3
0261      IF (NK3.LT.JZ) IZ=NK3
0262      M3=1
0263      IF (NK3.GT.0) GO TO 22
0264      23 CONTINUE
0265      24 CONTINUE
0266      DO 192 M=1,JJ
0267      DO 1915 N=1,JE
0268      NREZ(M)=NREZ(M)+NEZ(M,N)
0269      1915 FZ(M)=FZ(M)+FKZ(M,N)
0270      192 CONTINUE
0271      NTPN=NTP+NIZN-NFS-NAS
0272      NTPW=0.0
0273      FTOL=0.0
0274      DO 14 M=1,JF
0275      FTOL=FTOL+EF(M)
0276      14 NTPW=NTPW+NE(M)
0277      EWTOL=ETOL
0278      NTP=NTPN
0279      DAVG0=DAVG*ND
0280      DAVS=(DAVG0*1000.)*NTPW
0281      DSAVS=DSAVS+DAVS
0282      DLSG=1.2/DSAVS
0283      ANTP0=1.0/NTPW
0284      NSTP=NSTP+NTPW*0.5
0285      DTAVS=DTAVS+DAVS
0286      NTP5=NTP5+NTPN*DAVS
0287      NTP0=NTP
0288      EDAVG=ETOL*ANTP0
0289      NSQ=NS
0290      NSWQ=NSQ*WN2+0.5
0291      NTP=NTP+NIZN
0292      NTPP=NTP

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0293      DAVG=DTOT*0.5/NTP
0294      EAVG= EYDT/NTP
0295      DEG=ODEG/ETV
0296      TME=TME+DEG
0297      TMS=TMS+DEG
      C IF SOURCE PARTICLES SO(NS)< NES THEN ELIMINATE AS MANY RESIDUE PARTICLES AT TH
0298      NTPS=NTPW + NMS +NASW
0299      DCEN1= ALDA( EAVG)
0300      NTPW=NTPW +.5
0301      ODEG=DAVG*ND/SQRT( EAVG)
0302      SF2=1.0
0303      IF (ANIOS.NE.0) SF2=NTPS/ANIOS
0304      ESAVG=0.0
0305      IF(WNES.NE.0.0) ESAVG= EYEP/WNES
0306      EASAVG=0.0
0307      IF (WNAS.NE.0) EASAVG=ETAS/WNAS
0308      SWF=SWF
0309      RF=1.0
0310      IF ((WNES+WNAS).GT.0.0) RF=(WNIZN+SN)/(WNES+WNAS)
0311      CN5S=CN5S+ SN*DAVS
0312      CSF2= CSF2 +(ETEP+ETAS+40.*WNIZN)/TMS
0313      IF((MOD(NT,5).EQ.0).AND.(NT.NE.0)) SF1=CSF2/(50*EC*5.)
0314      CSF1= CSF1 + SF2*DAVS
0315      IF ((MOD(NT,5).EQ.0).AND.(NT.NE.0)) SF3=CSF1*D15G
0316      IF ((MOD(NT,5).EQ.0).AND.(NT.NE.0).AND.(SF3.LT.SF1*0.01)) SF3=
      XSQRT(SF1*SF3)
0317      WGHT=SF1/SF3
0318      SN=SN+ODEG*SF3*WGHT/ETV
0319      NSW=NSW+0.5
0320      IF (SN.GT.ANSW) NSW=SN +0.5
0321      SWF=NSW/5N
0322      NS=NSW*K2
0323      SF=SWF
0324      NMET=NMET +0.5
0325      NEXW=WNEXC +0.5
0326      W5VLE=W5VLE+ WVALUE*DAVS
0327      CETEP=CETEP + EYEP*DAVS
0328      CETAS=CETAS +ETAS*DAVS
0329      CN5IZ=CN5IZ+WNIZN*DAVS
0330      CNES=CNES+WNES*DAVS
0331      CNAS=CNAS+WNAS*DAVS
0332      CNMET=CNMET+NMET*DAVS
0333      CNEXC=CNEXC+WNEXC*DAVS
0334      CESAVG=CESAVG+ ESAVG*DAVS
0335      CEASG=CEASG+EASAVG*DAVS
0336      CLOSS= CLOSS+(WNIZN-WNES)*DAVS/(TMS*SF3*WGHT)
0337      WVLFI=W5VLE*D15G
0338      ALUSS=CLOSS*D15G
0339      ANSVG=CN5S*D15G
0340      ANIZVG=CN5IZ*D15G
0341      AETEP=CETEP/DTAVS
0342      AETAS=CETAS/DTAVS
0343      IF (ANIZVG.NE.0) WVLE2=(EO*ANSVG-AETEP-AETAS)/ANIZVG

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0344      ANES=CNES/DTAVS
0345      ANAS=CNAS/DTAVS
0346      AESAVG=CESAVG*DI5G
0347      AEASG=CEASG*DI5G
0348      ANMET=CNMET*DI5G
0349      ANEXC=CNEXC*DI5G
0350      ANS=ANES+ANAS-ANIZVG
0351      IF (ANS.LE.0.0) ANS=NSW/5NF
0352      IE (ANI7VG.NE.0.0) MVALUE=(SN*EO-AETEP-AETAS)/ANIZVG
0353      IF (WN17N.NE.0.0) WVC=(ANS*EO-ETEP-ETAS)/WN1Z
0354      IF (NS.GE.1) TMS=0.0
0355      NET=NES+NAS
0356      CALL SOURCE(NS,NTP)
0357      IF (NET.GT.0) CALL FILHOLINTP,NES,NAS
0358      90 NIZNT=NIZNT+NIZW
0359      IF (NUP.GT.0) CALL SEPROCINTP,K3
0360      IF (NDNR.GT.0) CALL RRPROCINTP,K3
0361      NAST=NAST+NWS
      C INTRODUCE THE SOURCE PARTICLES FOR THE NEXT CYCLE
      C EVERY FIFTH CYCLE RENORMALIZE NION AND OUTPUT QUANTITIES ARE PRINTED
      C RENORMALIZATION PROCESS OF NION
0362      ANSTP=1.0/NSTP
0363      ANTP5=DSAVS/NTP5
0364      NSTPN=NSTP
0365      DO 16 M=1,JZ
0366      CNREZ(M)=CNREZ(M)+NREZ(M)*DAVS
0367      16 CEZ(M)=CEZ(M)+EZ(M)*DAVS
0368      DO 18 M=1,JZ
0369      FNREZ(M)=NREZ(M)*ANTP5
0370      18 EZ(M)=FZ(M)/NREZ(M)
0371      SUMNE=0.0
0372      DO 193 M=1,JE
0373      SUMNE=SUMNE+NE(M)
0374      TNE(M)=TNE(M)+NE(M)/EDE(M)
0375      193 FNE(M)=NE(M)*ANTP5/EDE(M)
0376      EAV3=0.0
0377      DO 1755 M=1,JE
0378      EAV3=E3(M)+FNE(M)*EDE(M)+EAV3
0379      1755 CNE(M)=TNE(M)*ANSTP
0380      CNIO4=CMUPDP*2.0*AKT/SQRT(AKT*PI)
0381      ANION=0.0
0382      ANIO2=0.0
0383      DO 20 M=1,JE
0384      ANIO2=ANIO2+FM1(EG(M),CNE(M))*EDE(M)
0385      ANION=ANION+FM1(EG(M),FNE(M))*EDE(M)
0386      20 CONTINUE
0387      DNION=ANION*CNION
0388      DNIO2=ANIO2*CNION
0389      CJD=ALDSS
0390      IF (ALDSS.LT.0.0) CJD=0.0
0391      IF (ABS(CJD).GT.110.*50) CJD=110.*50
0392      TNION=ABS(ISO+CJD)/DNION
0393      TNIO2=ABS(ISO+CJD)/DNIO2

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0394      TNIO2=SQRT(TNIO2)
0395      TNION=SQRT(TNION)
0396      CNIO2=CNIO2 + TNIO2*DAVS
0397      ANIOS=CNIO2*0.156
0398      WF=1.0
0399      IF (ANES+ANAS.NE.0.0) WF=(ANIZVG+ANSVG)/(ANES+ANAS)
0400      IF (MOD(INT,NPM).NE.0) GO TO 1790
0401      NION=ANIOS
0402      IF ((WF.GT.0.005).AND.(WF.LT. 200.)) NION=ANIOS*WF**NRF
0403      1790 CONTINUE
0404      DAVT=DAVG*ND
0405      ALPHA=ANIZN/(NTPH*DAVT)
0406      ANEAM=NICN*0.33333/DEN
0407      1795 FORMAT('0','ANION=',E13.4,3X,'DNION=',E13.4,3X,'TNION=',E13.4,3X,
X'ALPHA=',F12.5,4X,'LEAK=',E13.4,/)
0408      WRITE(6,1785) ANION, DNIO2, TNIO2, ALPHA, ALOSS
0409      WRITE(6,171) EU, PDPO, D, DETZ, DEG, NDPN
0410      171 FORMAT('1','GAS: HE',2X,'SOURCE ENERGY:',F5.0,3X,'PRESSURE:',F4.0,
X'IDPR',3X,'SLAB THICKNESS:',F4.1, ' CM', ' DX:',F4.1, ' CM',4X,'DELT
X=',E12.5,3X,'#PROC',12,/)
0411      WRITE(6,180) NT, TME, NS,NSW
0412      180 FORMAT(' ', 'NUMBER OF TIME CYCLES=',I6,3X,'TIME =', F12.5, ' SEC',
13X, 'GENERATED SOURCE PART =', I3,3X, 'WGTED # OF SOURCE PART.=',
115)
0413      WRITE(6,1805) DAVT, DCEN, EAVG, EAVG, EAV3
0414      1805 FORMAT(' ', 'DAVG =',F12.5,3X, 'DCEN =',F12.5,3X, 'OVERALL EAVG=',
F10.5,3X, 'WGTED EAVG=',F10.5,3X, 'EAVG CHECK=',F10.5,/)
0415      WRITE(6,191)
0416      191 FORMAT(' ', 1X, 'NUMBER DISTRIBUTION',2X, 'FLUX DENSITY',4X, 'AVERAGE
1E ENERGY',2X, 'CUMULATED # DISTR',3X, 'ENERGY INTERVAL',2X, 'CUM FLUX
XDEN',4X, 'SDEN-DEVIATION',/)
0417      CESUM=0.0
0418      DO 281 M=1, JE
0419      CESUM=CESUM + CNE(M)*EDE(M)
0420      IF (NE(M).GT.0) DEAVG=EG(M)-EE(M)/NE(M)
0421      IF (NE(M).LE.0) DEAVG=0.0
0422      IF (NE(M).EQ.0) DNAVG=0.
0423      FNDV=CNE(M)-FNE(M)
0424      VFE=SQRT(EG(M))*ETV
0425      FLX1=FNE(M)*VFE
0426      FLX2=CNE(M)*VFE
0427      IF (NE(M).EQ.0) GO TO 281
0428      IF (((M.EQ.1).AND.(DEAVG.GT.0.0)).OR.((M.EQ.JE).AND.(DEAVG.LT.0.0))
X) DNAVG=-DEAVG*FNE(M)*0.5/EDE(M)
0429      IF ((DEAVG.LT.0.0).AND.((M+1).LE.JE)) DNAVG=-DEAVG*(FNE(M+1)-FNE(M
X))/(EG(M+1)-EG(M))
0430      IF ((DEAVG.GT.0.0).AND.(M.GT.1)) DNAVG=DEAVG*(FNE(M)-FNE(M-1))/(EG(
XM)-EG(M-1))
0431      281 WRITE(6,182) M,FNE(M),FLX1,EG(M),CNE(M),EDE(M),FLX2, DNAVG
0432      182 FORMAT(' ',I3,E12.5,5X,E12.4, 5X, F12.6,8X, E12.5,6X, F12.6,4X,E12
X.5,4X,E12.4)
0433      DET=DDEG/ETV
0434      WRITE(6,183) NTPP, NSTPN,NTPW,DET, DAVGO, NCDL

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0435      183 FORMAT(' ',//,3X,'TOTAL # OF PARTICLES PROCESSED =',I5,5X,'CUMULAT
        1ED # OF PARTICLES =',I10,3X,'WGTED TOTL # OF PARTICLES =',F12,3,/,/,
        X,'NEW DT =',E12,5,4X,'OLD DAVG =',F12,5,4X,'# OF COLLISIONS =',I5,/,/)
0436      WRITE(6,187)
0437      187 FORMAT(' ',//,1X,'ZONE NUMBER',5X,'AVERAGE ENERGY',5X,'NUMBER O
        1F ELECTRONS',3X,'CUMULATIVE # DEN(Z)',3X,'CUMU. AVG ENERGY',3X,
        1,'ZONE # DENSITY',/)
0438      ON 282 M=1, JZ
0439      ANZ=CNREZ(M)*ANSTP*DISG
0440      PCZ=CEZ(M)/CNREZ(M)
0441      282 WRITE(6,184) M, EZ(M), NREZ(M), ANZ, PCZ, FNREZ(M)
0442      184 FORMAT(' ',5X,15,5X,F14,6,8X,F14,6,5X,F14,5,7X,F14,6,7X,
        1E12,5/)
0443      WRITE(6,185) NION,WNIZN,NIZN, ANIZVG,WNES, NES,ANES, WNAS,NAS,
        XANAS,NTP, K1, K2, WF,BF
0444      185 FORMAT('O', 'NUMBER OF POSITIVE ION =',E14,6,/,/, 'NUMBER OF IONIZAT
        1IONS OCCURED =',F9,3,5X,'ACTUAL N =',I5,3X,'CUMULATIVE AVG =',F10,3,
        X,/,/, ' # OF ESCAPED PARTICLES =',F9,3,5X,'ACTUAL # =',I5,3X,'CUMULATIVE
        X AVG =',F9,2,/,/, ' # OF ABSORBED PARTIC =',E12,4,3X,'ACTUAL # =',I5,3X
        X,'CUMULATIVE AVG =',F9,2,/,/, ' TOTAL # OF PARTICLES IN THE NEW CYCL
        XE ',I5,/,/, ' K1 =',I2,3X,'K2 =',I2,3X,' BALANCE FACTOR WF =',
        XF10,5,3X,'BF2 =',F10,5,/)
0445      WRITE(6,186) ED1, NPM, ND, KELC
0446      186 FORMAT(' ENERGY DIVIDING LINE FOR RR PLAY E =',F10,4,4X,'EVERY',I2
        X,'TH CYCLE CHANGE N =',/,/, ' TIME INTERVAL MULTIPLY FACTOR =',F9,4,
        X,X, 'ELASTIC XSEC / BY',I2,/)
0447      NDR=NDN-NDN/K3
0448      WRITE(6,186) NSO, NSW, ANSVG,NDR,WNSE,WNHET,ANHET,WNEXC,ANEXC,NUP,
        XNGHT
0449      186 FORMAT(' ', ' SOURCE PARTICLES GENERATED OLD =',I10,/,/,
        1,' SOURCE PARTICLES (WEIGHTED) IN THE NEW CYCLE =',I10,3X,'CUMULAT
        1IVE AVG =',F10,3,/,/
        1,' # OF PARTICLES KILLED OFF IN RR PHNC =',I10,/,/,
        1,' # OF SUPER ELASTIC COLLISION =',F10,4,/,/,
        1,' # OF EXCITATIONS TO META STABLE LEVEL =',
        XF10,4,3X,'CUMULATIVE AVG =',F9,3,
        1,' # OF EXCITATIONS TO OTHER LEVELS =',F10,4,
        X,X, 'CUMULATIVE AVG =',F9,2,/,/, ' # OF PARTICLES CROSS OVER TO HIGH
        XENERGY REGION NUP =',I5,4X,'OVERALL WT =',E12,4,/)
0450      WRITE(6,195) EF, EDP, PDPO, SF1, SF2, SF3
0451      195 FORMAT(' E-FIELD =',F9,2,5X/CM',4X,'E/P =',F9,2,4X,'P/PO =',F7,1,
        X,X, 'SF1 =',E12,4,3X, 'SF2 =',E12,4,3X, 'SF3 =',E12,4,/)
0452      WRITE(6,170) SQ,WVC,SWED,WVALUE,WVLE1, WVLE2,ESAVG, AESAVG,EASAVG,
        XEASG
0453      170 FORMAT(' SOURCE RATE =',E11,3, ' PER SEC',4X,'W-VALUE(LOCAL ) =',
        XF10,3,3X,'SOURCE RED.FACT =',F13,3,/,/, ' W-VALUE =',
        XF10,3,4X,'CUM AVG WV1 =',F9,3,3X,'CUM AVG WV2 =',F9,3,
        1,' AVERAGE ENERGY OF ESCAPED PARTICLES =',F12,5,3X,'CUMULAT
        1IVE AVG =',F11,4,/,/, ' AVERAGE ENERGY OF ABSORBED PARTICLES =',F12,5,
        X,X, 'CUMULATIVE AVG =',F11,4,/)
0454      CKFNE=SUMNE*ANTPO
0455      WRITE(6,189) CKFNE, CESUM, THION, TMIQ2
0456      189 FORMAT(' ', 'THE SN # CHECK =',F12,4,3X,'THE CN # CHECK =',F12,4,/,/,

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MAIN

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1' CALCULATED N+ USE P='E14.6,3X,'CALC. N+ USE CN ='E14.6,3X, 'AR
X'IF CN='E14.6,771
0457 IF (MOD(NT,5).NE.0).OR.(MOD(NT,10).EQ.0) GO TO 3
0458 WRITE(9) EAVG, DAVG, TME, NTW, NT, NTP, NION, DDEG, NSW,SWF, E, U
0459 WRITE(9) Z, WT ,ANIOS, SF3
0460 ENDFILE 9
0461 REWIND 9
0462 3 CONTINUE
0463 IF (MOD(NT,10).NE.0) GO TO 4
0464 WRITE(10) EAVG, DAVG, TME, NTW, NT, NTP, NION, DDEG, NSW,SWF, E, U
0465 WRITE(10) Z, WT ,ANIOS, SF3
0466 ENDFILE 10
0467 REWIND 10
0468 4 CONTINUE
0469 IF (MOD(NT,NCUP).NE.0) GO TO 6666
0470 NSTP=0
0471 NTP5=0
0472 DSAVS=C,C
0473 NAST=0
0474 NIZNT=0
0475 WALE=C,C
0476 USS=0,0
0477 CSF1=0,C
0478 CNIO2=0,0
0479 CN512=0,0
0480 CN55=C,C
0481 CNMET=0,0
0482 CNEXC=C,C
0483 CESAVG=0,0
0484 CEASG=0,0
0485 DO 54 M=1,JZ
0486 CNREZ(M)=0,0
0487 54 CEZ(M)=C,C
0488 DO 55 M=1,JE
0489 TNE(M)=0,0
0490 55 CNE(M)=C,C
0491 IF (MOD(NT,10).NE.0) GO TO 6666
0492 LTAVS=0,0
0493 CETEP=0,0
0494 CETAS=0,0
0495 CNES=0,0
0496 CNAS=0,0
0497 6666 CONTINUE
0498 NT= NT + 1
0499 DO 56 M=1,500
0500 DIZN(M)=0,0
0501 IRN(M)=0
0502 TES(M)=0
0503 56 IAS(M)=0
0504 DO 51 N=1,50
0505 WSE(N)=C,C
0506 51 ESE(N)=0,0
0507 I=1

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C REINITIALIZE THE OUTPUT QUANTITIES

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0508      DO 52 M=1, J2
0509          RE(M)=0.
0510      52 NE(M)= 0.0
0511      DO 53 N=1, JZ
0512          NREZ(N)= 0.0
0513      53 EZ(M)=0.0
0514      DO 53 N=1, JE
0515          EKZ(N,M)=0.0
0516      53 NEZ(N,M)=0.0
0517          NIZN=0
0518          NOZN=0
0519          NIN=0
0520          NES=0
0521          NSE=0
0522          MCUL=0
0523          NAS=0
0524          DTOT=0.0
0525          NMET=0
0526          NEXC=0
0527          WNAS=0.0
0528          WNSF=0.0
0529          WNES=0.0
0530          WNIZN=0.0
0531          WNMET=0.0
0532          WNEXC=0.0
0533          ETAS=0.0
0534          FTEP=0.0
0535          ETUT=0.0
0536          NUP=0
0537          NONP=0
0538          NONR=0
0539          INIV=INIV+ 1100
0540          CALL RNINZ(INIV)
0541          WRITE(6,72)
0542          IF (NT.LT.NTC) GO TO 7777
0543      9999 STOP
0544          DEBUG SUBCHK
0545          END

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## VITA

Benjamin Shaw-hu Wang was born in [REDACTED]

[REDACTED] He is the son of Mr. C. H. Wang, the former Chinese Army General now the Congressman of the Republic of China. He received the B.S. degree in Electrical Engineering from the National Taiwan University of Taipei, Taiwan, China in 1960, he then served as an ensign in Chinese Naval Academy for one and a half years. He came to the United States in September, 1962 and enrolled in the University of Missouri at Rolla, Missouri. After he received the M.S. degree in Electrical Engineering in 1964, he joined the Westinghouse Research Laboratories and served as a research engineer for three and a half years.

In September 1967, he came to the University of Illinois at Urbana-Champaign to undertake studies for a Ph.D. in Computer Science Department. He has been a research assistant and teaching assistant in the Department of Computer Science and a research assistant in the Nuclear Engineering Program. He is the member of Kappa Mu Epsilon and Pi Mu Epsilon the two national honor societies of mathematics.